

Flavored Axion Models

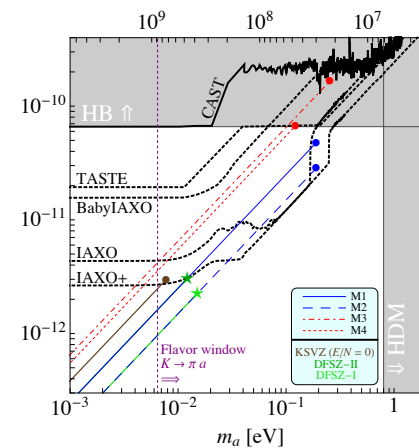


Robert Ziegler (CERN)

Outline

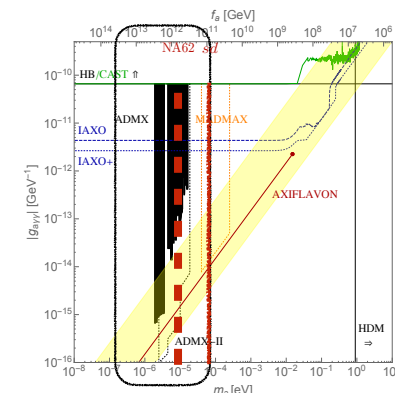
- The QCD Axion; DFSZ Models; Axion Couplings
- Non-Universal DFSZ Models: The Astrophobic Axion

arXiv: 1712.04940, with L. Di Luzio,
F. Mescia, E. Nardi, P. Panci



- Peccei-Quinn = Froggatt-Nielsen: The Axiflavoron

arXiv: 1612.08040, with L. Calibbi,
F. Goertz, D. Redigolo, J. Zupan



The Strong CP Problem

Why don't we observe CPV in strong interactions

$$\bar{\theta} \equiv \theta + \arg \det(M_u M_d) < 10^{-10} \quad ?$$

- Can impose CP and break it spontaneously such that large CKM phase induced but determinants stay real

Nelson-Barr solution [pure UV, difficult to test]

- If $\bar{\theta}$ would be dynamical field without any other potential, non-perturbative dynamics would generate potential for $\bar{\theta}(x)$ with trivial minimum

Axion solution [new ultra-light, decoupled, stable particle]

The Peccei-Quinn Mechanism

[Peccei, Quinn '77]

Introduce new global $U(1)_{PQ}$ symmetry with fermion charges such that have QCD anomaly

Break global $U(1)_{PQ}$ symmetry spontaneously at scale f by vev of complex scalar field Φ

The diagram shows the equation for the complex scalar field Φ within a rectangular box. The equation is
$$\Phi = \frac{f + \phi(x)}{\sqrt{2}} e^{ia(x)/f}$$
 Three blue arrows point to different parts of the equation: one from the text 'U(1)_{PQ} breaking scale' to the f in the numerator; one from the text 'radial mode' to the $\phi(x)$ in the numerator; and one from the text 'axion' to the $a(x)$ in the exponent.

U(1)_{PQ} breaking scale

radial mode

axion

$$\Phi = \frac{f + \phi(x)}{\sqrt{2}} e^{ia(x)/f}$$

$U(1)_{PQ}$ non-linearly realized as shift symmetry of axion

The Peccei-Quinn Mechanism

Effective Lagrangian at scales $\ll f$ contains only Goldstone boson $a(x)$, all other fields take mass at f

$$\mathcal{L}_{\text{eff}} = \underbrace{\bar{\theta} \frac{\alpha_s}{8\pi} G_a^{\mu\nu} \tilde{G}_{a,\mu\nu}}_{\text{Interactions with SM fermions}} + \underbrace{\mathcal{L}_{\text{a,int}}\left[\frac{\partial_\mu a}{f}, \psi_{\text{SM}}\right]}_{\text{ABJ term for QCD anomaly}} + \underbrace{\frac{a}{f} \xi \frac{\alpha_s}{8\pi} G_a^{\mu\nu} \tilde{G}_{a,\mu\nu} + \mathcal{L}_{\text{anom}}\left[\frac{a}{f} F \tilde{F}\right]}_{\text{ABJ term for other anomalies}}$$

[respects shift symmetry] [breaks shift symmetry] [cf. $\frac{\pi^0}{f_\pi} F \tilde{F}$]

Depends on $\bar{\theta}$ only through $\frac{\bar{a}(x)}{f_a} \equiv \bar{\theta} + \frac{a(x)}{f} \xi$

[have essentially made $\bar{\theta}$ dynamical field]

The Axion

Effective Lagrangian induces axion potential

$$V_{\text{eff}} = -\frac{\bar{a}(x)}{f_a} \frac{\alpha_s}{8\pi} G_a^{\mu\nu} \tilde{G}_{a,\mu\nu} \xrightarrow[\text{effects}]{\text{non-PT}} V(a) \sim -m_\pi^2 f_\pi^2 \left| \cos \frac{\bar{a}(x)}{f_a} \right|$$

Potential minimized at CP-conserving point $\langle \bar{a}(x) \rangle = 0$

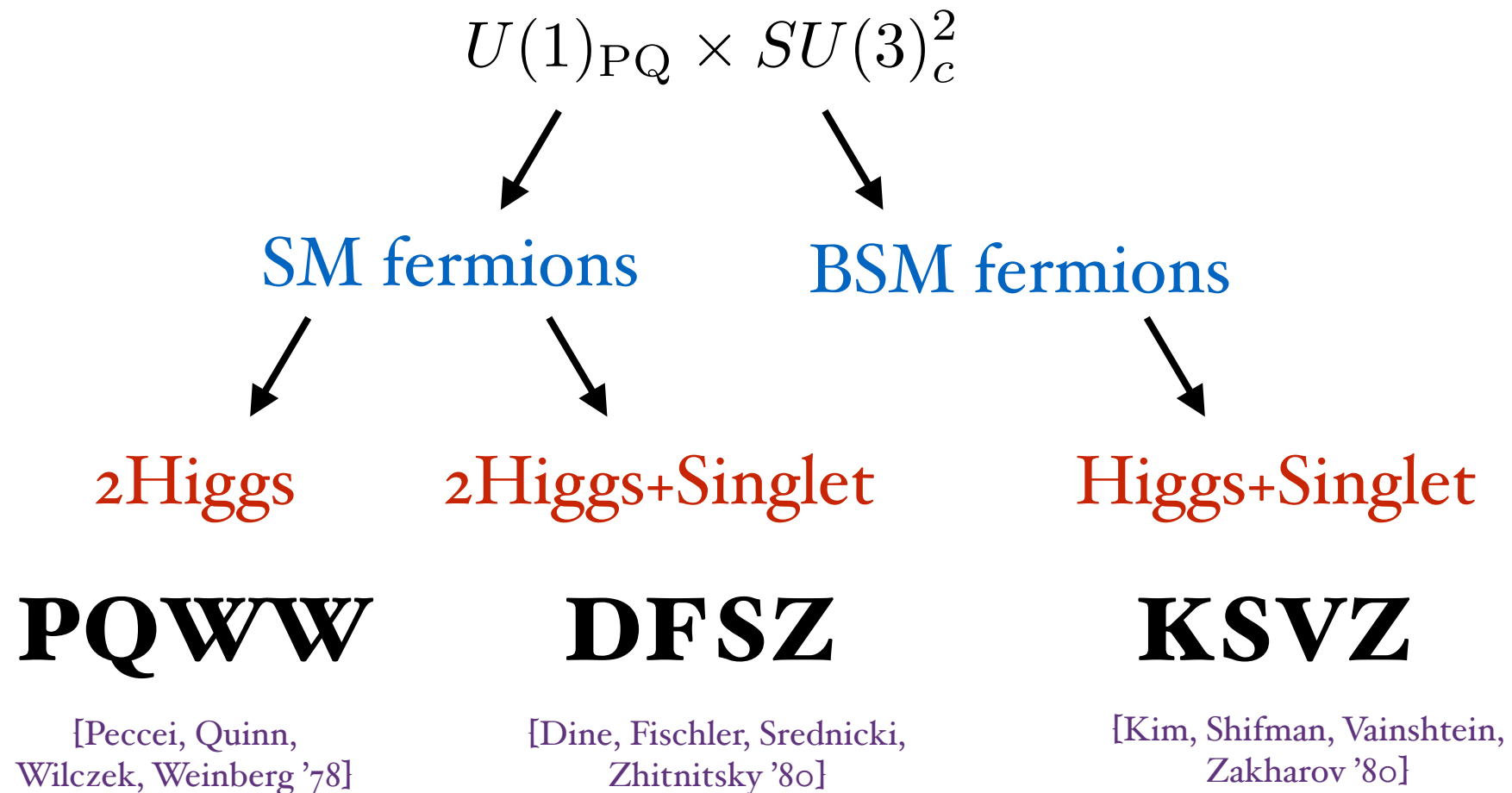
QCD θ -term dynamically relaxed to zero

Axion gets mass $m_a \sim m_\pi f_\pi / f_a$ and couples
to photons and SM fermions $\propto 1/f_a \propto m_a$

For large f_a axion is ultra-light and decoupled

Axion Models

Specify **anomalous** breaking of PQ (**fermion sector**)
and **spontaneous** breaking of PQ (**scalar sector**)



excluded

$J/\psi \rightarrow \gamma a$
see however 1710.03764

$\langle \text{Singlet} \rangle \gg v$: “Invisible” axion models

DFSZ Models

$$\text{SM fermions} + \text{2Higgs} + \text{Singlet} \left\{ \begin{array}{l} \langle H_1 \rangle = c_\beta v \quad \langle H_2 \rangle = s_\beta v \\ \langle \Phi \rangle = v_{\text{PQ}} \gg v \end{array} \right.$$

Construct 2HDM Lagrangian invariant under single $U(1)$

$$\mathcal{L}_{\text{yuk}} = y_{ij}^u \bar{Q}_i U_j \left\{ \begin{array}{l} H_1 \\ H_2 \end{array} \right. \xrightarrow[\text{U(1) charges}]{\text{flavor-universal}} \begin{array}{l} \bar{Q}_i U_j H_1 \\ \bar{Q}_i D_j \tilde{H}_2 \\ \bar{L}_i E_j \tilde{H}_{1 \text{ or } 2} \end{array}$$

Break residual $U(1)$ by H-Singlet
couplings $\mathcal{L} \sim H_1^\dagger H_2 \Phi^2$

Ensure $GB_{\text{PQ}} \perp GB_Z$
 $0 = X_1 c_\beta^2 + X_2 s_\beta^2$

All PQ charges fixed in terms of vacuum angle

Axion Couplings

Higgs contain GB: axion-fermion couplings from Yukawas

$$H_i \sim v_i e^{i\textcolor{red}{X}_i a(x)/v_{\text{PQ}}}$$

Can remove by PQ transformation acting only on fermions

$$f \rightarrow f e^{i\textcolor{blue}{X}_f a(x)/v_{\text{PQ}}}$$

Anomalous: generate axion couplings to field strengths

$$\mathcal{L}_{\text{anon}} \sim N \frac{a}{v_{\text{PQ}}} \frac{\alpha_s}{4\pi} G\tilde{G} + E \frac{a}{v_{\text{PQ}}} \frac{\alpha_{\text{em}}}{4\pi} F\tilde{F}$$

Local: generate derivative axion couplings to fermions

$$\mathcal{L}_{\text{ferm}} \sim \frac{\partial_\mu a}{v_{\text{PQ}}} \bar{u}_i \gamma^\mu P_R \textcolor{blue}{X}_{u_i} u_i + \dots$$

$$[\text{mass basis: } \textcolor{blue}{X}_u \rightarrow V_{UR}^\dagger \textcolor{blue}{X}_u V_{UR}]$$

DFSZ Predictions

Anomaly coefficients only depend on charge differences

$$\left. \begin{aligned} 2N &= \text{Tr} (X_{q_L} - X_{q_R}) = 3 (X_1 - X_2) \\ E &= \dots \dots = 4X_1 - X_2 - 3X_{1 \text{ or } 2} \end{aligned} \right\} \begin{array}{l} \rightarrow 6 \\ \rightarrow 8 \text{ or } 2 \end{array}$$

Fermion couplings flavor-universal & depend only on t_β

$$X_u = -\underbrace{X_1}_{2s_\beta^2} \mathbb{1}_{3 \times 3} \quad X_d = \underbrace{X_2}_{-2c_\beta^2} \mathbb{1}_{3 \times 3} \quad X_e = X_{1 \text{ or } 2} \mathbb{1}_{3 \times 3}$$

Also u+d and u+e sums depend only on charge differences

$$X_u + X_d \propto X_1 - X_2 \quad X_u + X_e \propto X_1 - X_{1 \text{ or } 2}$$

[standard normalization is $v_{\text{PQ}} = 2N f_a$]

DFSZ Predictions

Anomaly coefficients only depend on charge differences

$$\left. \begin{aligned} 2N &= \text{Tr} (X_{q_L} - X_{q_R}) = 3 (X_1 - X_2) \\ E &= \dots \dots = 4X_1 - X_2 - 3X_{1 \text{ or } 2} \end{aligned} \right\} \begin{aligned} &\rightarrow 6 \\ &\rightarrow 8 \text{ or } 2 \end{aligned}$$

Fermion couplings flavor blind & depend only on t_β

$$X_u = -X_1 \mathbb{1}_{3 \times 3} \quad X_e = X_{1 \text{ or } 2} \mathbb{1}_{3 \times 3}$$

**only 2 parameters!
(& 1 discrete choice)**

Also u+d and e sums depend only on charge differences

$$X_u + X_d \propto X_1 - X_2 \quad X_u + X_e \propto X_1 - X_{1 \text{ or } 2}$$

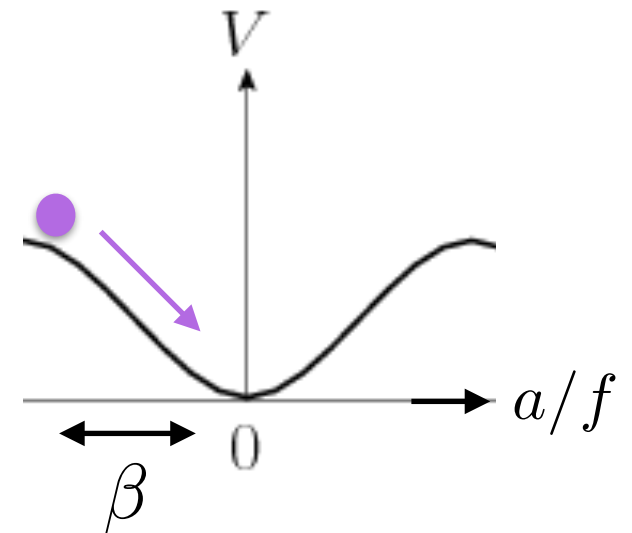
[standard normalization is $v_{\text{PQ}} = 2N f_a$]

Axions as Dark Matter

[axion essentially stable for $m_a \lesssim 20 \text{ eV}$]

When PQ breaking before inflation axion can be dark matter through “misalignment mechanism”

At QCD phase transition axion starts oscillating around minimum:
energy stored in oscillations
contributes to DM relic density



$$\Omega_{\text{DM}} h^2 \approx 0.1 \left(\frac{10^{-5} \text{ eV}}{m_a} \right)^{1.18} \beta^2$$

Correct abundance for
 $10^{-7} \text{ eV} \lesssim m_a \lesssim 10^{-4} \text{ eV}$

Axion Searches

Axion searches typically rely on axion-photon coupling

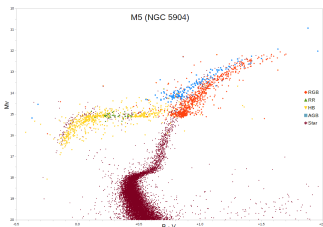
$$\mathcal{L} \supset g_{a\gamma\gamma} a \vec{E} \cdot \vec{B} \quad g_{a\gamma\gamma} \sim \frac{E}{N} \frac{1}{10^{16} \text{GeV}} \frac{m_a}{\mu\text{eV}}$$

Haloscopes: DM axion $\xrightarrow{\text{static B-field}}$ photon $E_\gamma \sim m_a$
ADMX
resonantly detected by
microwave cavities of size $1/m_a$

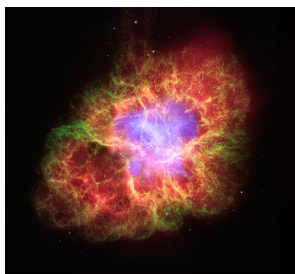
Helioscopes: Solar photon $\xrightarrow{\text{nuclear E-field}}$ axion $E_a \sim \text{keV}$
CAST/IAXO
converted back to X-ray
photons with static B-fields

Astrophysical Bounds

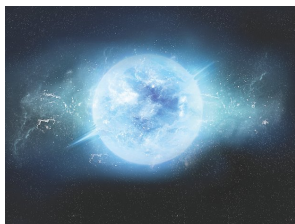
Large axion couplings to matter would allow to radiate off energy in astrophysical objects



Evolution of Horizontal Branch stars: $m_a < \frac{3 \cdot 10^{-1} \text{ eV}}{C_\gamma}$
constrain photon couplings

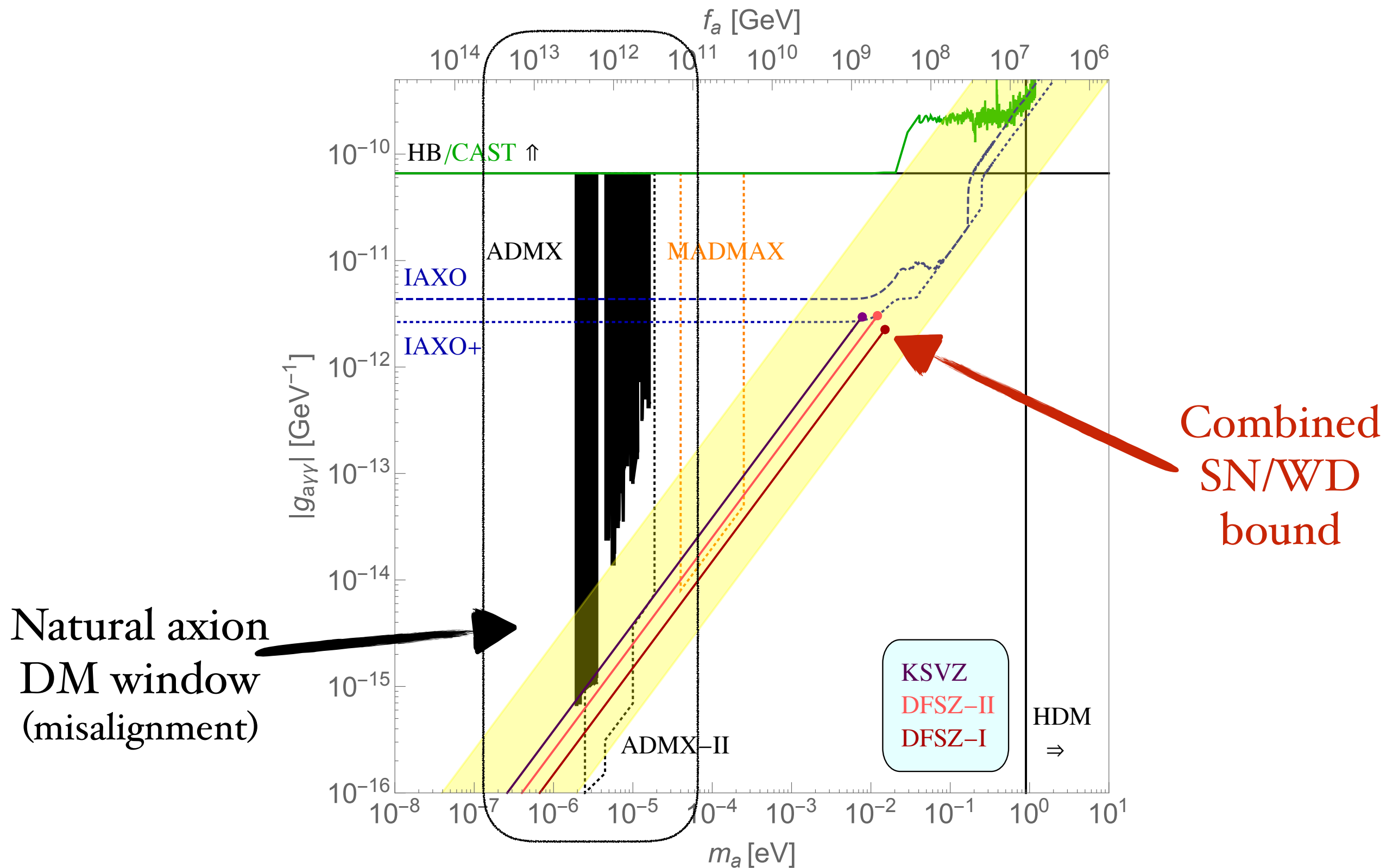


Supernova neutrino burst duration: $m_a < \frac{4 \cdot 10^{-3} \text{ eV}}{|C_N|}$
constrain nucleon couplings



White Dwarf cooling: $m_a < \frac{3 \cdot 10^{-3} \text{ eV}}{|C_e|}$
constrain electron couplings

Status of Axion Searches



Flavored Axions

Tree-level flavor-violating axion couplings are strongly constrained

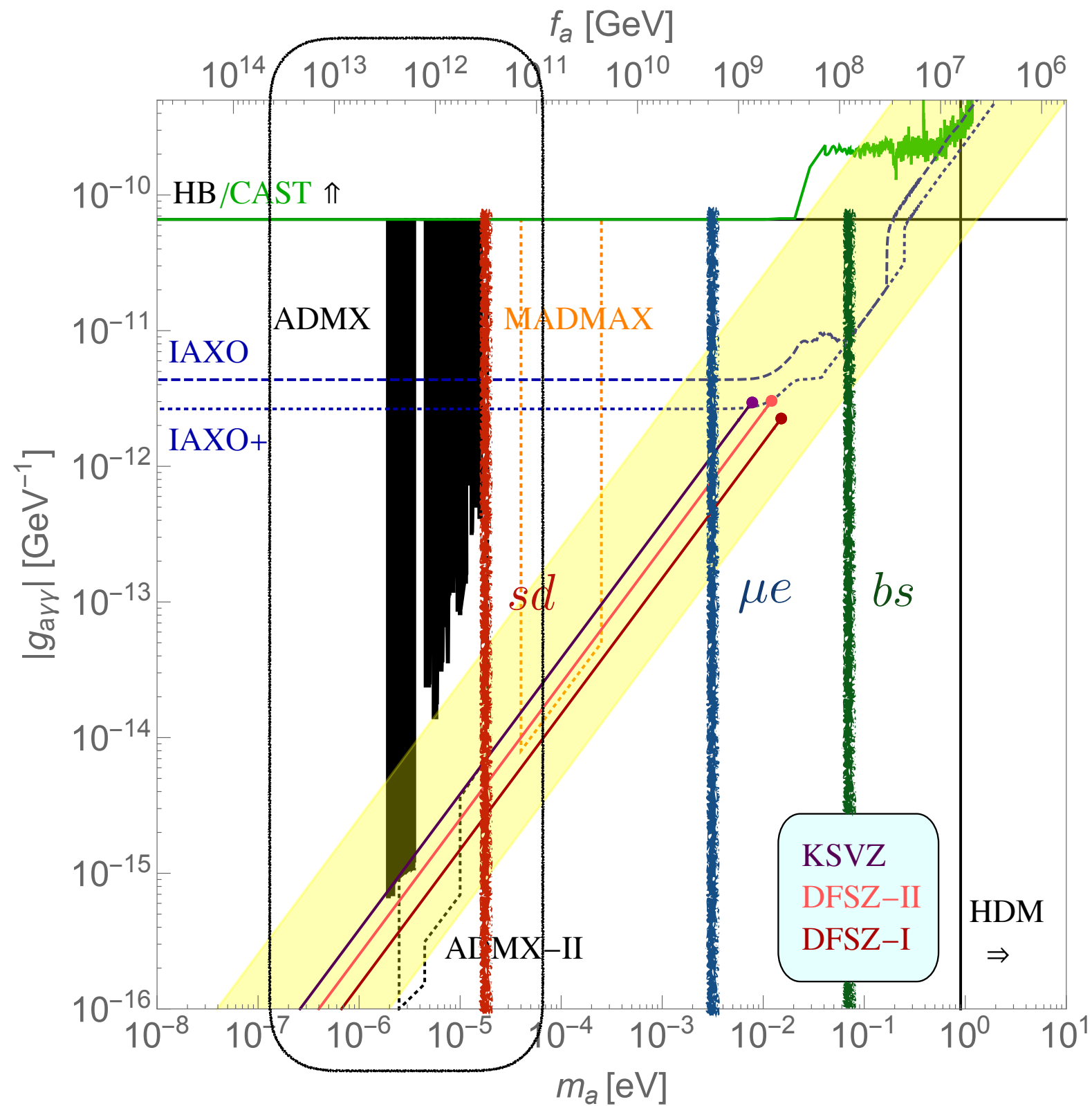
$$\frac{\partial_\mu a}{2f_a} \bar{f}_i \gamma^\mu (C_{ij}^V + C_{ij}^A \gamma_5) f_j$$

$$\begin{array}{lll} \mu \rightarrow ea\gamma & m_a < \frac{3 \cdot 10^{-3} \text{ eV}}{|C_{\mu e}|} & \xrightarrow{?} \text{MEG, Mu3e} \\ \text{(Crystal Box, '88)} & & \end{array}$$

$$\begin{array}{lll} K \rightarrow \pi a & m_a < \frac{2 \cdot 10^{-5} \text{ eV}}{|C_{sd}^V|} & \xrightarrow{\times 1/8} \text{NA62} \\ \text{(E787+E949, '08)} & & \end{array}$$

$$\begin{array}{lll} B \rightarrow Ka & m_a < \frac{9 \cdot 10^{-2} \text{ eV}}{|C_{bs}^V|} & \xrightarrow{\times 1/10} \text{BELLE II} \\ \text{(CLEO, '01)} & & \end{array}$$

Flavor Bounds



(for $C_{i \neq j} = 1$)

Flavor Violation in DFSZ

In DFSZ flavor-violating axion couplings vanish at tree-level since Yukawas and PQ aligned

$$[\text{PQ}_u, Y_u^\dagger Y_u] = 0$$

Flavor Violation loop&CKM suppressed (MFV)

$$C_{sd}^V \sim C_t \frac{y_t^2}{16\pi^2} V_{ts}^* V_{td} \sim C_t \frac{y_t^2}{16\pi^2} \lambda^5 \longrightarrow m_a \lesssim 10 \text{ eV}$$

Flavored Axion Models

In general get flavor-violating axion couplings at tree-level from misalignment of PQ charges and Yukawas

$$[\text{PQ}_u, Y_u^\dagger Y_u] \neq 0$$

Here discuss two examples:

2+1 DFSZ

misalignment is
free parameter

PQ = FN

misalignment from
fermion masses

Non-universal DFSZ Models

Generalized DFSZ-type models with 2Higgs
PQ charges universal for two generations

$$X_f = \text{diag}(X_1, X_1, X_3) \rightarrow V_f^\dagger X_f V_f = X_1 \delta_{ij} + (X_3 - X_1) \epsilon_{ij}^f$$

$$\epsilon_{ij}^f \equiv (V_f)_{i3}^* (V_f)_{j3} \quad f = u_L, u_R, d_L, d_R$$

Depend on 2 misalignment parameters in each sector

$$0 \leq \epsilon_{ii}^f \leq 1$$

$$\sum_i \epsilon_{ii}^f = 1$$

$$C_{ii}^f = X_1 + (X_3 - X_1) \epsilon_{ii}^f$$

$$|C_{i \neq j}^f| = |X_3 - X_1| \sqrt{\epsilon_{ii}^f \epsilon_{jj}^f}$$

Can have large flavor violation in all sectors!

Nucleophobic DFSZ Models

Can e.g. use to construct **nucleophobic** axion models

$$C_p + C_n = 0.50(5) (C_u + C_d - 1) - 2\delta \quad C_p - C_n = 1.273(2) \left(C_u - C_d - \frac{1-z}{1+z} \right)$$



UV couplings



$G\tilde{G}$



RG contrib from heavy
quark couplings $\sim 5\%$



isospin breaking contrib
 $z = m_u/m_d \approx 1/2$

Nucleon (and pion) couplings suppressed for

$$C_u + C_d = 1 \quad C_u = 1/(1+z) \approx 2/3$$

The SN bound relaxed by factor ~ 20 ;
can relax also WD bound when $C_e \ll 1$

An Astrophobic Axion Model

$$\mathcal{L} \sim \frac{\bar{f}_{L3} f_{R3}}{\bar{f}_{L3} f_{Ra}} \begin{cases} h_1 & \textcolor{red}{u} \\ \tilde{h}_2 & \textcolor{blue}{d} \\ \tilde{h}_1 & \textcolor{green}{e} \end{cases} + \frac{\bar{f}_{La} f_{Rb}}{\bar{f}_{La} f_{R3}} \begin{cases} h_2 & \textcolor{red}{u} \\ \tilde{h}_1 & \textcolor{blue}{d} \\ \tilde{h}_2 & \textcolor{green}{e} \end{cases}$$

Only PQ charges of RH fields universal: couplings depend on misalignment parameters (LH rotations)

$$\begin{aligned} C_{u_i u_i}^A &= c_\beta^2 - \epsilon_{ii}^{u_L}, & C_{u_i u_j}^{V,A} &= \pm \epsilon_{ij}^{u_L} \\ C_{d_i d_i}^A &= s_\beta^2 - \epsilon_{ii}^{d_L}, & C_{d_i d_j}^{V,A} &= \pm \epsilon_{ij}^{d_L} \\ C_{e_i e_i}^A &= -c_\beta^2 + \epsilon_{ii}^{e_L}, & C_{e_i e_j}^{V,A} &= \mp \epsilon_{ij}^{e_L}, \end{aligned}$$

Assume at most CKM like rotations in LH quark sector

$$\epsilon_{11}^{u_L, d_L} \lesssim \lambda^6, \quad \epsilon_{22}^{u_L, d_L} \lesssim \lambda^4, \quad \epsilon_{33}^{u_L, d_L} \simeq 1$$

$$\epsilon_{ij}^{u_L} \equiv (V_{UL})_{i3}^* (V_{UL})_{j3}$$

Astrophobic Axion Models

$$C_u = c_\beta^2 \quad C_d = s_\beta^2$$



nucleophobic for
 $c_\beta^2 \simeq 2/3$ (5% tuning)

$$C_e = -c_\beta^2 + \epsilon_{11}^{eL}$$



electrophobic for
 $\epsilon_{11}^{eL} \simeq 2/3$ (5% tuning)

Note that in 3HDM 2/3 of **astrophobic** conditions

$$C_u + C_d = 1$$

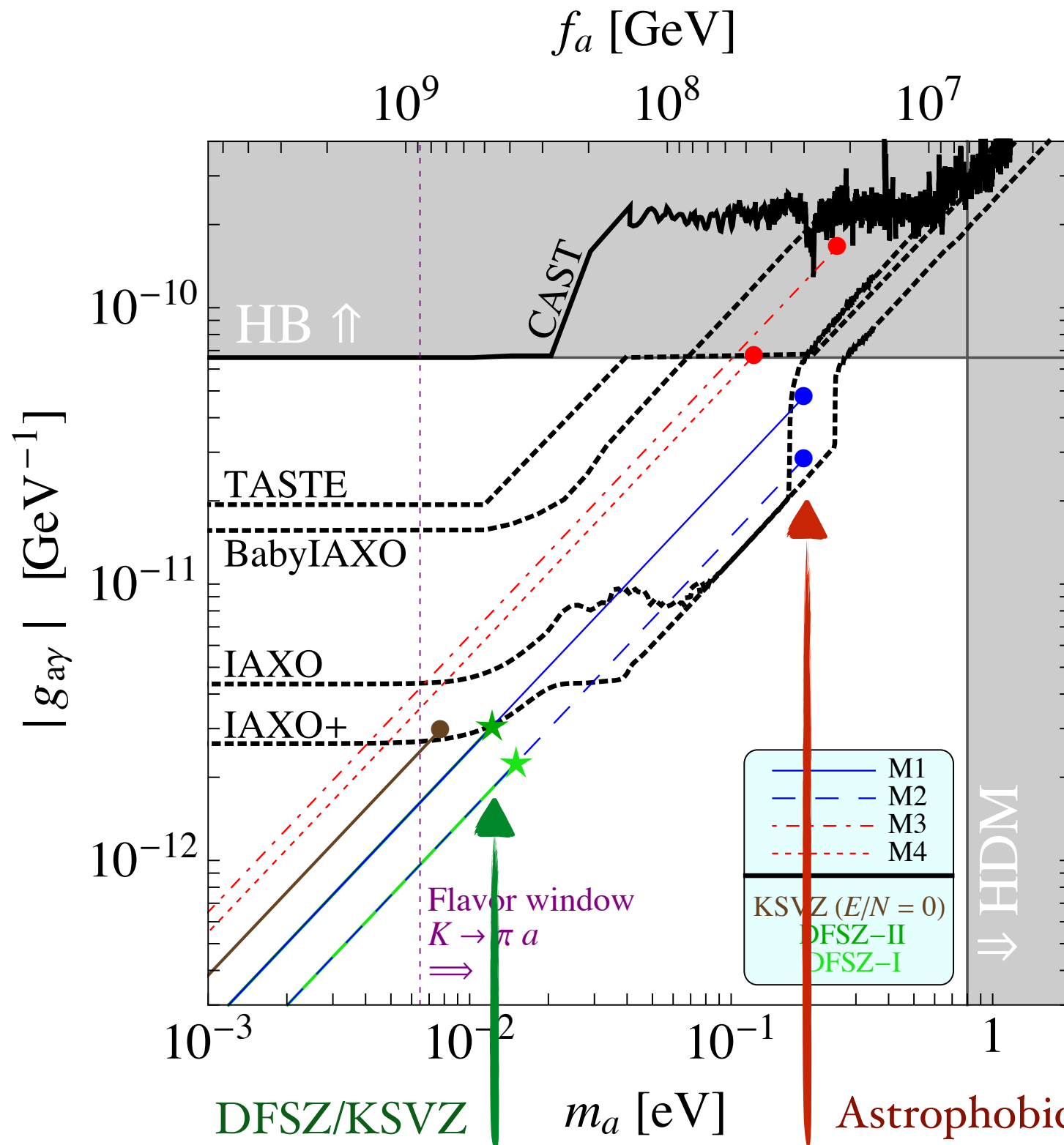
$$C_u + C_e = 2/3$$

can be realized by charge assignment (no tuning), while

$$C_u \approx 2/3$$

only with adjustment of vacuum alignment angle

Astrophobic Phenomenology



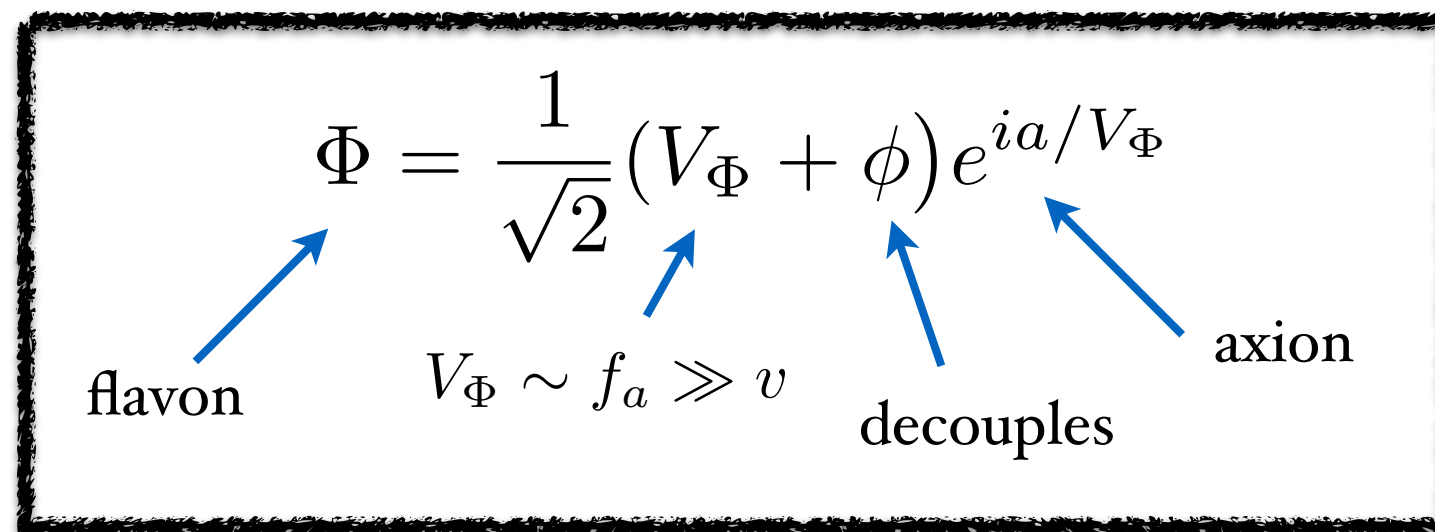
Get 4 astrophobic 2HDM models with different E/N

- Axion can be **as heavy as 0.2 eV**
- Testable at IAXO and e.g. NA62
- Predict $\mathcal{B}_{\tau \rightarrow ea} = 6.7 \cdot 10^{-6} \left(\frac{m_a}{0.2 \text{ eV}} \right)^2$
- Can fit stellar cooling anomalies

PQ as a Flavor Symmetry

PQ could be subgroup of flavor symmetry
that explains Yukawa hierarchies [F. Wilczek '82]

Smallness of Yukawas arise from spontaneously
broken global flavor symmetry, associated
Goldstone can be identified with QCD axion



The diagram shows the decomposition of a complex scalar field Φ into its magnitude and phase. The equation is $\Phi = \frac{1}{\sqrt{2}} (V_\Phi + \phi) e^{ia/V_\Phi}$. Blue arrows point from labels to parts of the equation: 'flavon' points to Φ , ' $V_\Phi \sim f_a \gg v$ ' points to V_Φ , 'decouples' points to ϕ , and 'axion' points to a .

$$\Phi = \frac{1}{\sqrt{2}} (V_\Phi + \phi) e^{ia/V_\Phi}$$

flavon $V_\Phi \sim f_a \gg v$ decouples axion

Simplest realization: $PQ = U(1)_{FN}$

Flavor Symmetries

SM fields charged under flavor symmetry G ,
which is spontaneously broken by “**flavon**” field Φ

Effective Yukawa Lagrangian needs flavon
insertions in order to be invariant under G

The diagram shows the effective Yukawa Lagrangian $\mathcal{L}_{\text{eff}} \sim a_{ij} \left(\frac{\Phi}{\Lambda_F} \right)^{x_{ij}} h \bar{q}_i u_j$ enclosed in a hand-drawn rectangular box. Three blue arrows point to specific parts of the equation: one from the text 'O(1) coefficients' to the coefficient a_{ij} , one from 'cutoff scale' to the denominator Λ_F of the fraction, and one from 'from G selection rules' to the exponent x_{ij} .

$$\mathcal{L}_{\text{eff}} \sim a_{ij} \left(\frac{\Phi}{\Lambda_F} \right)^{x_{ij}} h \bar{q}_i u_j$$

Yukawas given by powers of small order parameter $\epsilon \equiv \frac{\langle \Phi \rangle}{\Lambda_F}$

U(1) Flavor Symmetry

Simplest U(1) symmetry works: Froggatt-Nielsen

	ϕ	$\bar{q}_i$	u_i	d_i	h
U(1)	-1	q_i	u_i	d_i	0

$$y_{ij}^U = a_{ij}^U \epsilon^{q_i + u_j}$$

$$y_{ij}^D = a_{ij}^D \epsilon^{q_i + d_j}$$

Can easily reproduce all hierarchies, e.g.

$$u_i = (4, 2, 0)$$

$$q_i = (3, 2, 0)$$

$$d_i = (4, 3, 3)$$

$$y^U \sim \begin{pmatrix} \epsilon^7 & \epsilon^5 & \epsilon^3 \\ \epsilon^6 & \epsilon^4 & \epsilon^2 \\ \epsilon^4 & \epsilon^2 & 1 \end{pmatrix}$$

$$y^D \sim \begin{pmatrix} \epsilon^7 & \epsilon^6 & \epsilon^6 \\ \epsilon^6 & \epsilon^5 & \epsilon^5 \\ \epsilon^4 & \epsilon^3 & \epsilon^3 \end{pmatrix}$$

$$\epsilon \approx 0.2$$

Get unitary rotations as $V_f \sim \epsilon^{|f_i - f_j|}$

Axion-Photon Couplings

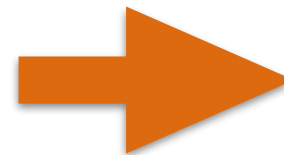
Although have considerable freedom in fermion
U(1) charges **can sharply predict E/N**

$$\frac{E}{N} \in [2.4, 3.0]$$

as direct consequence of fermion mass hierarchies

$$\underbrace{\det m_u \det m_d / v^6}_{\approx 10^{-20}} = \underbrace{[\det a_u \det a_d]}_{\mathcal{O}(1)} \epsilon^{2N}$$

$$\underbrace{\det m_d / \det m_e}_{\approx 0.7} = \underbrace{[\det a_d / \det a_e]}_{\mathcal{O}(1)} \epsilon^{\frac{8}{3}N - E}$$



$$\frac{E}{N} = \frac{8}{3} - 2\delta$$

$$\delta = \frac{\log \frac{\det m_d}{\det m_e} - \log \alpha_{de}}{\log \frac{\det m_u \det m_d}{v^6} - \log \alpha_{ud}}$$

≈ -0.4 ≈ 0
 ≈ -44 ≈ 0

Axion-Fermion Couplings

Axion-fermion couplings determined by fermion masses and mixings up to $\mathcal{O}(1)$ coefficients

$$X_f = \text{diag}(f_1, f_2, f_3) \rightarrow V_f^\dagger X_f V_f \quad V_f \sim \epsilon^{|f_i - f_j|}$$

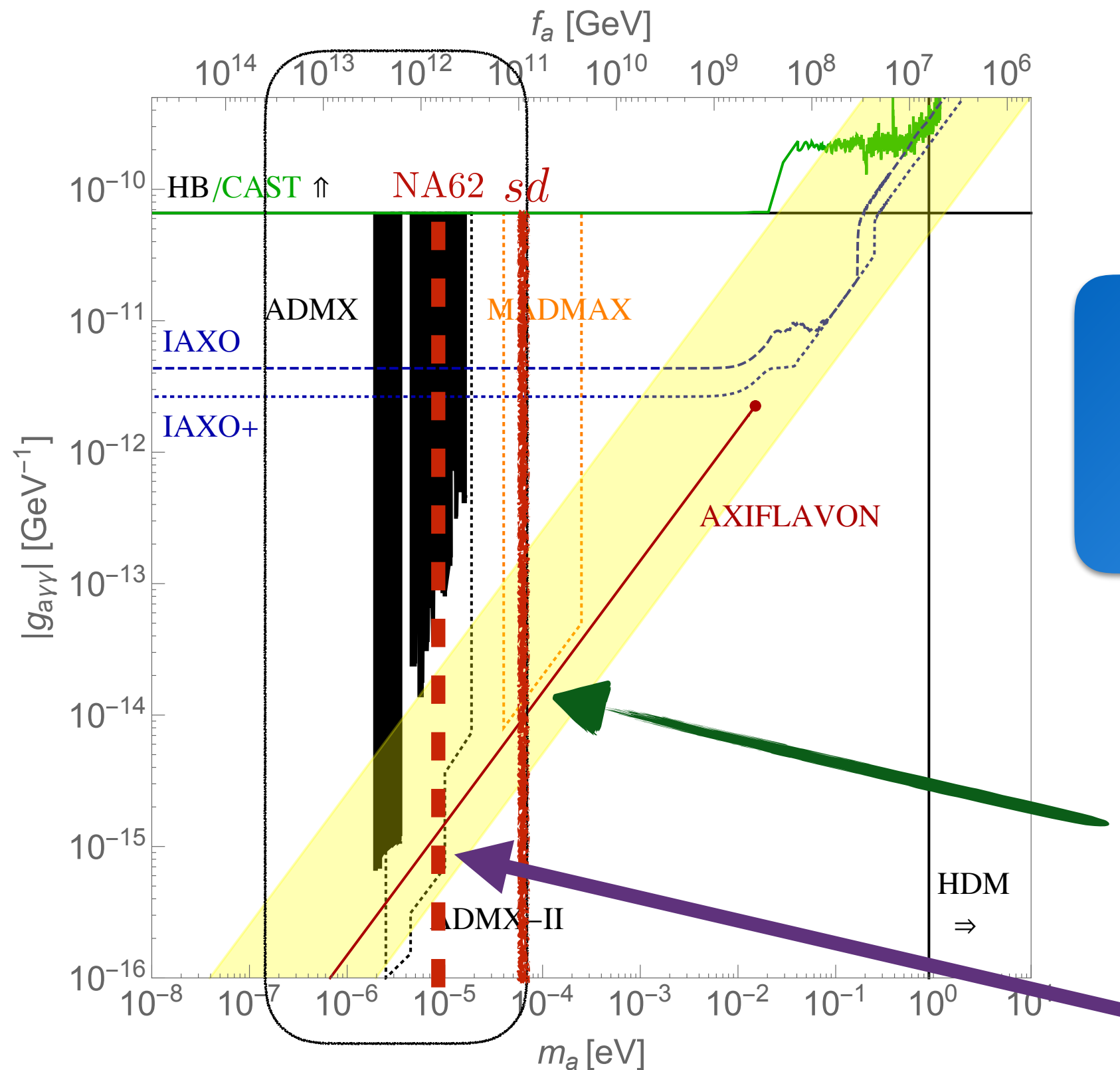
Flavor-violating couplings governed by CKM

$$C_{ij}^q \sim V_{ij}^{\text{CKM}} \quad C_{i < j}^u \sim \frac{m_i}{m_j V_{ij}^{\text{CKM}}}$$

Get large a-s-d couplings of order Cabibbo angle

$$C_{sd}^V \sim \lambda \approx 0.2$$

The Axiflavoron



**Natural axion DM
window testable at
NA62 (and ADMX-II)**

Present bound
from E787+E949

Expected future
bound from NA62

Summary

- Flavored Axion models allow for complementary axion searches with precision flavor experiments
- Generalized 2+1 DFSZ models with free flavor alignment parameters can realize nucleophobic & electrophobic axions up to 0.2 eV
- Well-motivated axion model from $PQ = FN$, giving flavor alignment parameters in terms of fermion masses and mixings