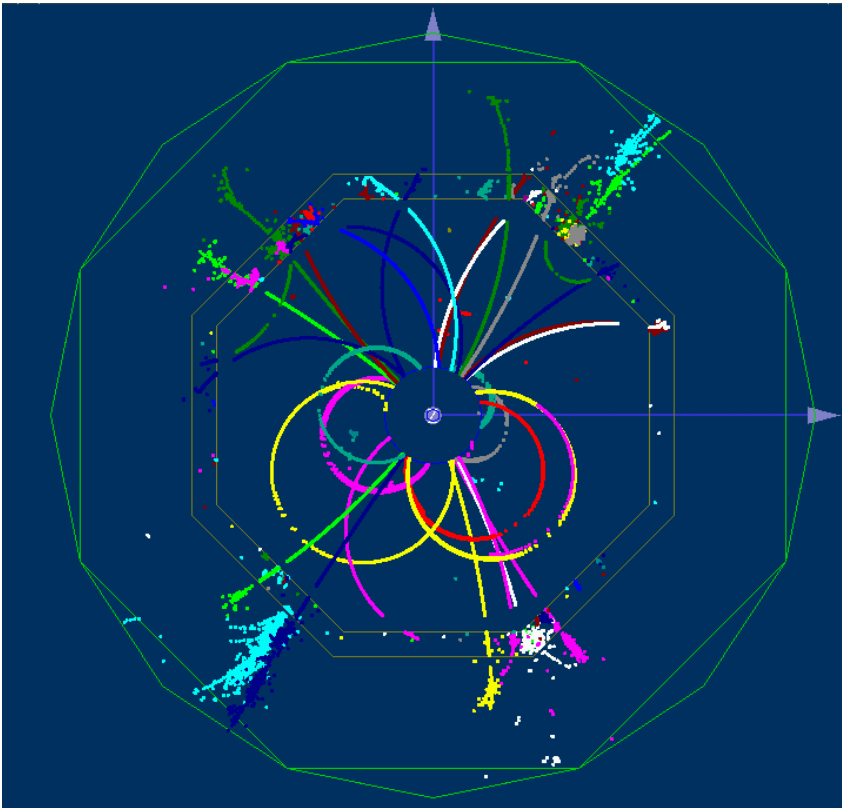


Particle Flow

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This Talk:

- ★ Software needs for Detector Optimisation
- ★ Particle Flow Algorithms
- ★ Current Results
- ★ Conclusions/Outlook

1 Detector Optimisation

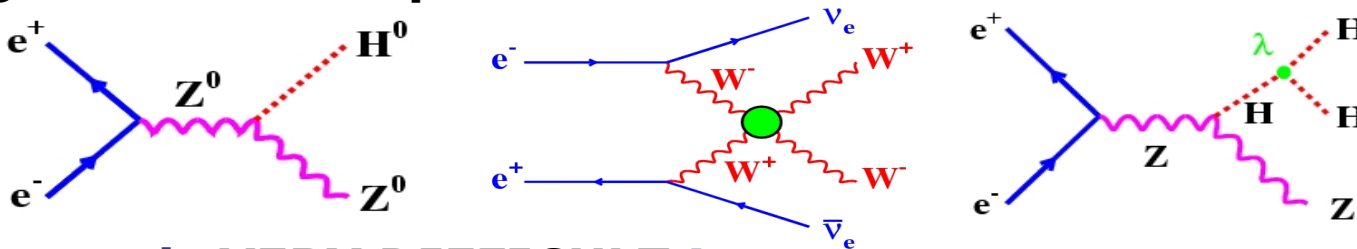


General consensus that **Calorimetry and PFA** drives overall **ILC** detector design

BUT: Don't really know what makes a good detector from point of view of PFA (plenty of personal biases – but little hard evidence)

How to optimise compare ILC detector design(s)

- ★ **Optimize** detector design using **key** physics processes
- ★ Need to choose the key “benchmark” processes (**DONE**)
e.g. the usual suspects +



★ The rest is **VERY DIFFICULT !**

★ Need unbiased comparison

- Same/very similar **reconstruction algorithms**
- **these need to realistic (i.e. start-of-art)**
- **Need Multiple PFAs** : avoid trap of optimising detector to flaws of particular algorithm
- This is a lot of work – need user friendly software

Detector Optimisation : Software Tools

- ★ Until **very** recently we did not have the software tools to optimise the detector from the point of view of Particle Flow
- ★ **This has changed !**
- ★ The basic tools are mostly there:
 - ★ **Mokka** : now has scalable geometry for the LDC detector
 - ★ **MARLIN**: provides a nice (and simple) **reconstruction framework**
 - ★ **LCIO**: provides a common format for worldwide PFA studies
 - ★ **Reconstruction**: in MARLIN framework already have **ALGORITHMS**

What is needed in MARLIN:

- ✦ **Digitisation: (take simulated hits → hits)**
 - ✓ simple MARLIN processors exist (more work needed)
- ✦ **Tracking: (two options currently in MARLIN)**
 - ✓ Full LEP like fit: TPC hits + currently being extended to VTX..
 - ✓ "Cheated" tracks: TPC/FTD/VTX use MC to assign hits to track. Track parameters from a Helix fit
- ✦ **Clustering: (two options)**
 - ✓ TrackWiseClustering (Alexei R. et al)
 - ✓ MAGIC (Chris Ainsley)
- ✦ **PFA: now (nearly) have two algorithms !**
 - ✓ Wolf (Alexei R.)
 - ✓ PandoraPFA (Mark Thomson) – will be released in January

★ All the necessary tools exist !

- that doesn't mean that its time to stop work...
- things aren't perfect **yet**



We are now in the position to start to learn how to optimise the detector for PFA

But first...

learning from ongoing studies of Perfect Particle Flow (P. Krstonosic)

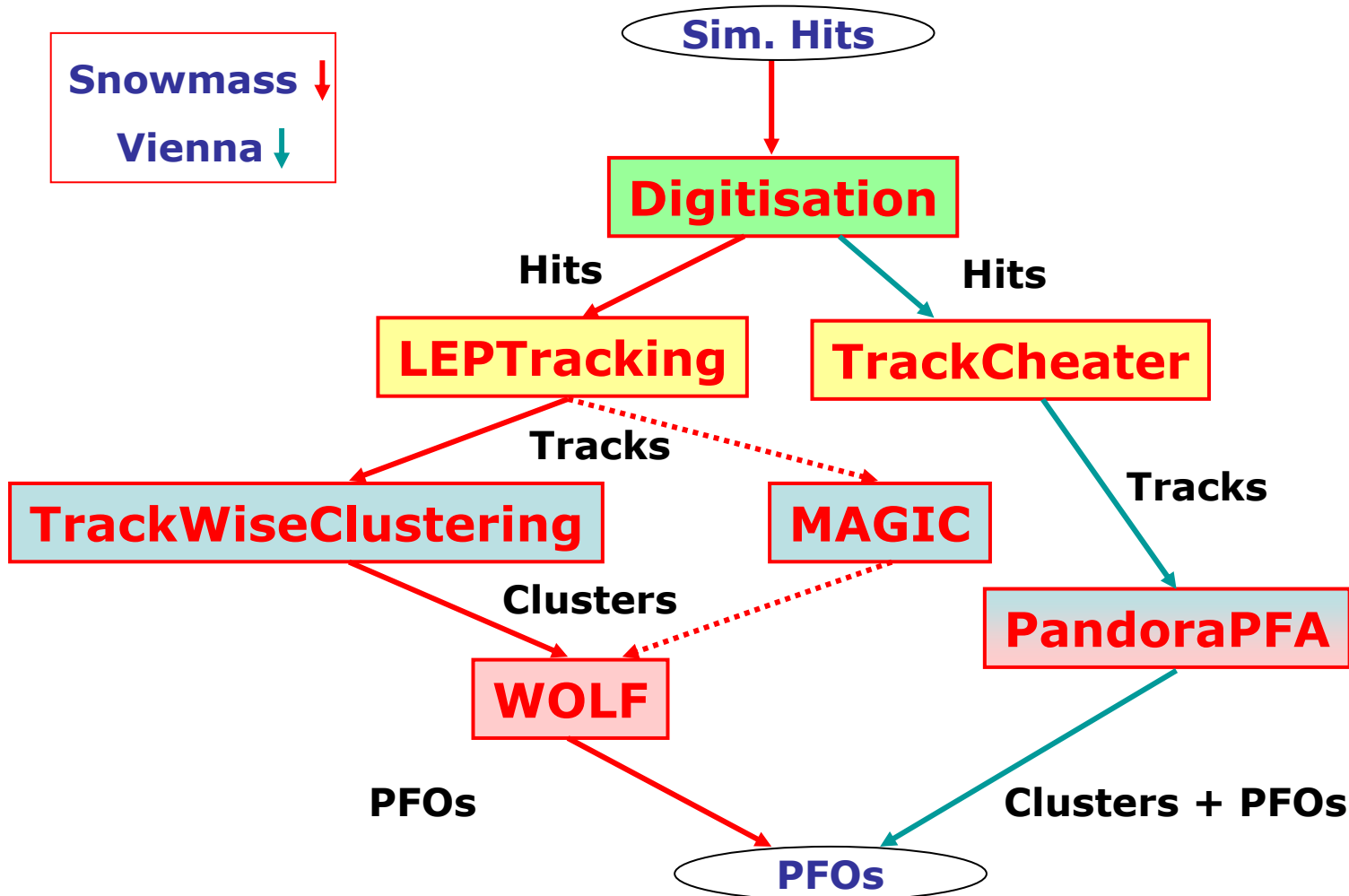
e.g. $e^+e^- \rightarrow Z \rightarrow qq$ at 91.2 GeV

To be reviewed

Effect	σ [GeV] separate	σ [GeV] not joined	σ [GeV] total (% / \sqrt{E})	σ % to total
$E_v > 0$	0.84	0.84	0.84 (8.80%)	12.28
$Cone < 5^\circ$	0.73	1.11	1.11(11.65%)	9.28
$P_t < 0.36$	1.36	1.76	1.76(18.40%)	32.20
σ_{HCAL}	1.40	1.40	2.25(23.53%)	34.12
σ_{ECAL}	0.57	1.51	2.32(24.27%)	5.66
$M_{neutral}$	0.53	1.60	2.38(24.90%)	4.89
$M_{charged}$	0.30	1.63	2.40(25.10%)	1.57

(assumed sub-detector resolutions: **ECAL 11%/√E**, **HCAL 50%/√E +4%**)

2 Particle Flow Algorithms in MARLIN

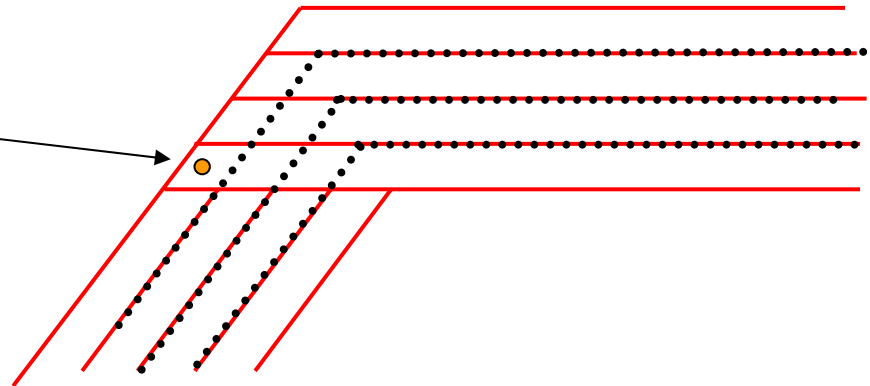


- ★ PandoraPFA/WOLF/MAGIC share many common features
- ★ Will briefly discuss some of the main points of the new Algorithm

PandoraPFA Clustering I

- ★ All **current MARLIN** clustering algorithms are “forward projecting”
 - Form clusters starting from inner CAL layer – working outwards
- ★ Arrange hits into **PSEUDOLAYERS** (same done in **MAGIC**)
 - i.e. order hits in increasing depth within calorimeter
 - PseudoLayers follow detector geometry

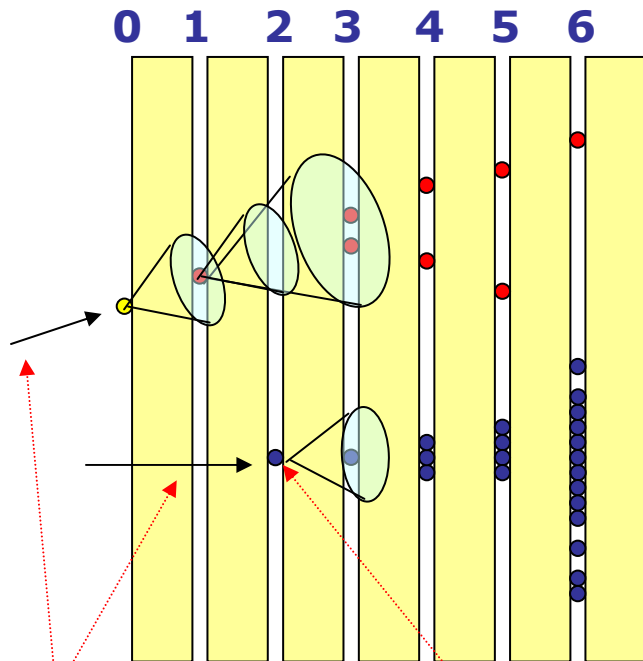
- Hit in early layer
- But high PseudoLayer



(WOLF orders hits by distance from IP)

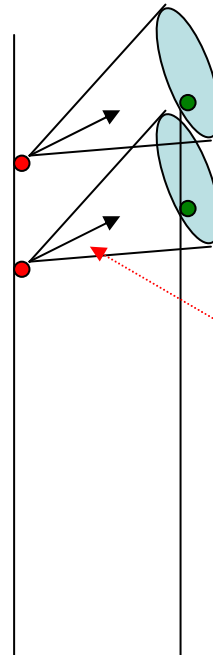
PandoraPFA Clustering II

- ★ Start at inner layers and work outward
- ★ Associate Hits with existing Clusters
- ★ If multiple clusters "want" hit then **Arbitrate**
- ★ Step back **N** layers until associated
- ★ Then try to associate with hits in current layer (M pixel cut)
- ★ If no association made form new Cluster
- ★ + tracks used to seed clusters



Initial cluster direction

Unmatched hits seeds new cluster



Simple cone algorithm based on current direction + additional N pixels

Cones based on either: initial PC direction or current PC direction

WOLF/MAGIC do things slightly differently but same basic idea

PandoraPFA Cluster Association

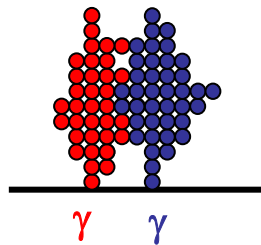
- ✦ By design clustering errs on side of caution
i.e. clusters tend to be split
- ✦ **Philosophy:** easier to put things together than split them up
- ✦ Clusters are then associated together in two stages:
 - 1) Tight cluster association - clear topologies
 - 2) Loose cluster association - catches what's been missed but rather crude



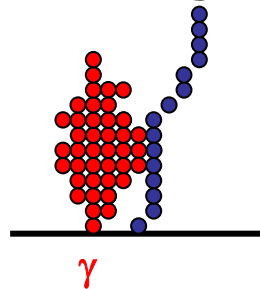
Photon ID

- ★ Photon ID plays important role
- ★ Simple "cut-based" photon ID applied to all clusters
- ★ Clusters tagged as photons are immune from association procedure - just left alone

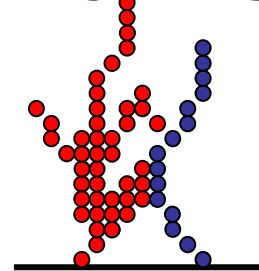
Won't merge



Won't merge

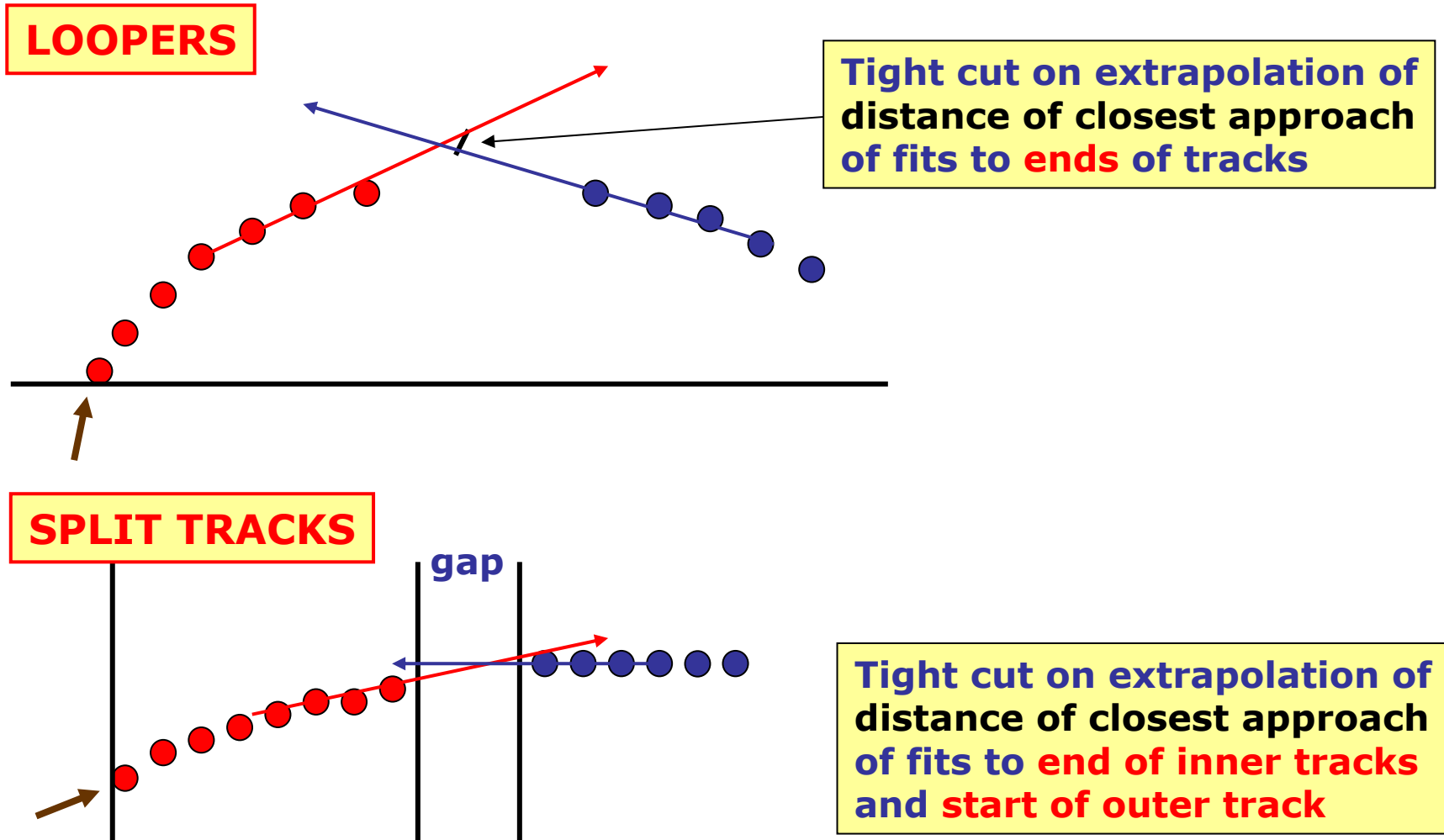


Could get merged



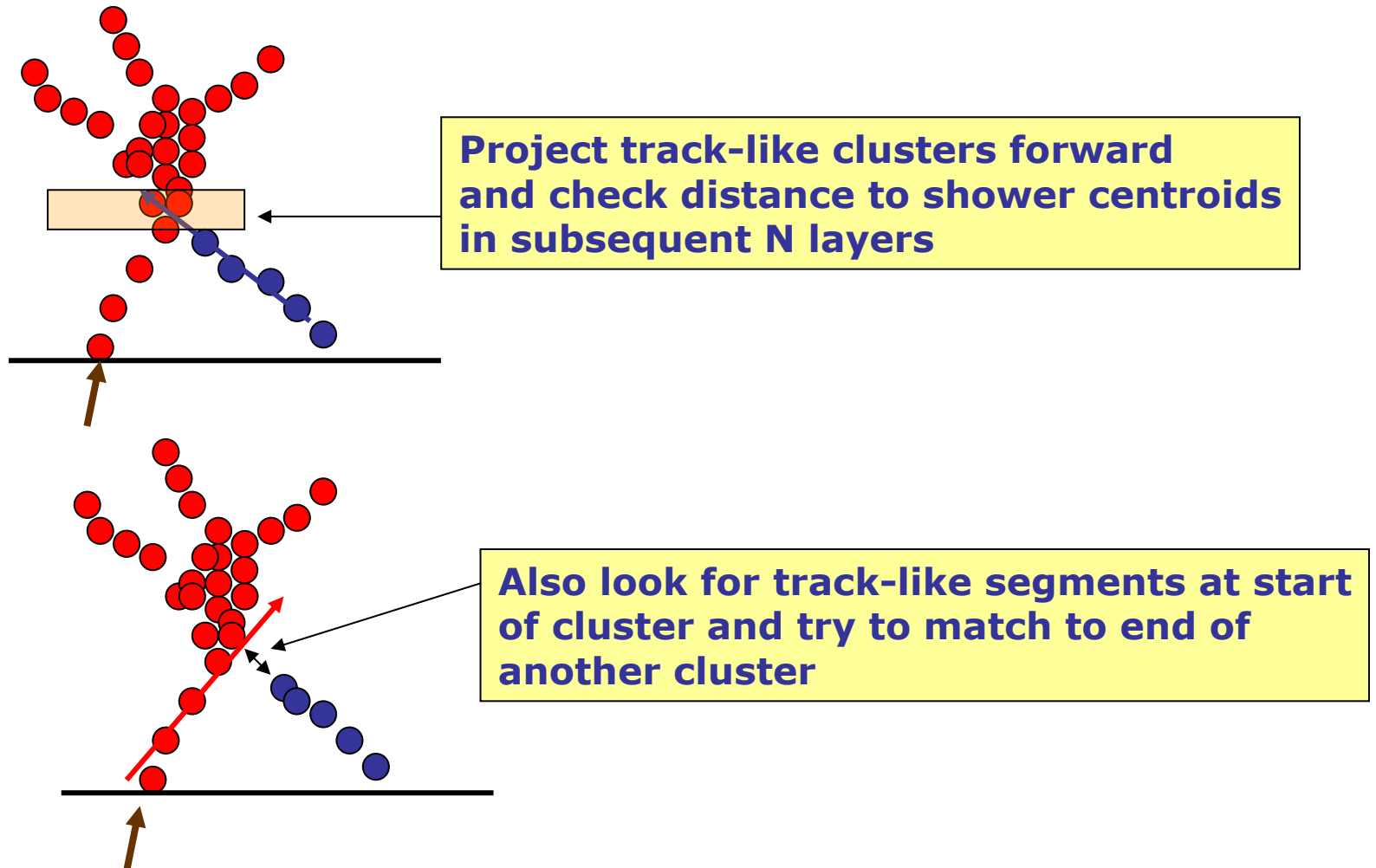
WOLF/MAGIC do things differently but both perform cluster merging

Cluster Association I : track merging



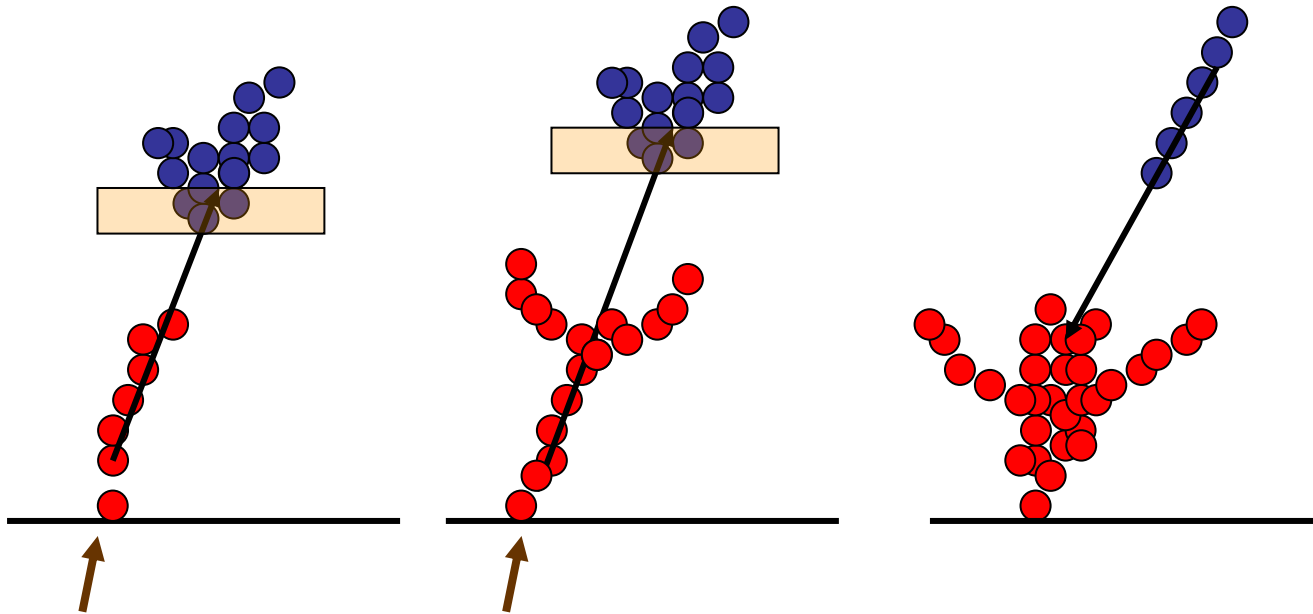
Cluster Association II : Backscatters

- ★ Forward propagation clustering algorithm has a major drawback: back scattered particles form separate clusters



Cluster association III : MIP segments

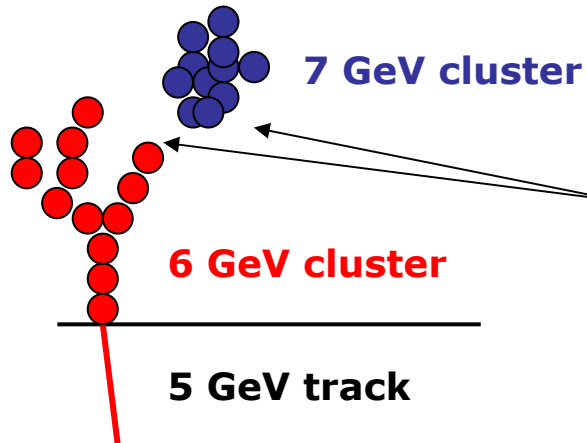
- ★ Look at clusters which are consistent with having tracks segments and project backwards/forward



- ★ Apply tight matching criteria on basis of projected track
[NB: + track quality i.e. χ^2]

Cluster Association Part II

- Have made very clear cluster associations
- Now try “cruder” association strategies
- **BUT first associate tracks to clusters (temporary association)**
- Use track/cluster energies to “veto” associations, e.g.

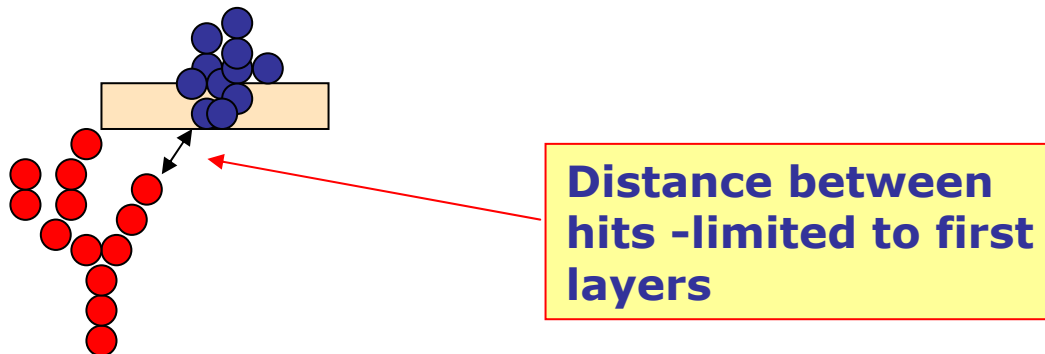


Provides some protection against “silly” mistakes

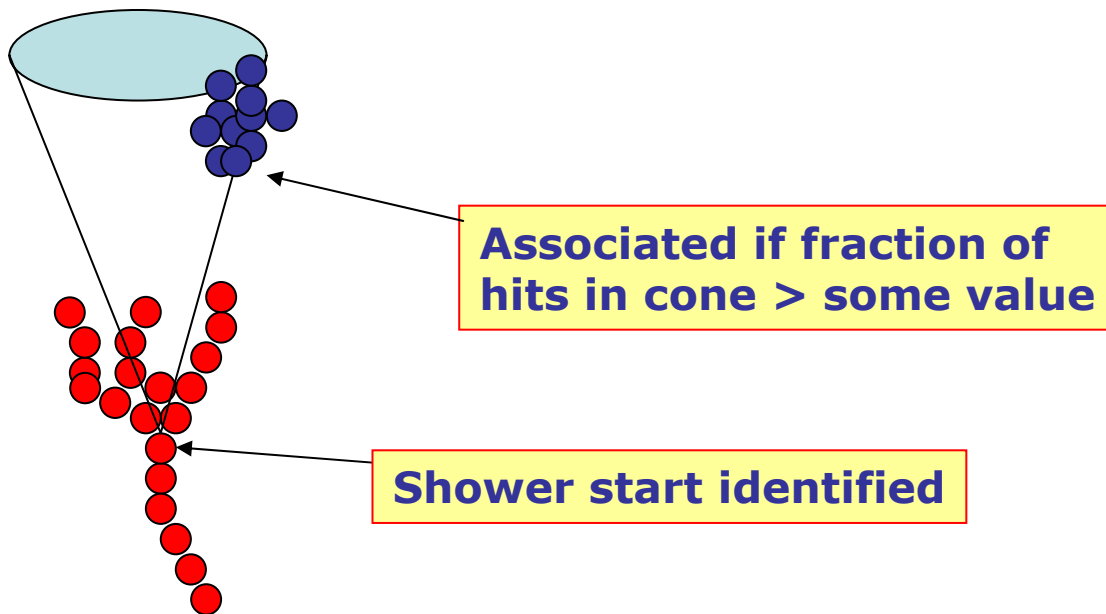
★ Clustering and PFA not independent

Sledgehammer Cluster Association

Proximity



Shower Cone



+Track-Driven Shower Cone

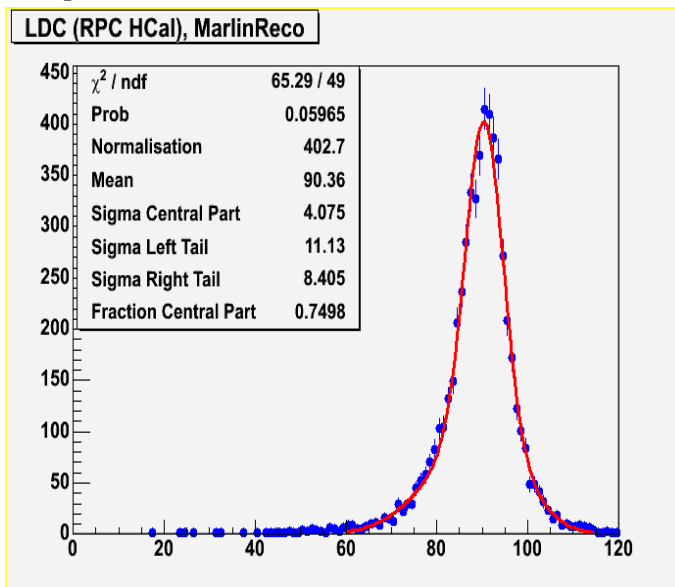
Apply looser cuts if have low E cluster associated to high E track

3 PFA Results

- ★ Currently PFA performance only investigated for $Z \rightarrow qq$ at 91.2 GeV
 - Good place to start as relatively simple (spread out jets)
- ★ Need to define figure of merit

Snowmass-style

Fit central Gaussian + asymmetric tails

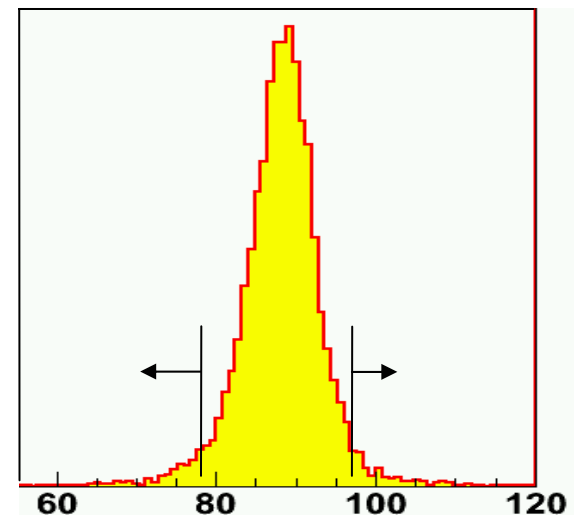


OK, but requires care:

- \Rightarrow sigma + fraction in peak
- + results will depend on fit region

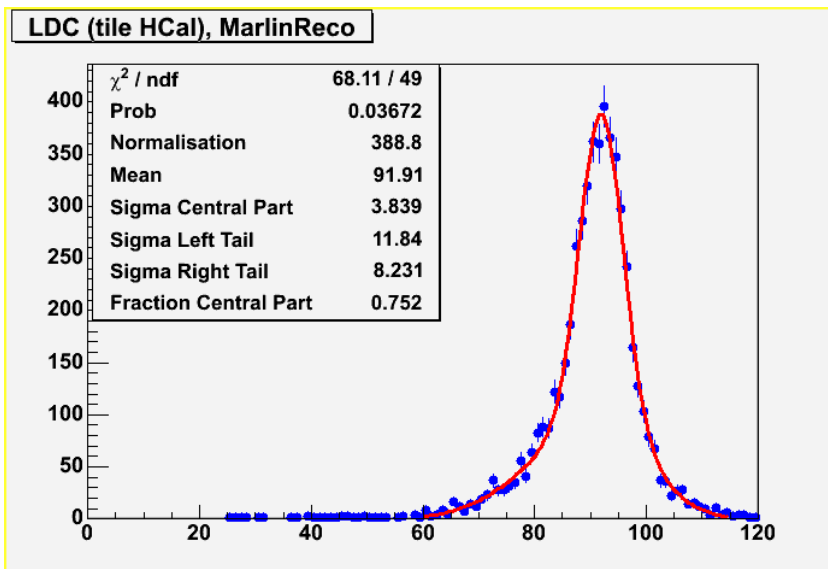
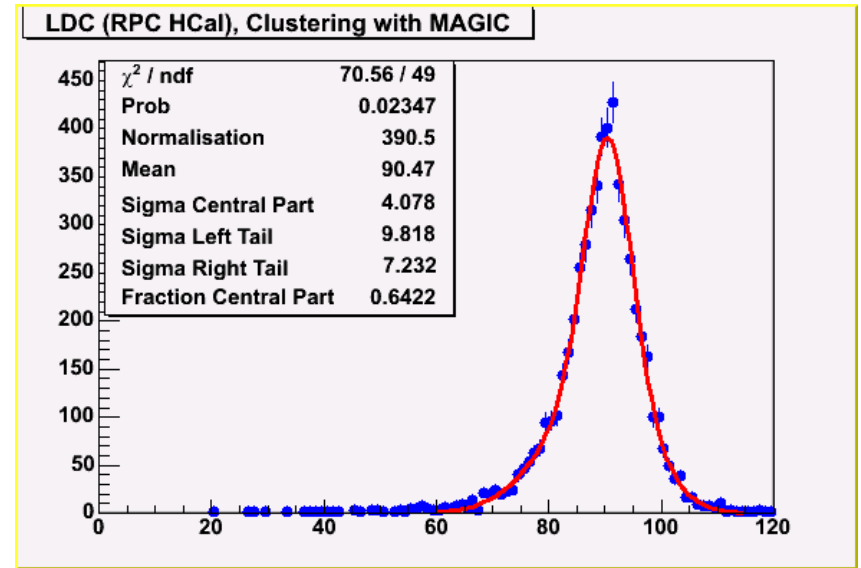
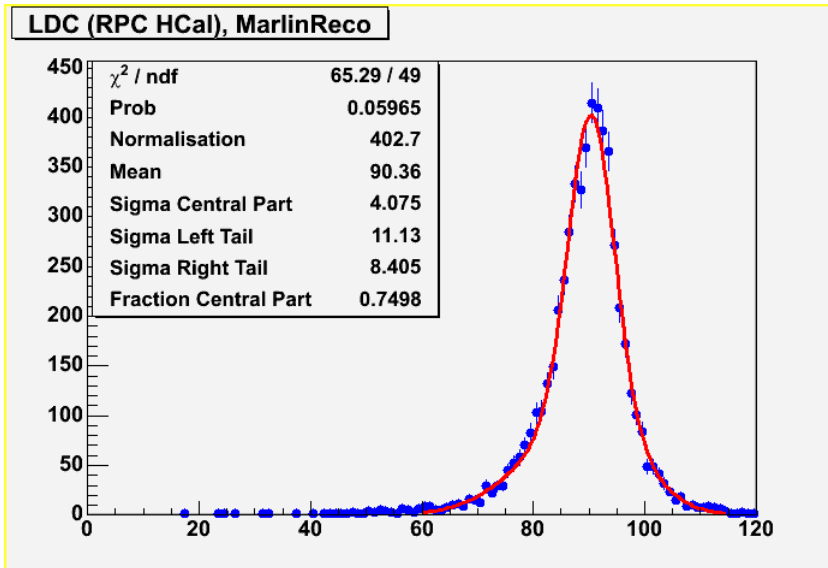
New Proposed Figure of Merit:

- ★ Find smallest region containing 90 % of events
- ★ Determine rms in this region



★ More robust

Wolf Results ($Z \rightarrow uds$)



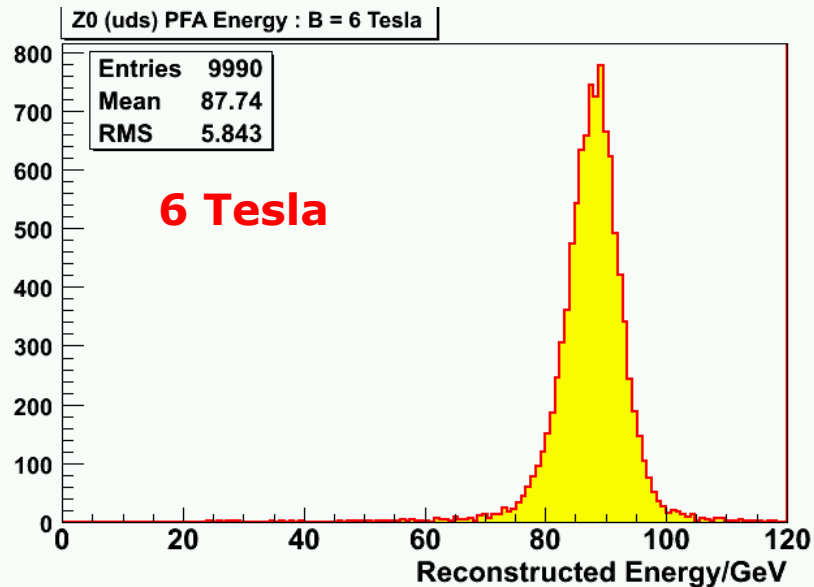
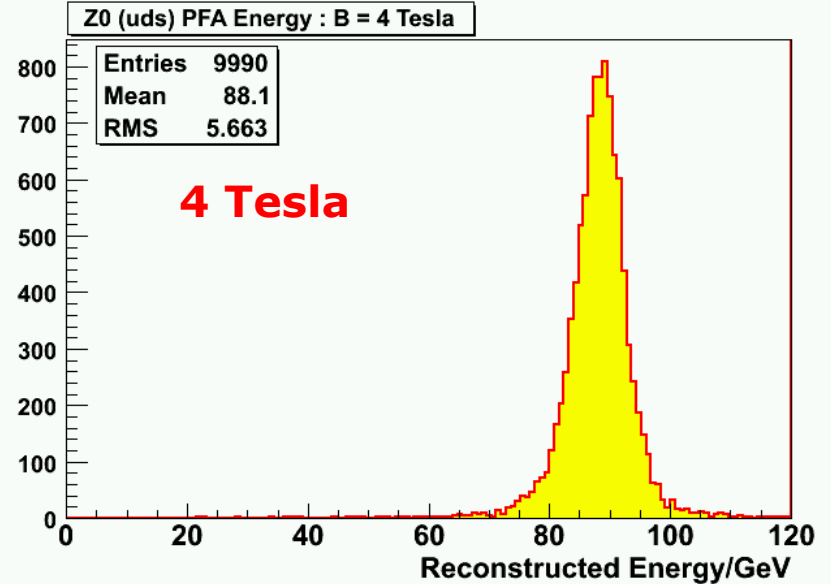
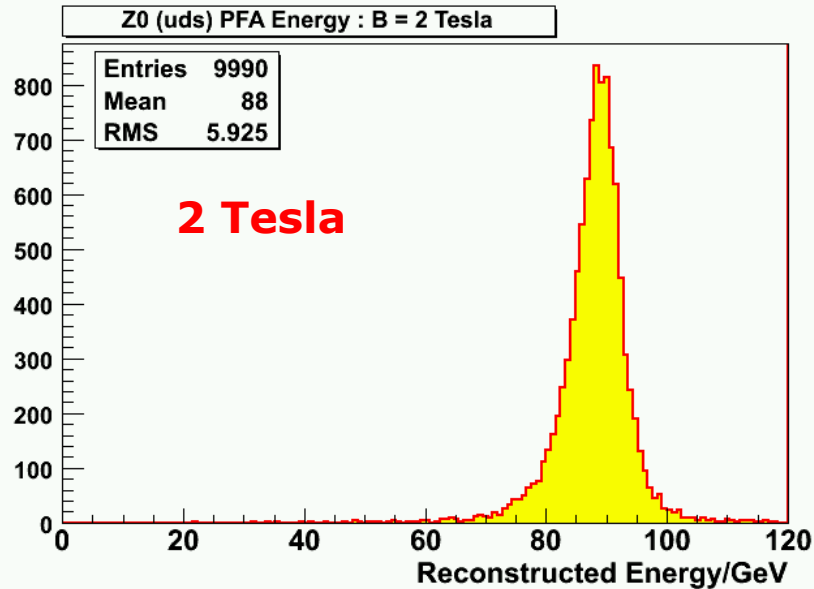
★ RMS of Central 90 % of Events

	RMS (90%)
RPC HCAL	4.3 GeV
Tile HCAL	4.1 GeV
RPC (MAGIC)	4.4 GeV

- RMS (90 %) is somewhat larger than width of fitted peak

(Results for Reco Tracks)

PandoraPFA Results ($Z \rightarrow uds$)



* RMS of Central 90 % of Events

B-Field	$\sigma_E/E = \alpha\sqrt{(E/\text{GeV})}$
2 Tesla	$35.3 \pm 0.3\%$
4 Tesla	$35.8 \pm 0.3\%$
6 Tesla	$37.0 \pm 0.3\%$

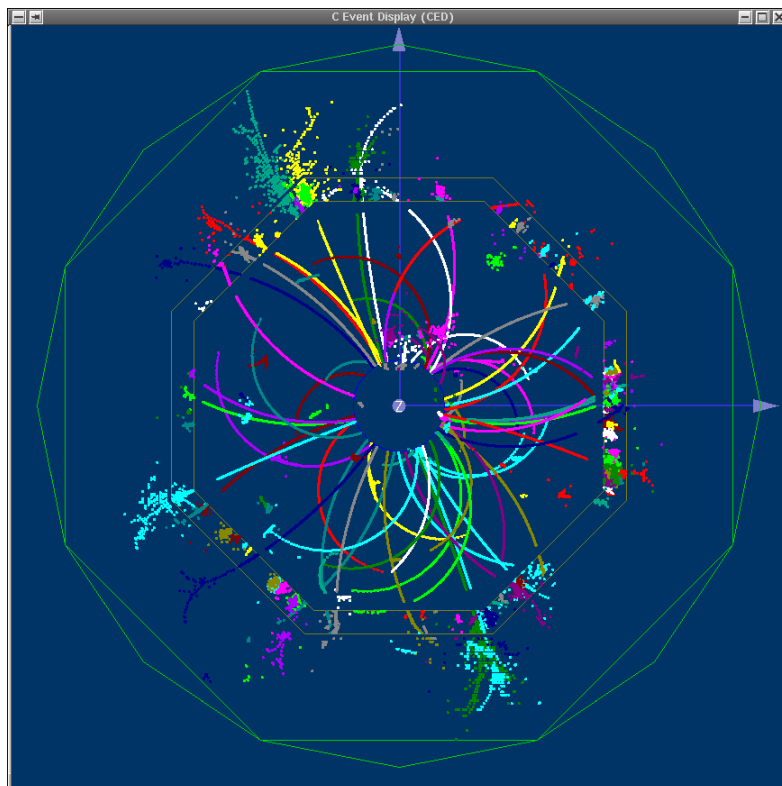
✦ only weakly depends on B

(Results for Cheated Tracks)

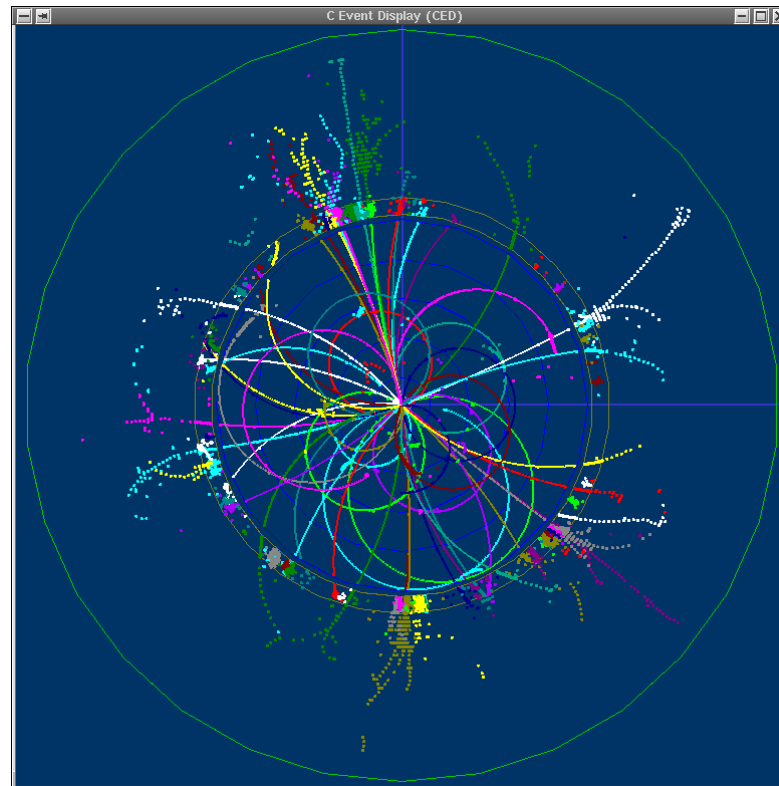
Towards detector optimisation

- ★ Both WOLF and PandoraPFA designed to work for different detector parameters / detectors !

e.g. tt event in LDC

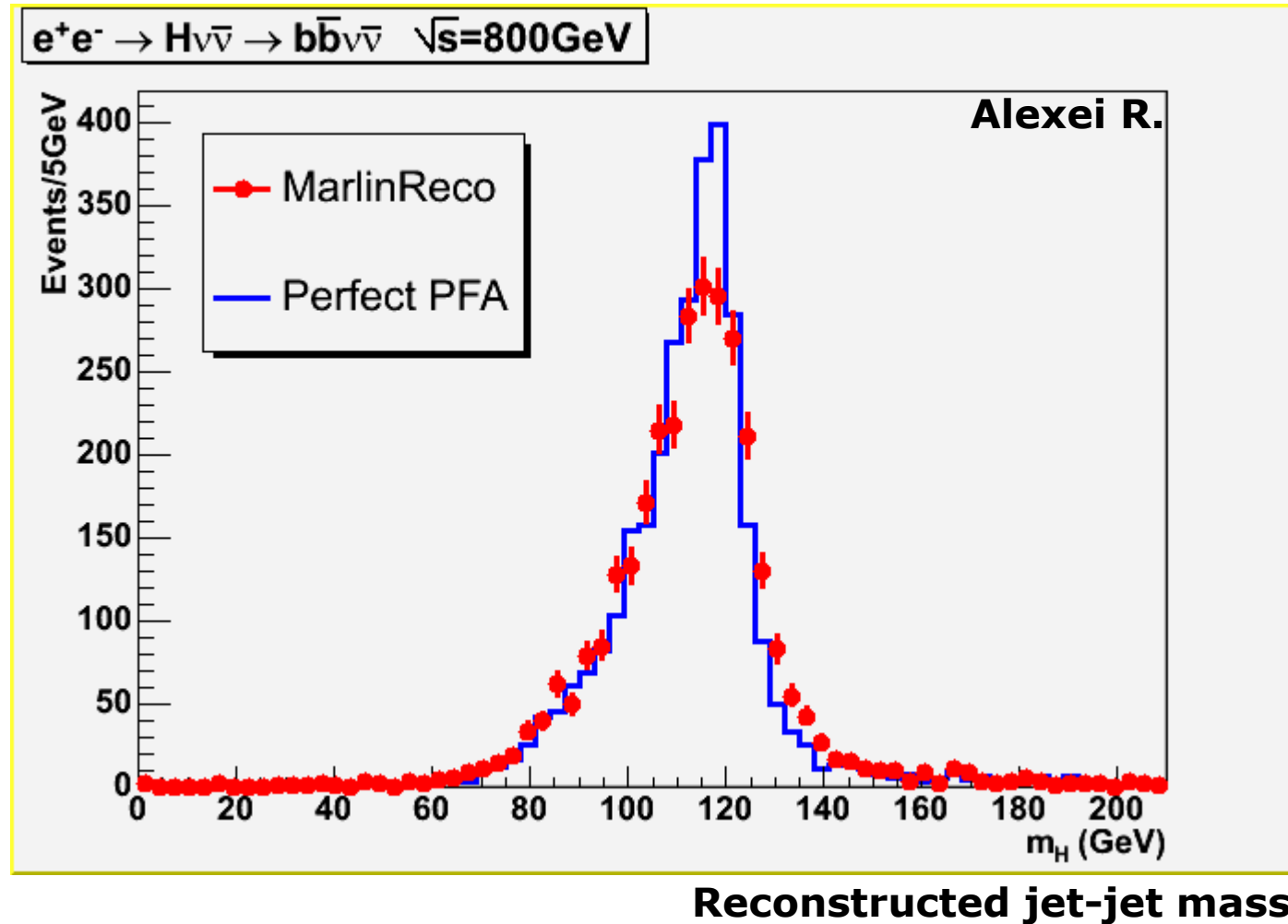


e.g. tt event in SiD



- ★ really are in a position to start optimising the ILC detector design

also possible to perform physics studies....



Conclusions

- ★ Huge amount of progress in the last year
 - ★ **MARLIN** provides a very convenient framework to “plug in” reconstruction modules
 - ★ **Realistic PFAs** now exist
 - plenty of room for development/improvement
 - ★ Can now seriously start to **optimise the ILC detector(s)**
 - ★ **THIS NEEDS CARE** – need to be sure not just seeing flaws in algorithms (Multiple Algorithms help)
 - ★ + possible to pick up off-the-shelf software and perform full- simulation physics studies
 - ★ **Need to ensure that the software development and detector optimisation/physics studies are performed in a coherent manner**
- ★ This is an excellent time to start using **MARLIN**
 - ★ It is **EASY** to get going, you can be up and running in days !