

Sensitivity of ZZ- \rightarrow llnu to Anomalous Couplings

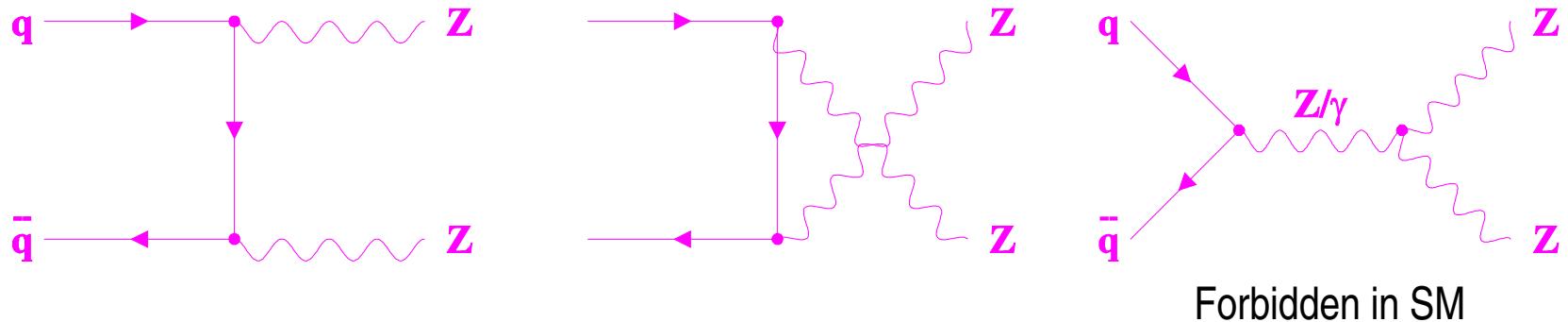
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Introduction

- Advantage of $ZZ \rightarrow ll nunu$ c.f. $ZZ \rightarrow llll$
 - Larger branching ratio: 6 times as many events before cuts
- Disadvantage
 - Very large backgrounds from $Z+jets$ and $t\bar{t}$
- Investigate sensitivity of limits on anomalous couplings to background level and systematic error on background
- This is very preliminary ‘work in progress’

Anomalous Couplings



- ZZZ and ZZg vertices forbidden in SM
- Production of on-shell ZZ probes ZZZ and ZZg anomalous couplings:
 $f4Z, f5Z, f4g f5g$
- $\text{All} = 0$ in SM

Anomalous Couplings

- f4 violate CP; helicity amplitudes do not interfere with SM; cross-sections depend on f_4^{**2} and sign cannot be determined
- f5 violate P; do interfere with SM
- Couplings depend on energy. Usual to introduce a form factor to avoid violation of unitarity:

$$f(s') = f_0 / (1 + s'/\Lambda^{**2})^{**n}$$

- Studies below use n=3, $\Lambda = 2 \text{ TeV}$
- Also assume couplings are real and only one non-zero

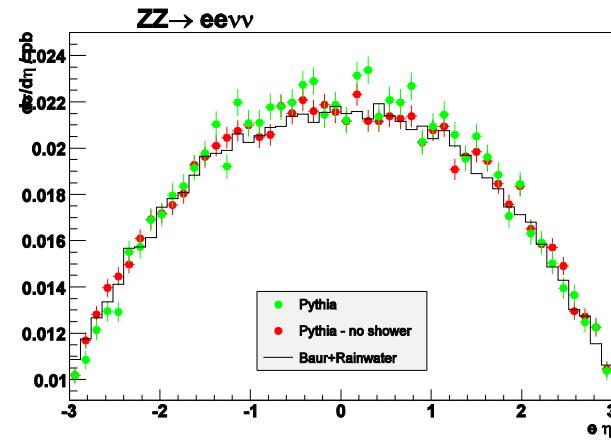
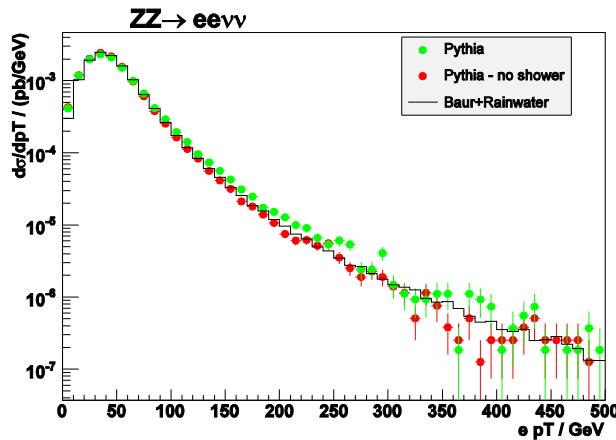
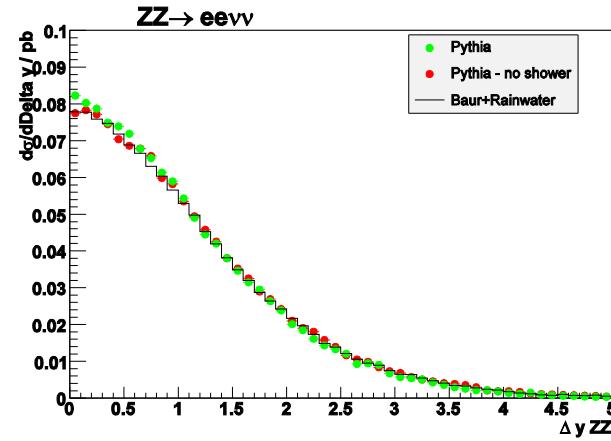
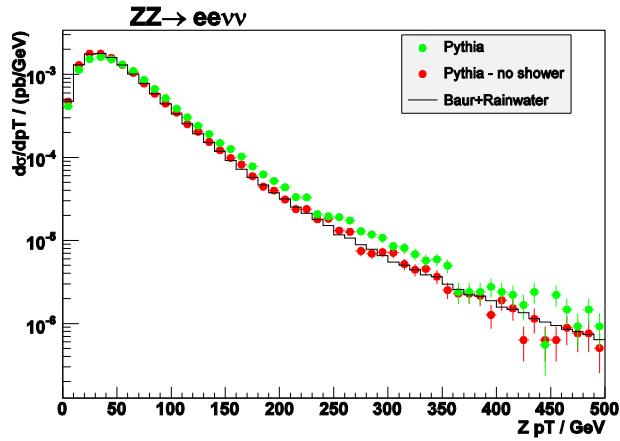
AC Monte Carlo

- Study AC using LO Monte Carlo of Baur and Rainwater
- First compare SM predictions for ZZ->eeenunu cross-section with Pythia:

CTEQ4L pdfs; $76 < m_Z < 106$ GeV

	BR	Pythia (no showers)	Pythia
No cuts	125.5 fb	126.3 fb	126.3 fb
$pT(e) > 15$ GeV $ eta(e) < 2.5$ $pT_{miss} > 50$ GeV	35.1 fb	34.5 fb	39.5 fb

Comparison with Pythia

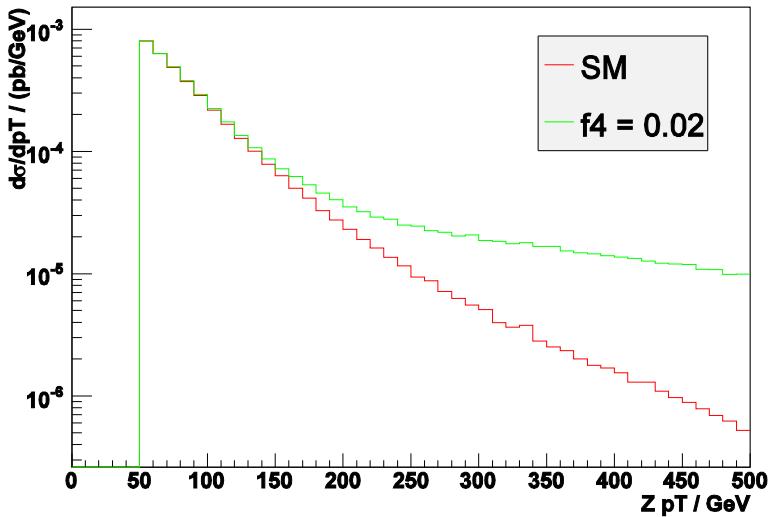


PDF dependence

- Comparisons with Pythia used out-of-date CTEQ4L because BR program used pdflib
- Have now modified it to use LHAPDF
- Following results use CTEQ6LL

pdf set	ZZ->eenunu with cuts
CTEQ4L	36.85 fb
CTEQ6LL	36.51 fb
MRST2001LO	36.90 fb

Signature of Anomalous Couplings



e.g above for ZZ->e enunu
with $pT(e) > 15$ GeV,
 $|eta(e)| < 2.5$

- Anomalous couplings produce increase in ZZ invariant mass, Z pT and lepton pT distributions
- For ZZ->llnunu can use high pT(Z) cross-section to obtain limit, or fit Z pT distribution

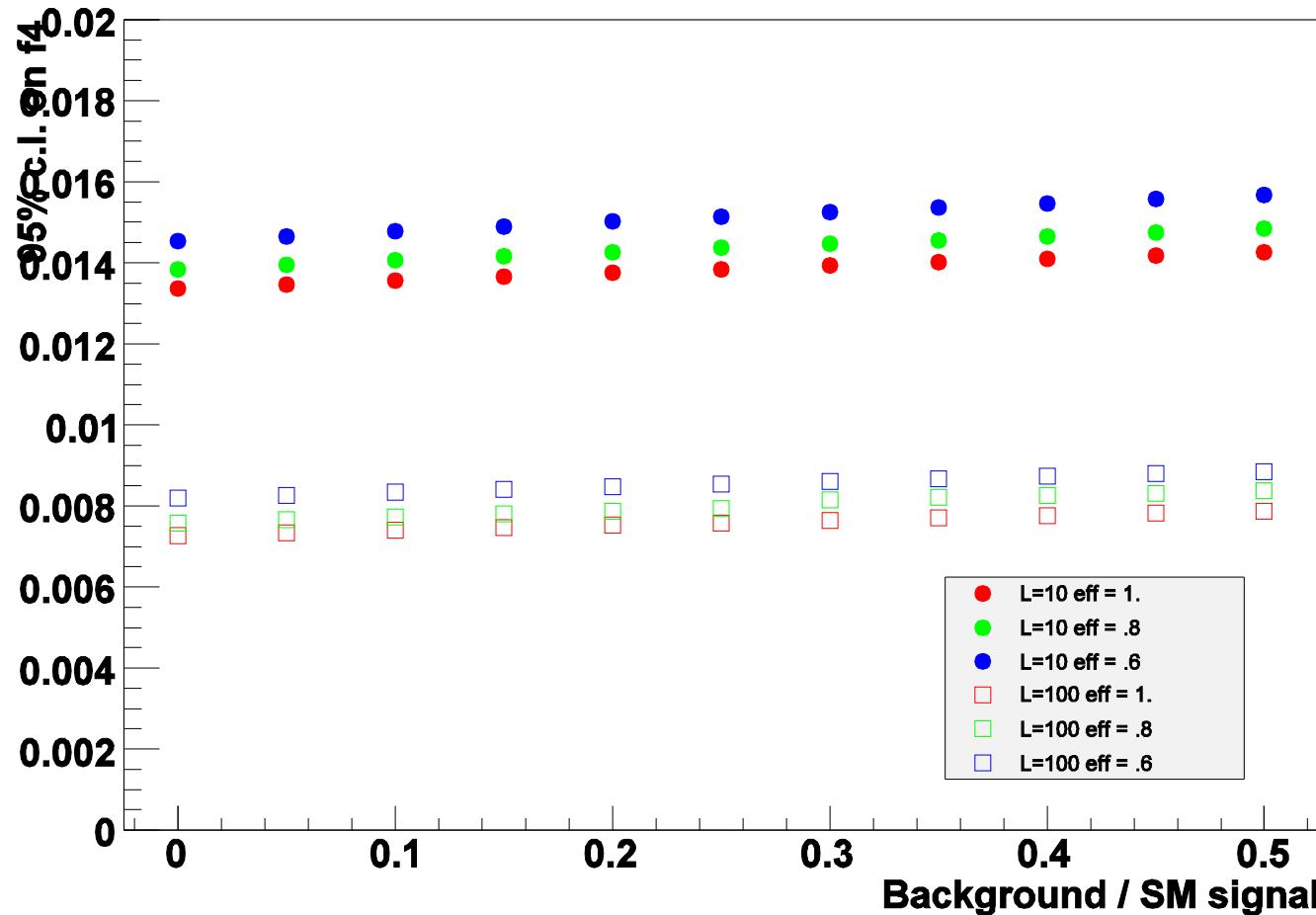
Limits from Cross-section Measurement

- First consider measurement of $ZZ \rightarrow ll\nu\nu$ cross-section for $pT(Z) > pT_{min}$
- N.B. this is LO: $pT(l) = pT_{miss}$
- Take $pT(e) > 15 \text{ GeV}$, $|eta(e)| < 2.5$, $pT_{min} = 50 \text{ GeV}$
- SM: $72.7 \text{ fb} \rightarrow 727$ signal events for 10 fb^{-1}
- Calculate cross-section, hence expected events as function of f_{4Z}
- E.g. $f_{4Z} = 0.01: 76.2 \text{ fb} \rightarrow 762$ signal events

Limits from Cross-section Measurement

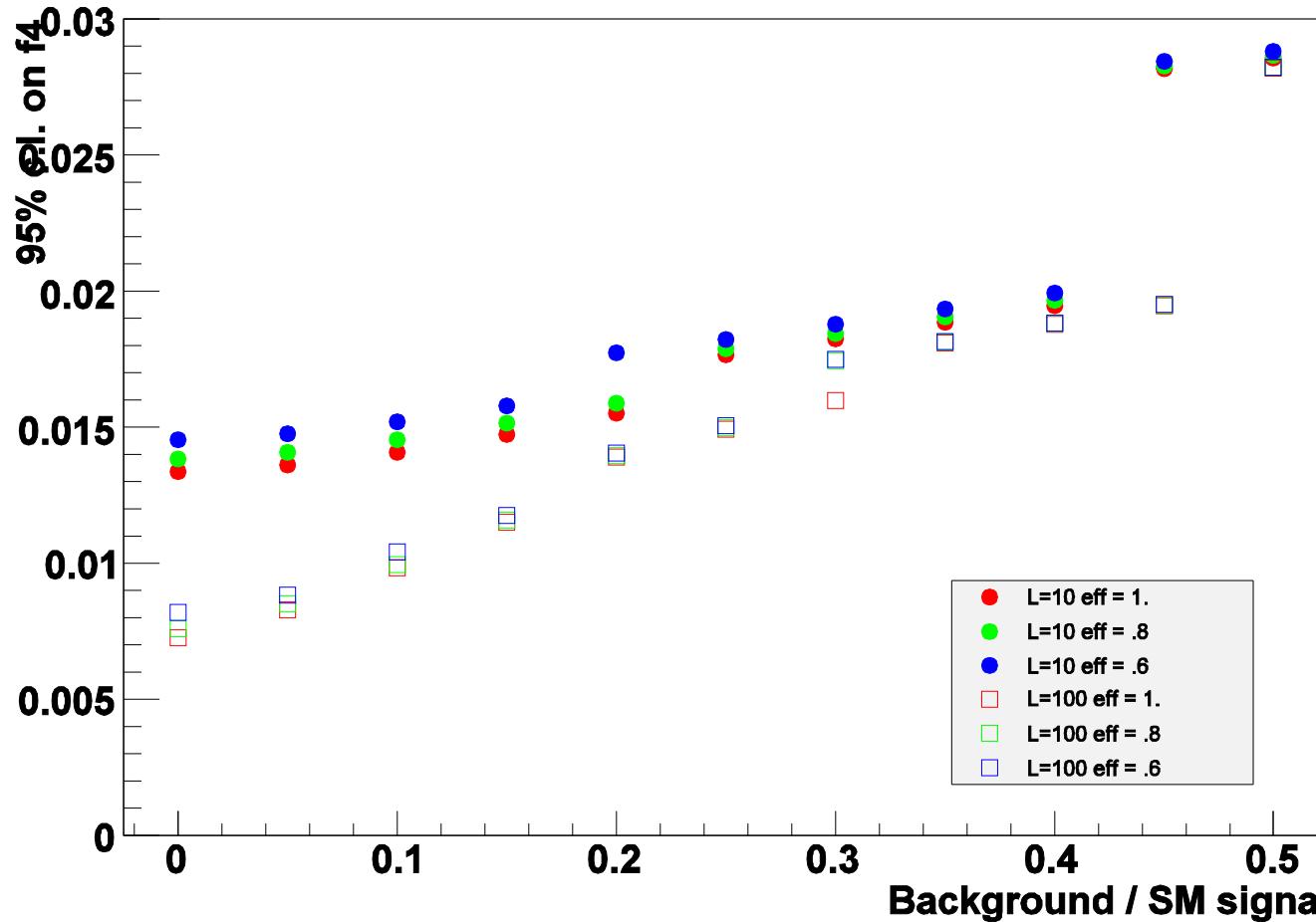
- Use chi-squared comparison between expected and ‘observed’ (=SM) numbers of events to determine 95% c.l. on coupling
(assume only one coupling non-zero)
- Calculate limit as function of ratio of background to SM signal
- First assume statistical errors only, then consider effect of a systematic error on the background

$pT_{miss} > 50 \text{ GeV}$; statistical errors only
Little dependence on background fraction



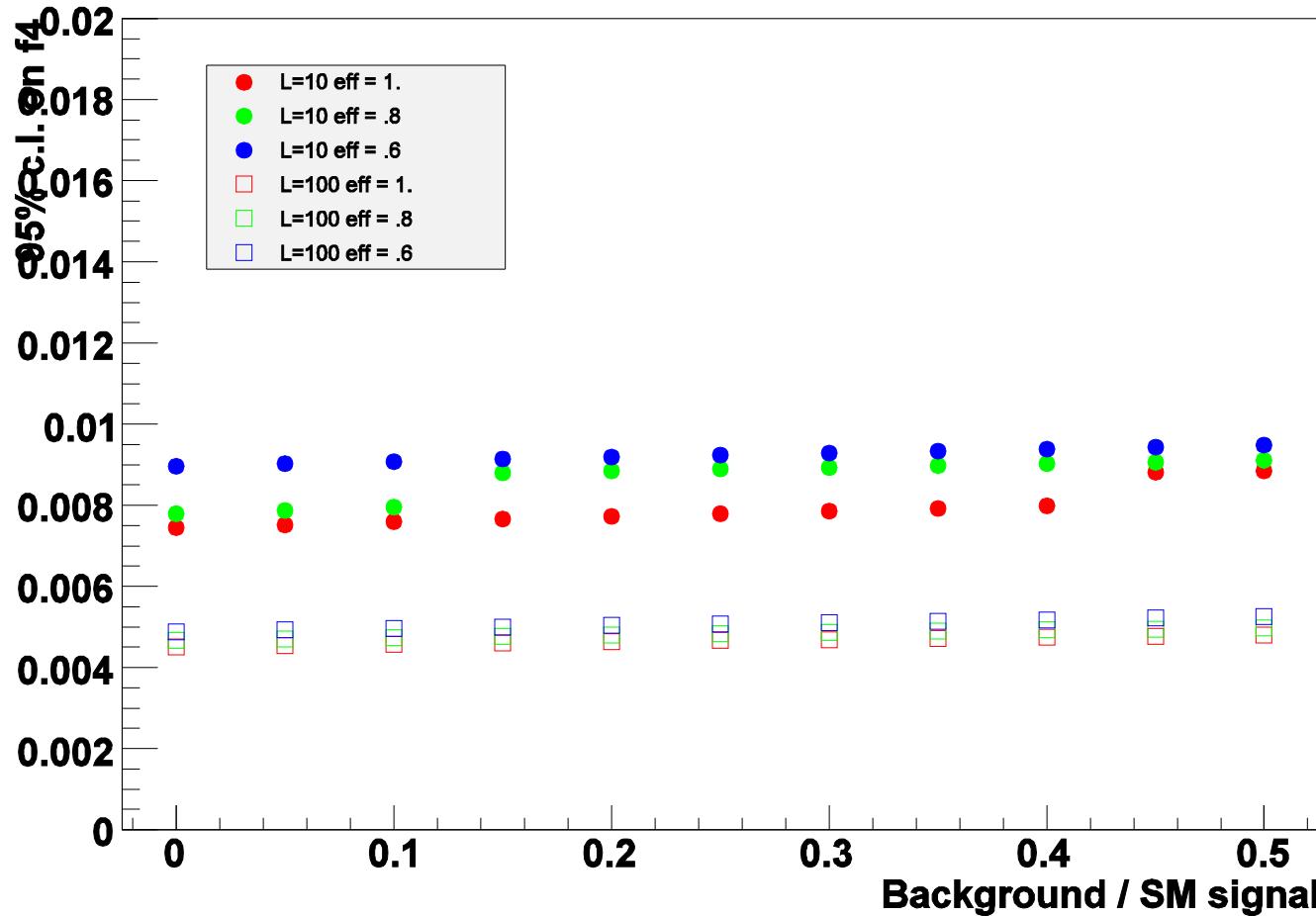
$pT_{miss} > 50 \text{ GeV}$; 20% systematic error on background

Strong dependence on background: limits independent of luminosity for high background



$pT_{miss} > 150$ GeV; statistical errors only

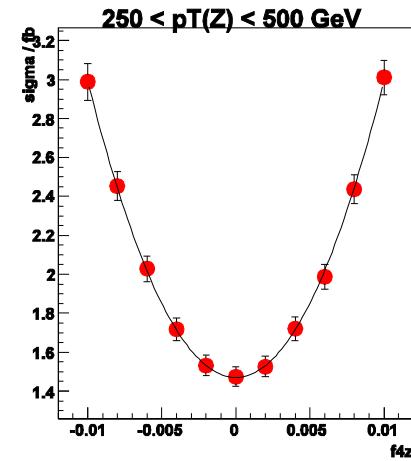
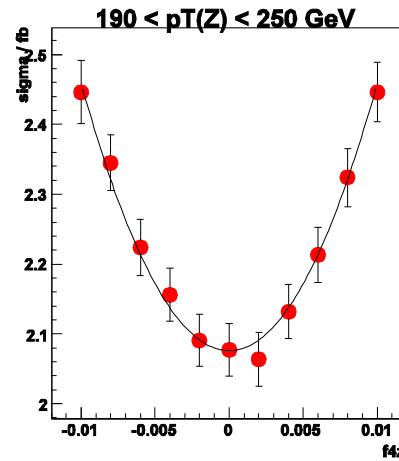
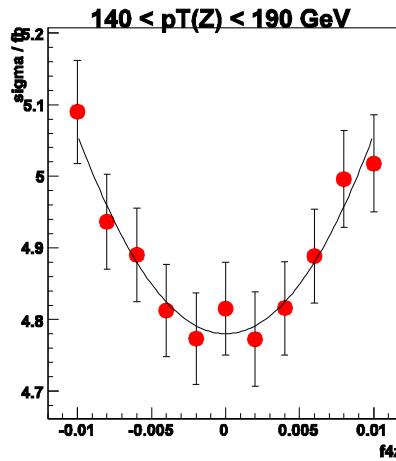
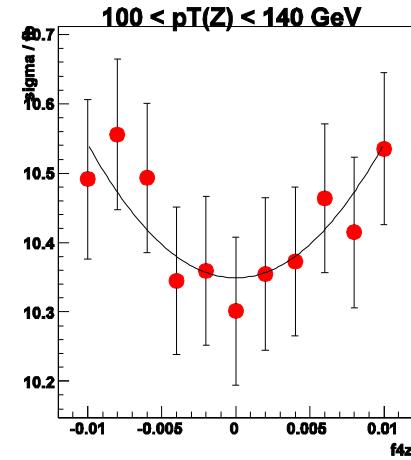
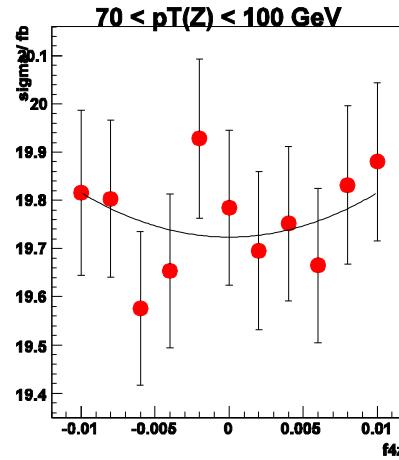
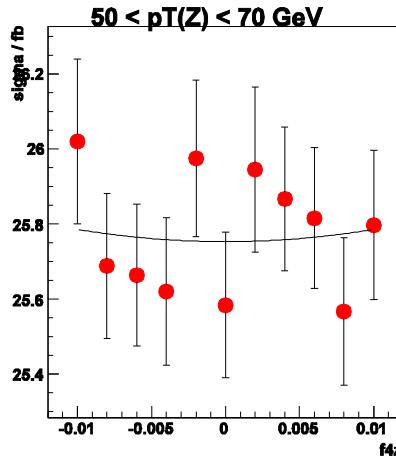
Limits much better than using $pT_{miss} > 50$ GeV



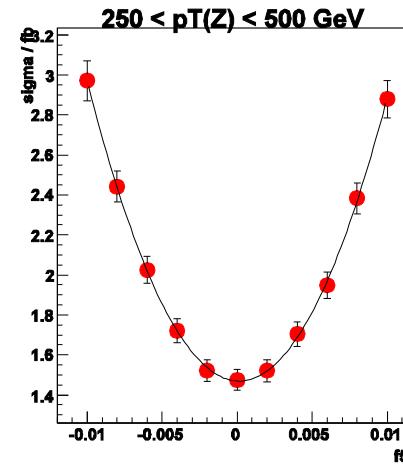
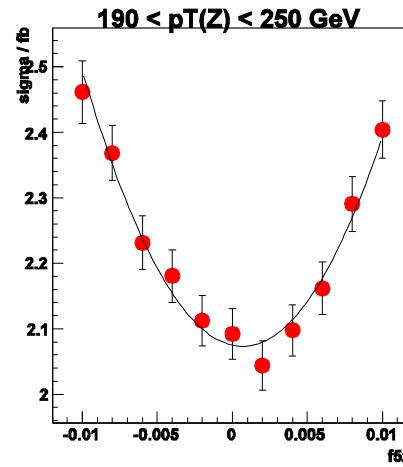
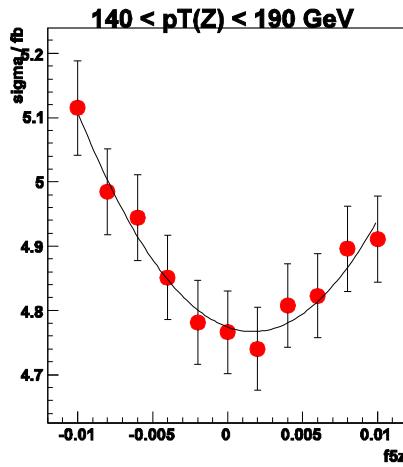
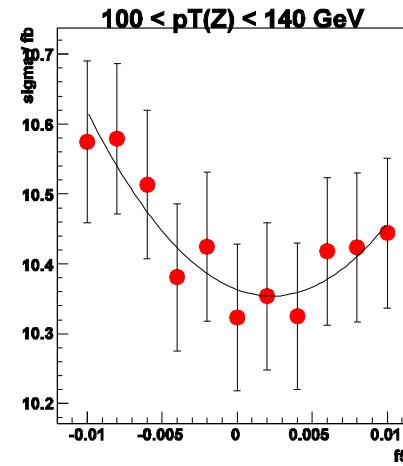
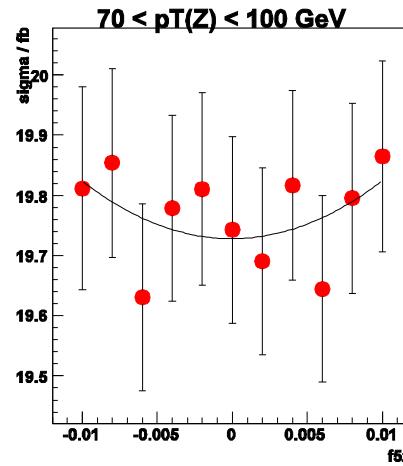
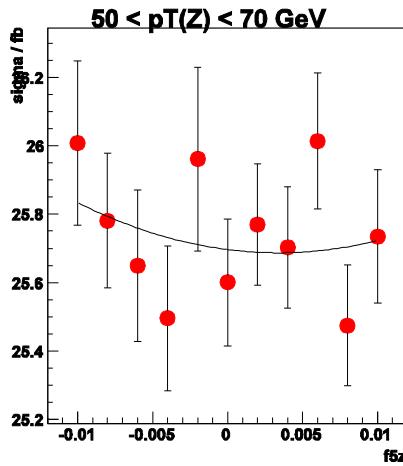
Limits from Fits to pT Distribution

- Limits from a simple cross-section measurement depend on pT cut – harder pT cut can give better limit despite much lower statistics
- Therefore better to fit pT distribution
- Results below are for ZZ->llnunu with $pT(l) > 20 \text{ GeV}$, $|eta(l)| < 2.5$
- Use BR program to generate pT distributions for several values of couplings (only one non-zero at a time)
- In each pT bin fit cross-section to quadratic in coupling to obtain distribution at arbitrary value

Cross-section v f4Z in pT bins



Cross-section v f5Z in pT bins



Limits from Fits to pT Distribution

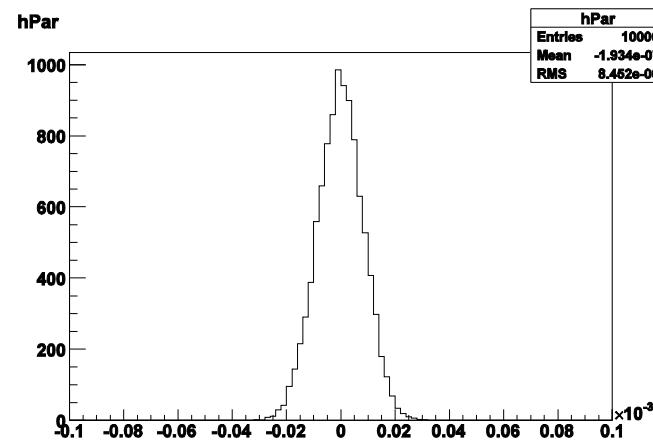
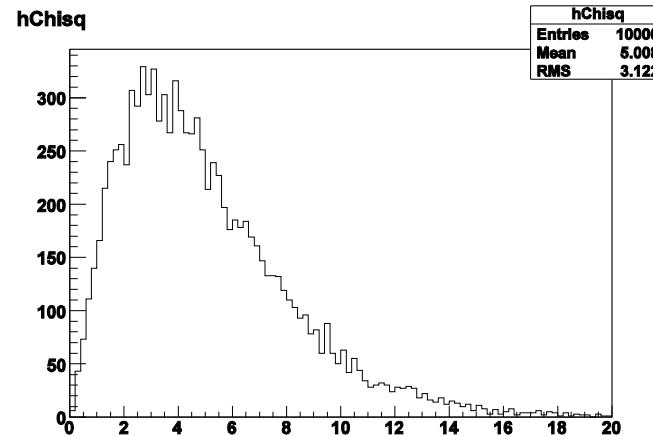
- Create ‘fake data’ sample:
 - Calculate expected SM events in each pT bin
 - Add background – constant fraction of SM
 - Apply Gaussian smearing
- Construct error matrix
 - Statistical errors plus systematic error on background assumed fully correlated
- Fit fake data sample
 - One parameter fit to $f4Z^{**2}$ or $f5Z$
 - 95 % c.l. from $X^{**2} - X^{**2\min} = 3.84$

Limits from Fits to pT Distribution

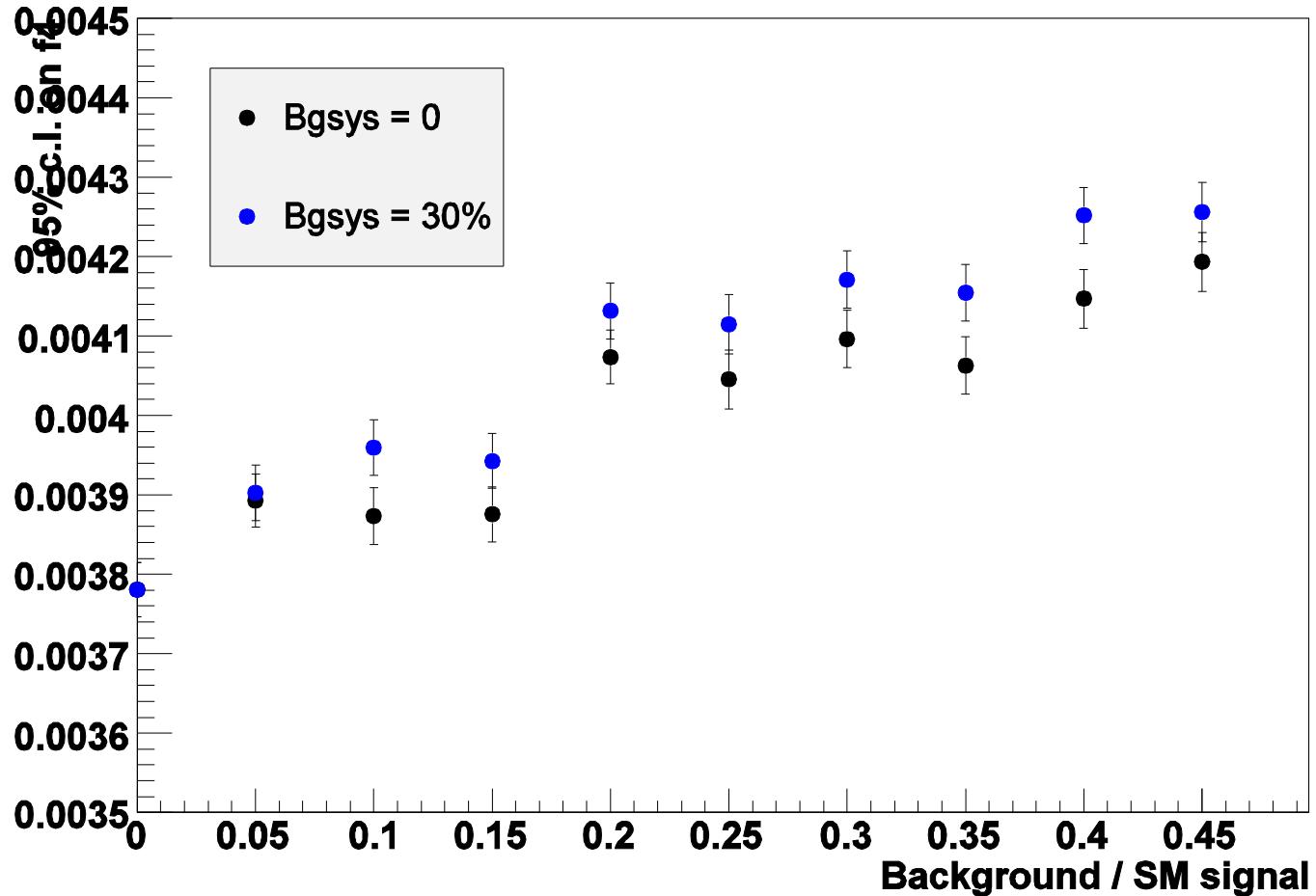
- Generate 1000 fake data samples for each value of background fraction and each value of background systematic

- Mean $X^{**2}/dof = 1$
- Mean $f4^{**2} = 0$

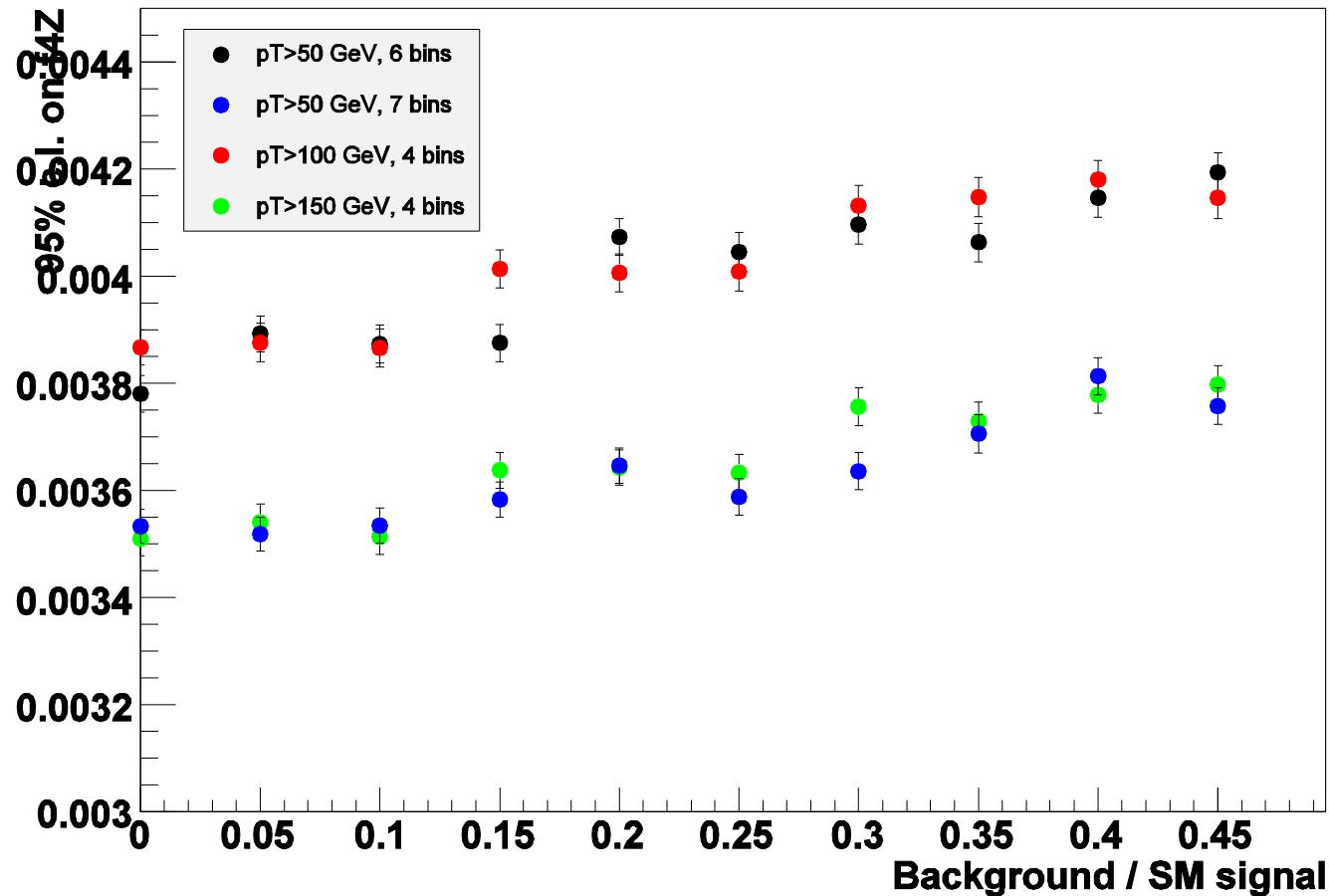
As expected



Results for 100 fb⁻¹, eff = 1.0 from Fit in Range 50 GeV < pT < 500 GeV



Results for 100 fb-1, eff = 1.0 from Different Fit Ranges and Binning



Summary so far....

- Limits worse with more background – but not dramatically so
- Limits worse with large systematic error on background – but not dramatically so
- BUT this was assuming constant background – unrealistic
- AC have little effect at low pT, so limits not too sensitive to overall normalization of pT distribution
- Limits depend on binning of pT distribution – will need to be optimized for given integrated luminosity

Next Steps

- Investigate effect of more realistic (steeply falling) background distribution
- Investigate optimal binning
- Estimate limits for expected experimental efficiency/background
- Set up framework for 2-D couplings
- Think how we are going to predict expected pT distribution (reweighting, fast MC etc.)