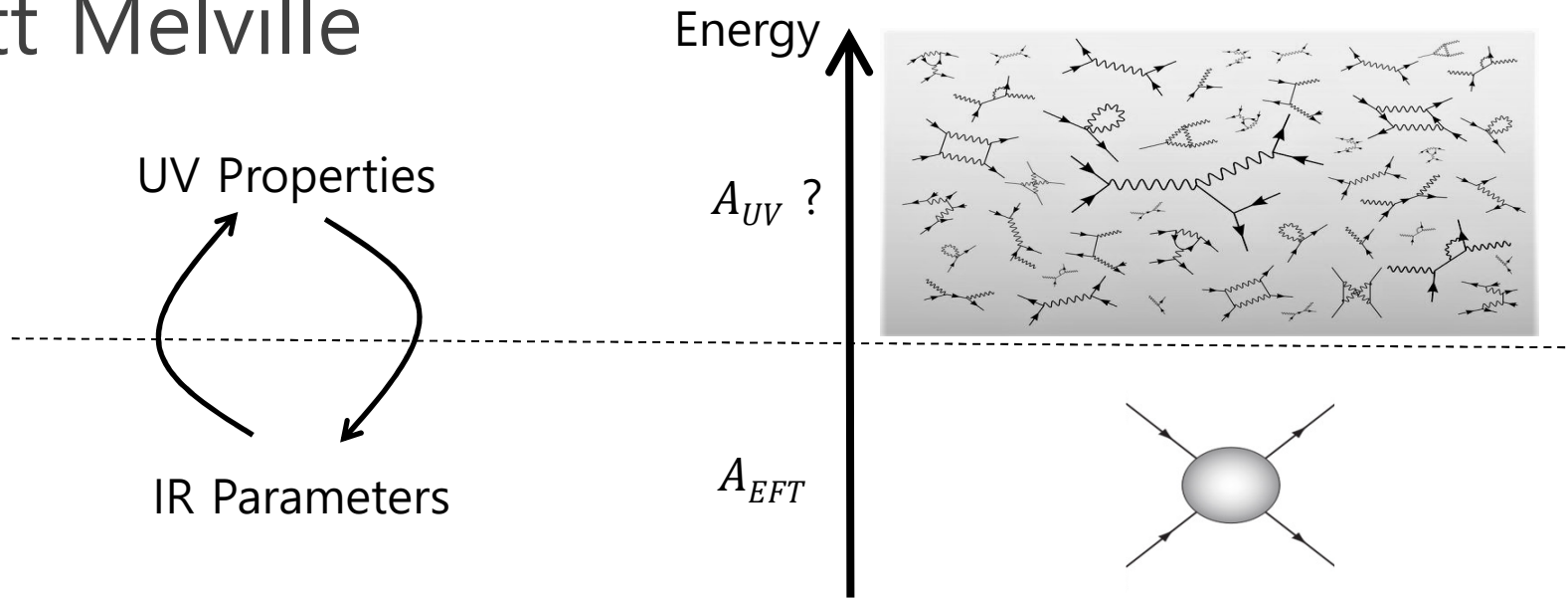


Low Energy? Think Positive!

Scott Melville



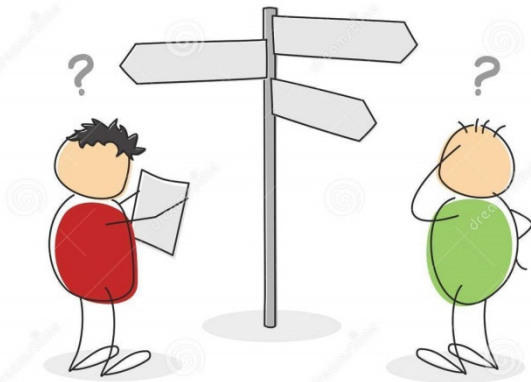
- Big Picture
- Positivity Constraints
- SMEFT
- Massive Gravity

Only some EFTs have sensible UV completions

... and this gives us constraining power!

bounds on dimension-8 operators for VBS

requires enhanced cutoff and weak coupling



Deriving bounds:

hep-th/0602178 Adams et al

1702.06134 SM et al

1706.02712 SM et al

Applying bounds:

1608.02942 Cheung et al

1710.02539 Bellazzini

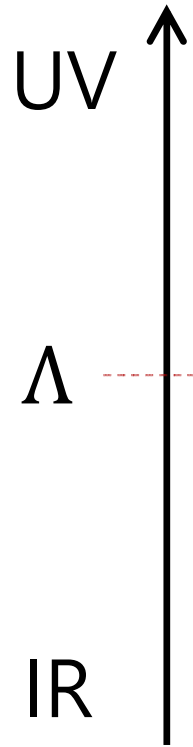
1710.09611 SM et al

1804.10624 SM et al

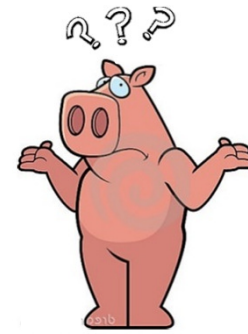
1808.00010 Zhang et al

Big Picture





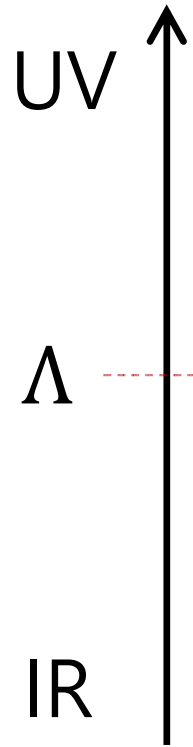
$$\mathcal{L}_{UV} = ?$$



$$\mathcal{L}_{EFT} = \sum^n \frac{c_n}{\Lambda^n} \mathcal{O}_n$$



Finite Number of Local Operators



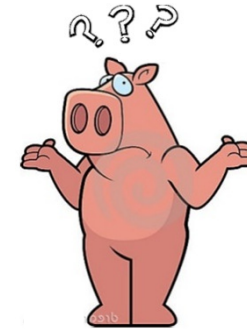
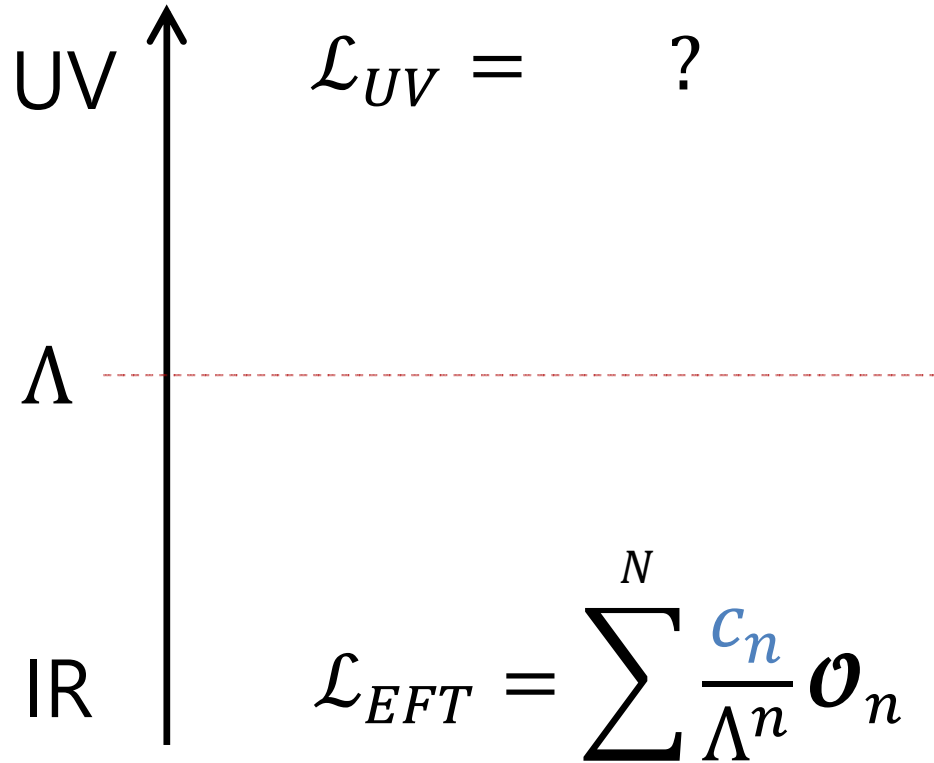
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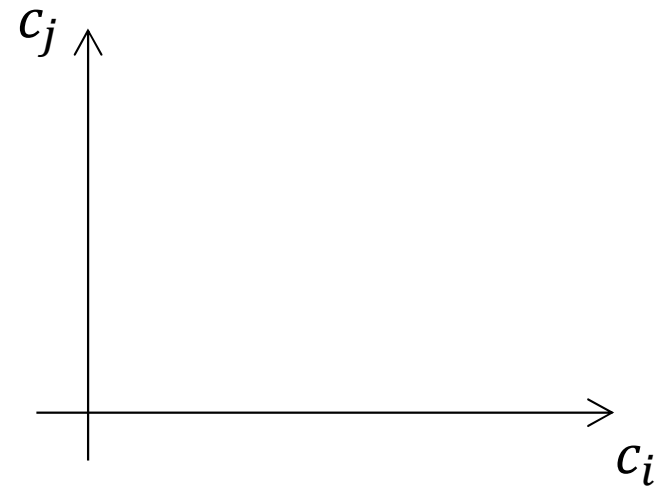
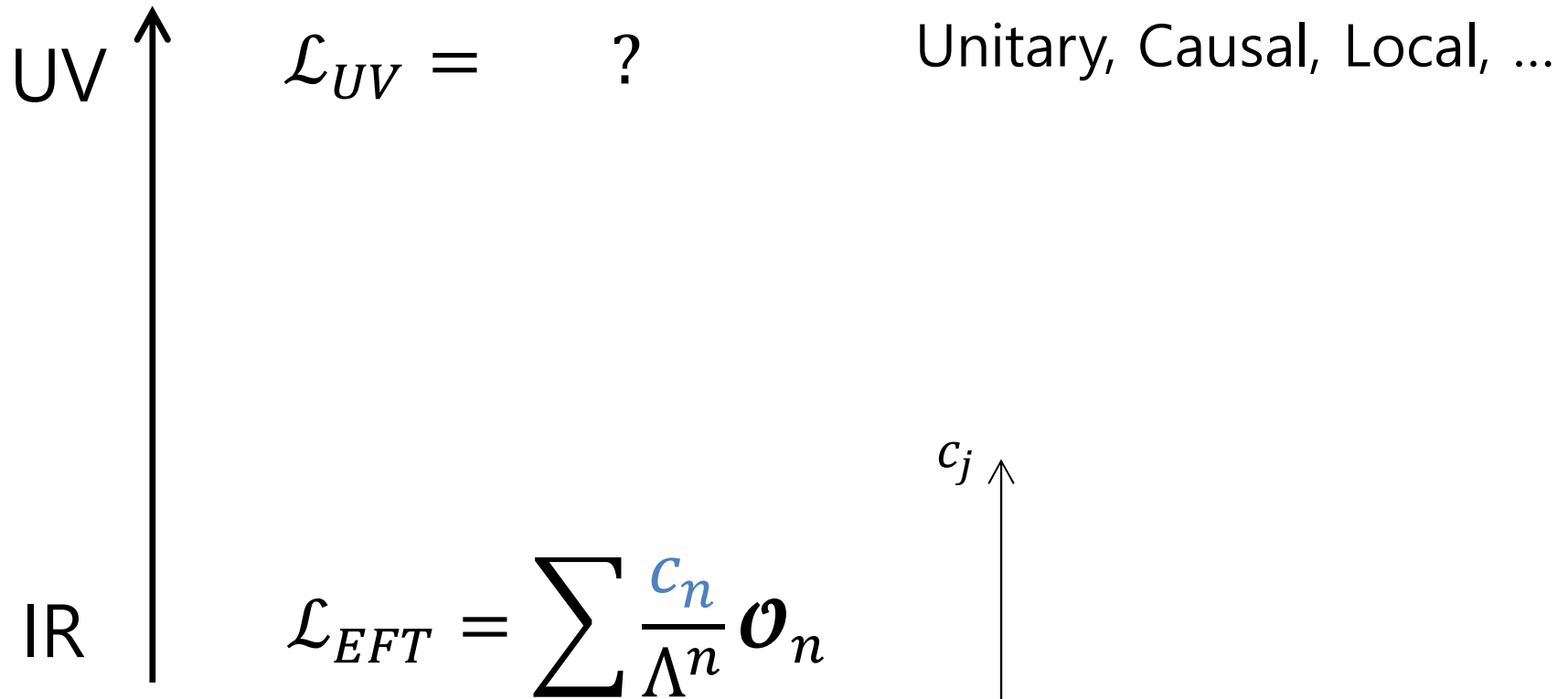


Finite Number of
Local Operators
with undetermined
coefficients

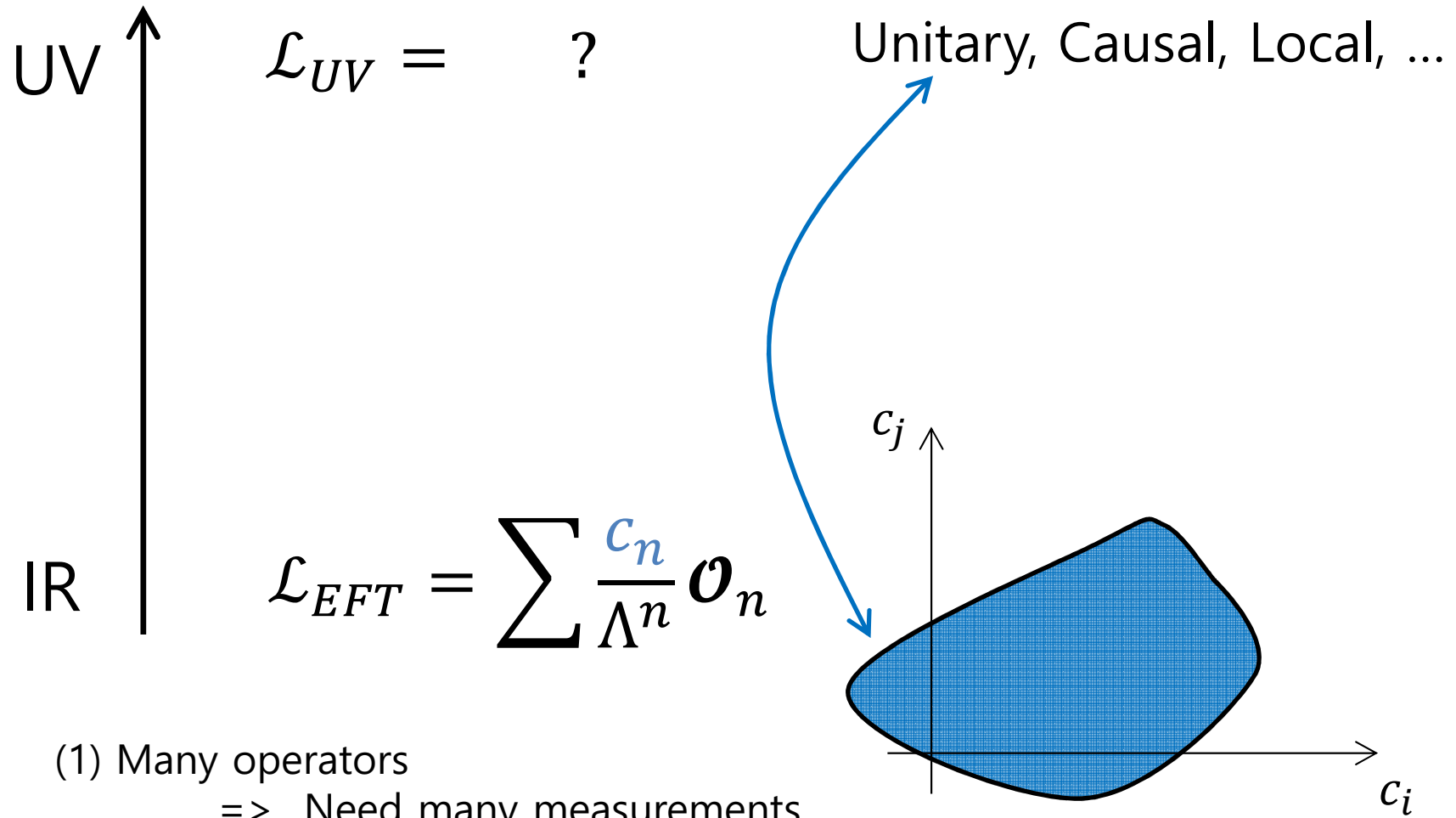


- (1) Many operators
=> Need many measurements
- (2) No deeper understanding of UV

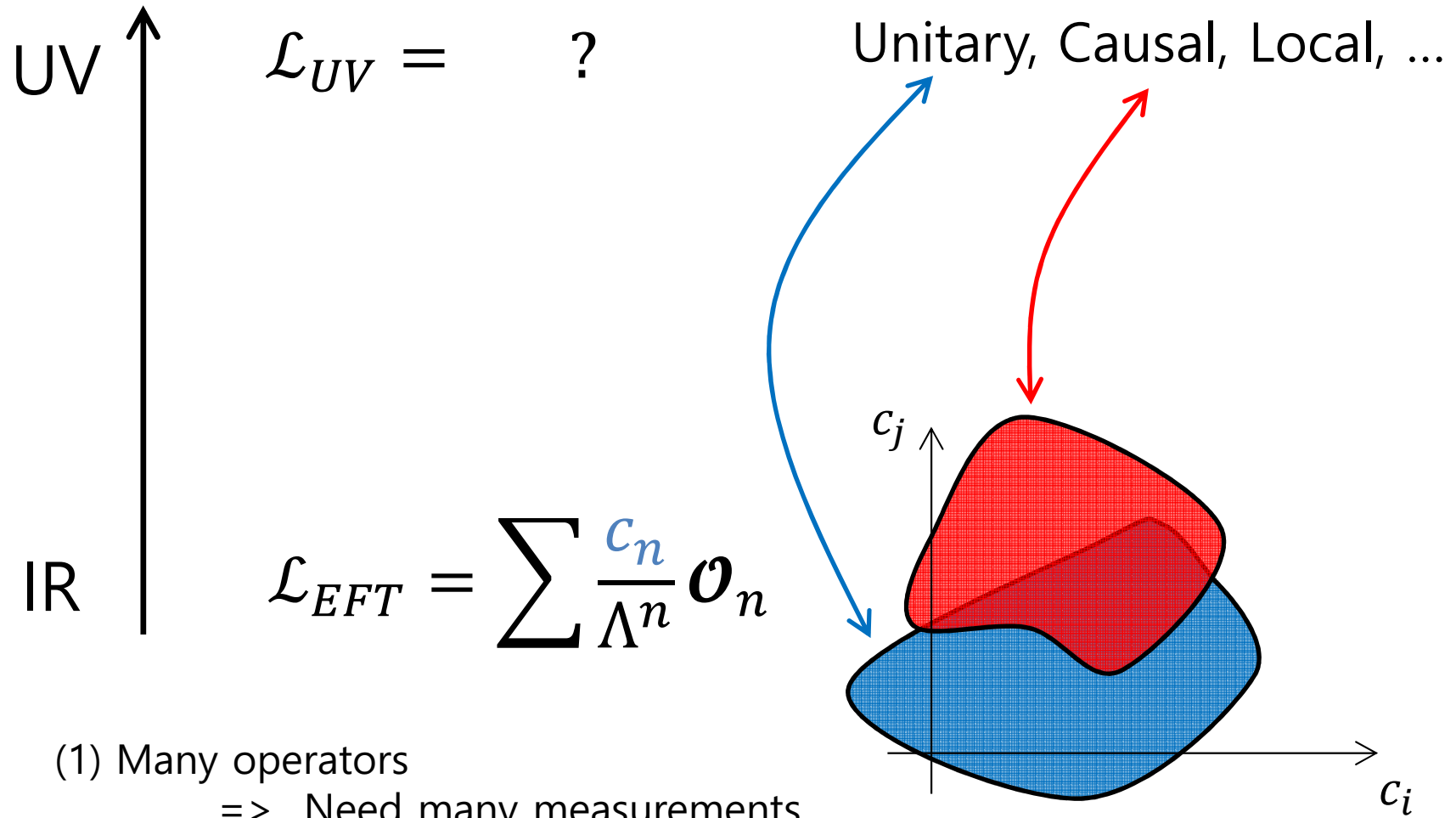
Finite Number of
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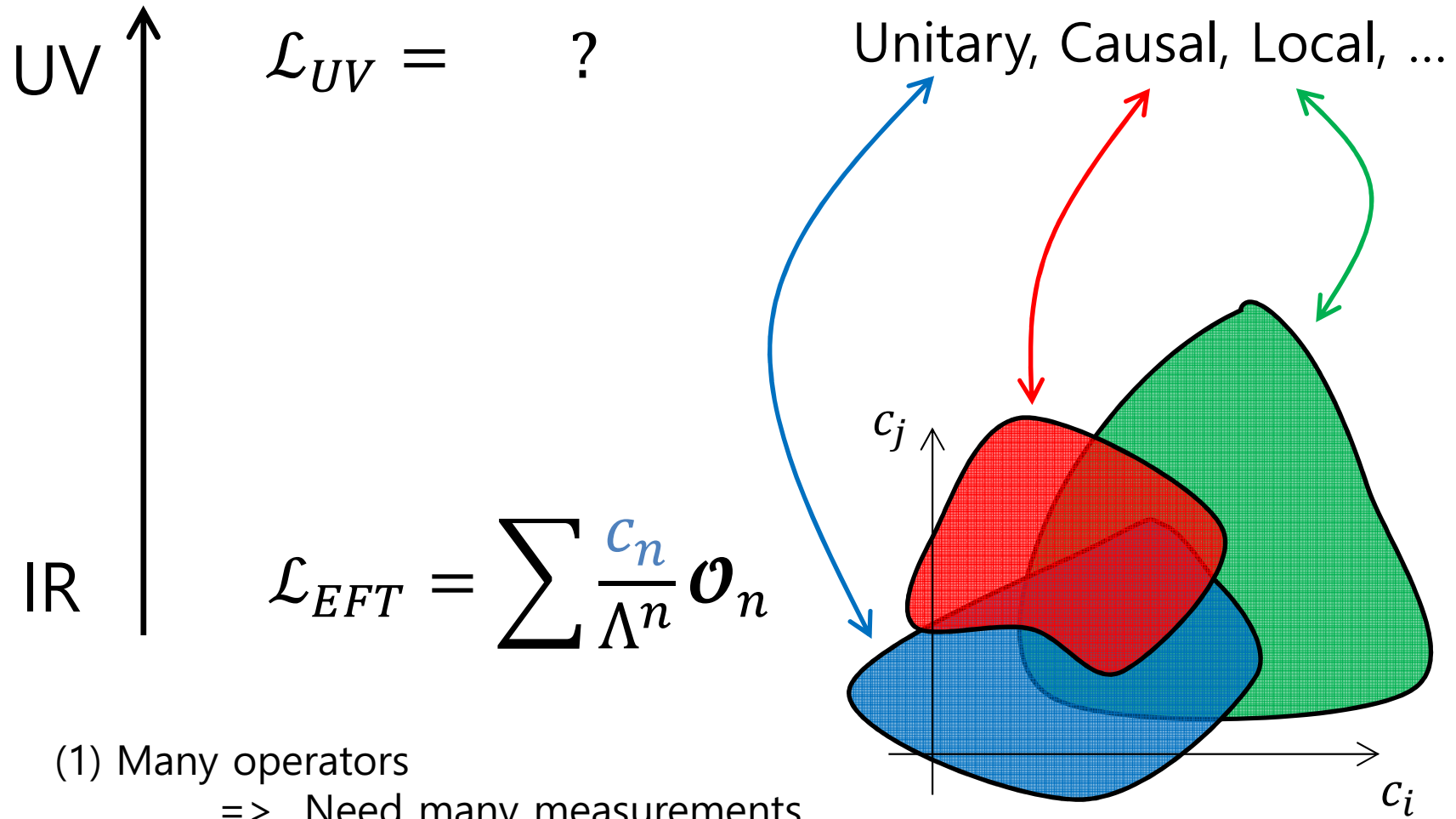


- (1) Many operators
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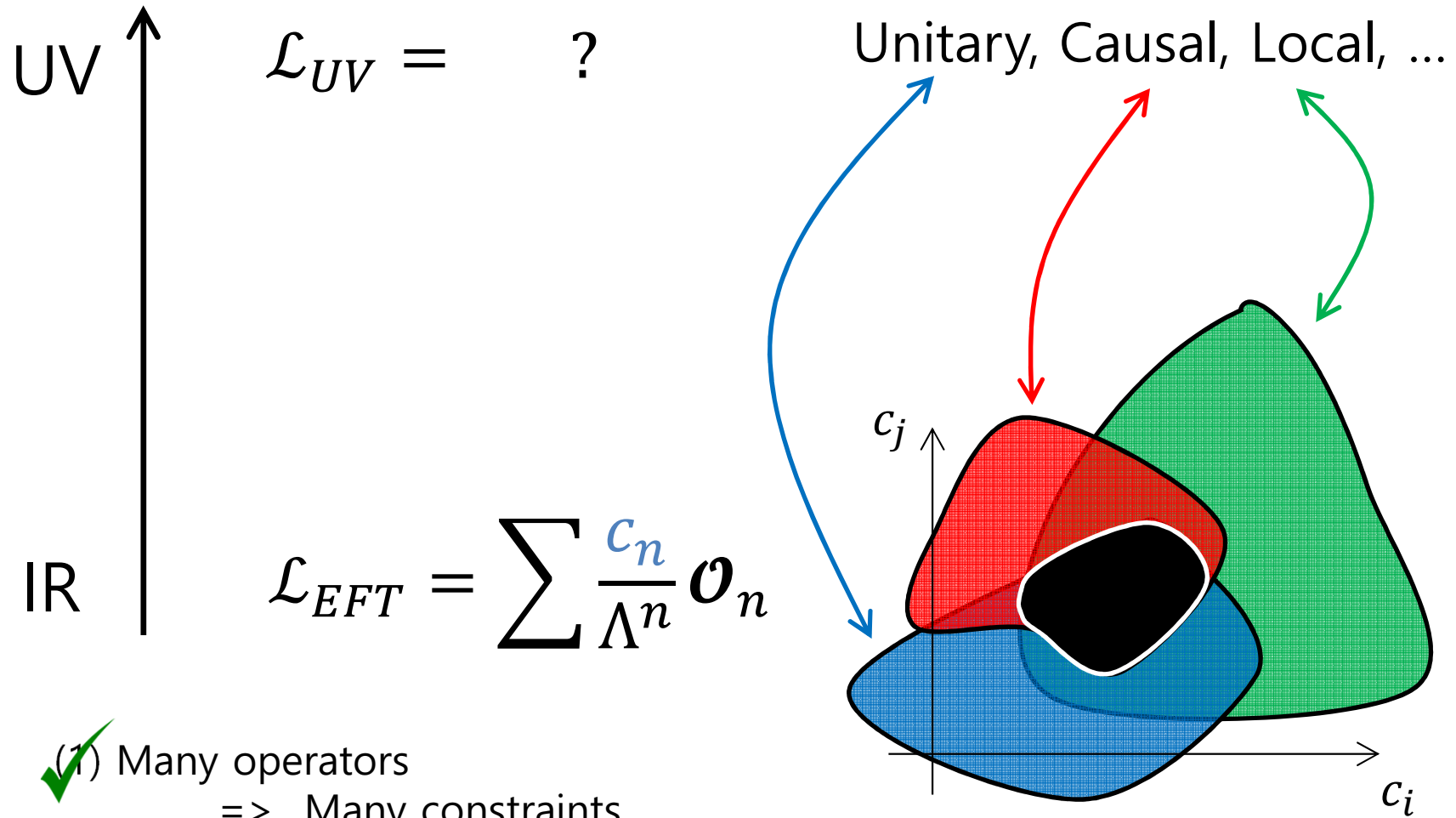


(1) Many operators
=> Need many measurements

(2) No deeper understanding of UV



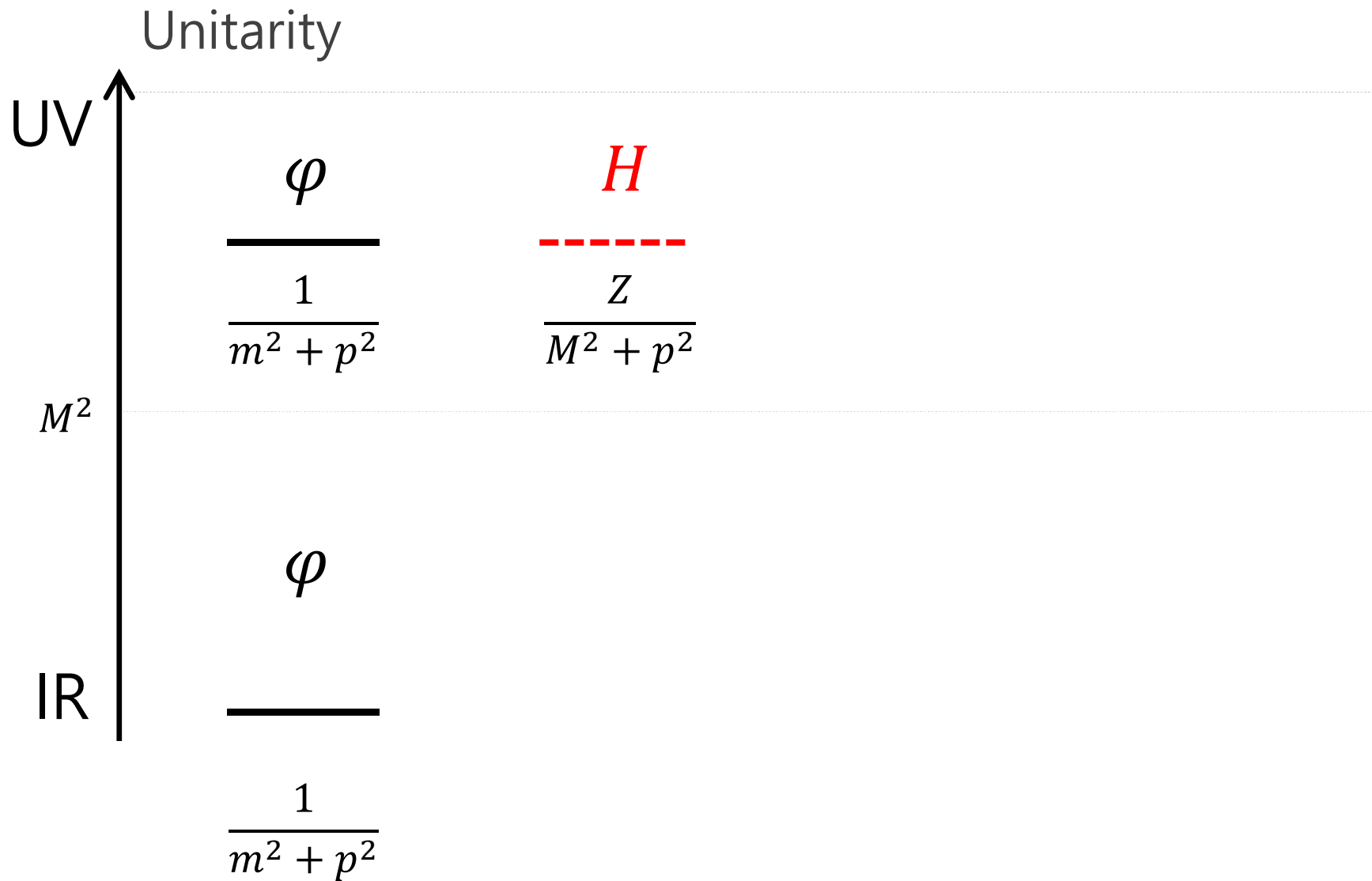
- (1) Many operators
=> Need many measurements
- (2) No deeper understanding of UV

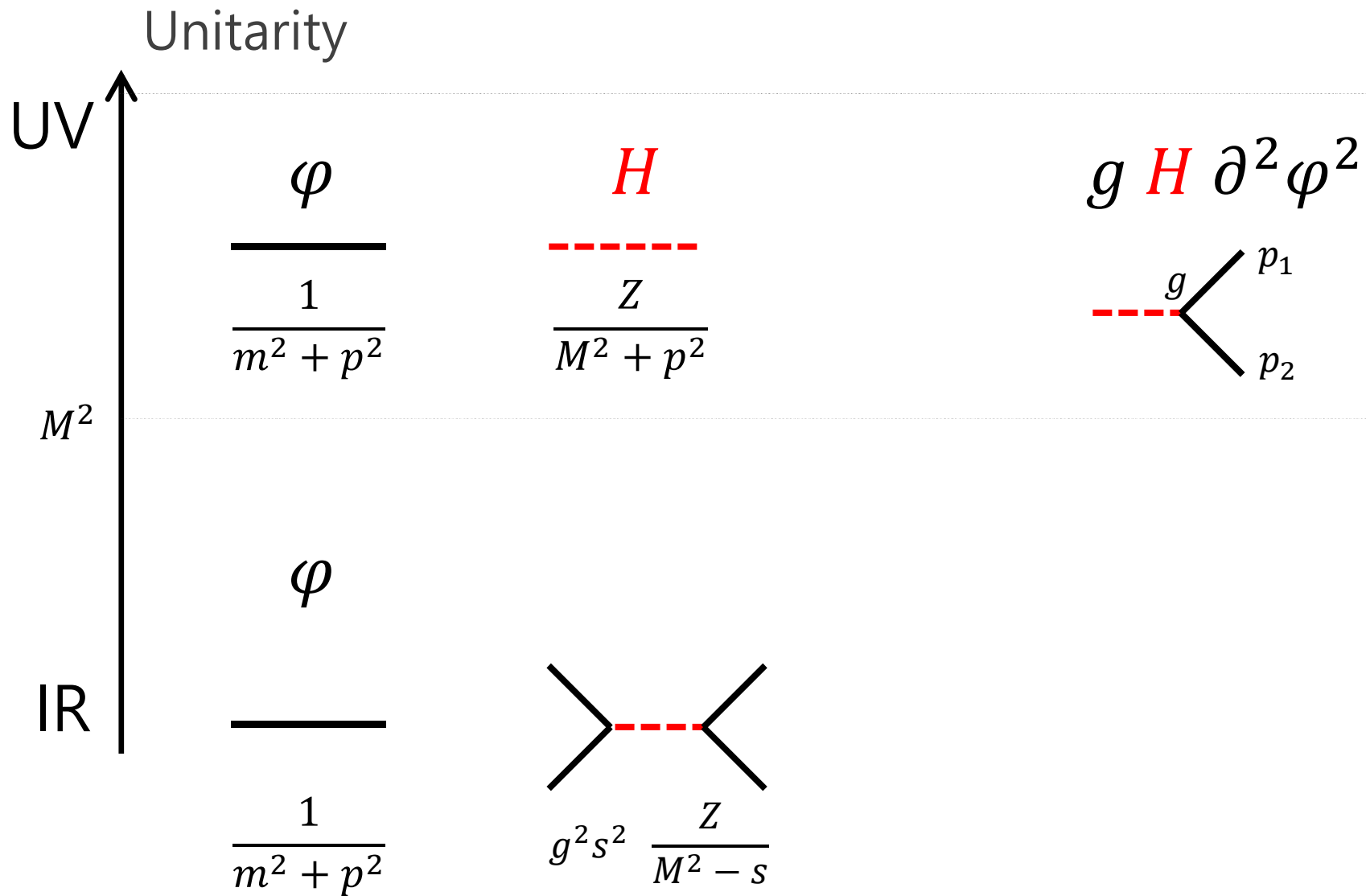


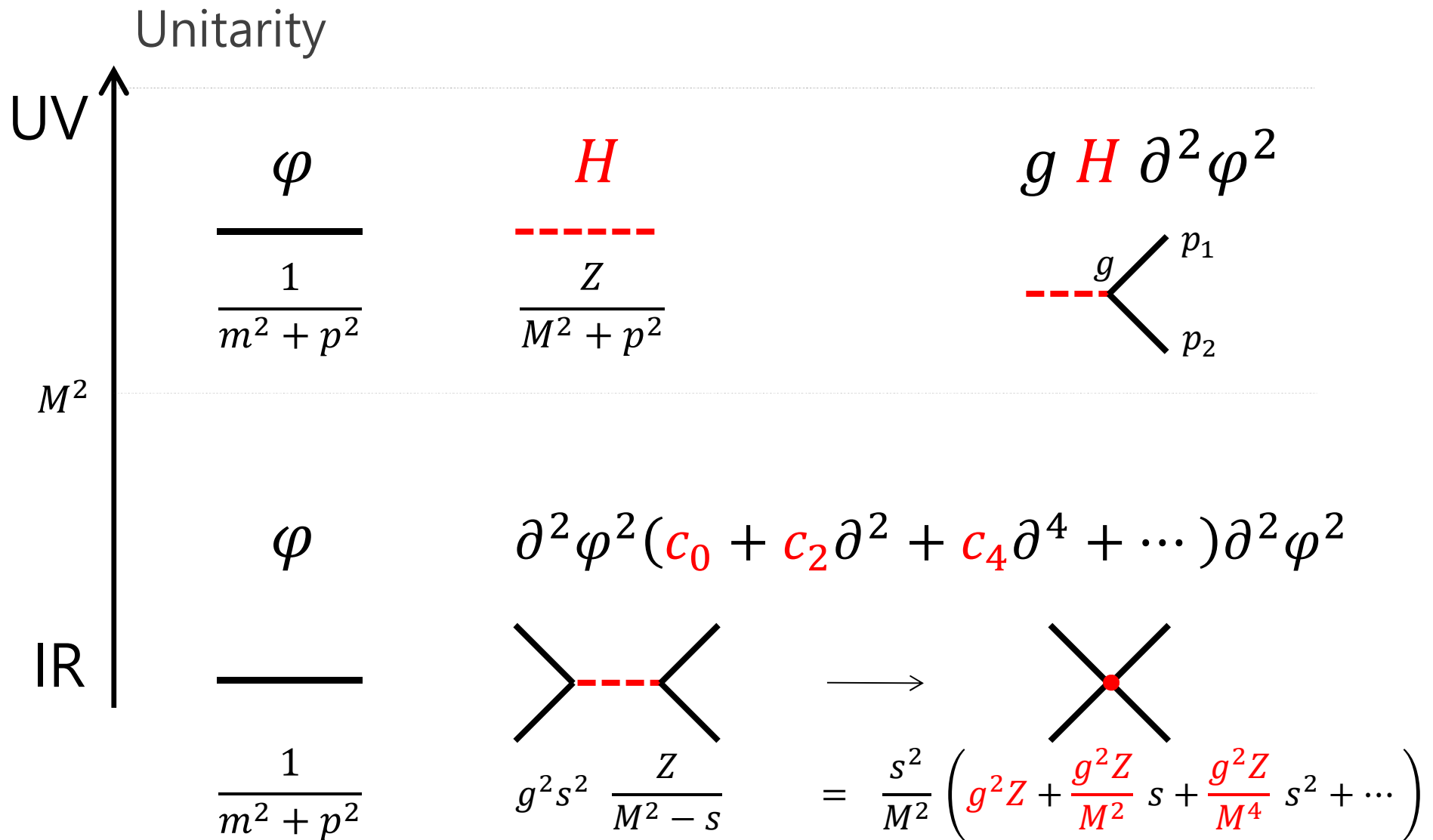
- ✓ (1) Many operators
=> Many constraints
- ✓ (2) Can infer UV properties from IR measurements

Positivity Constraints

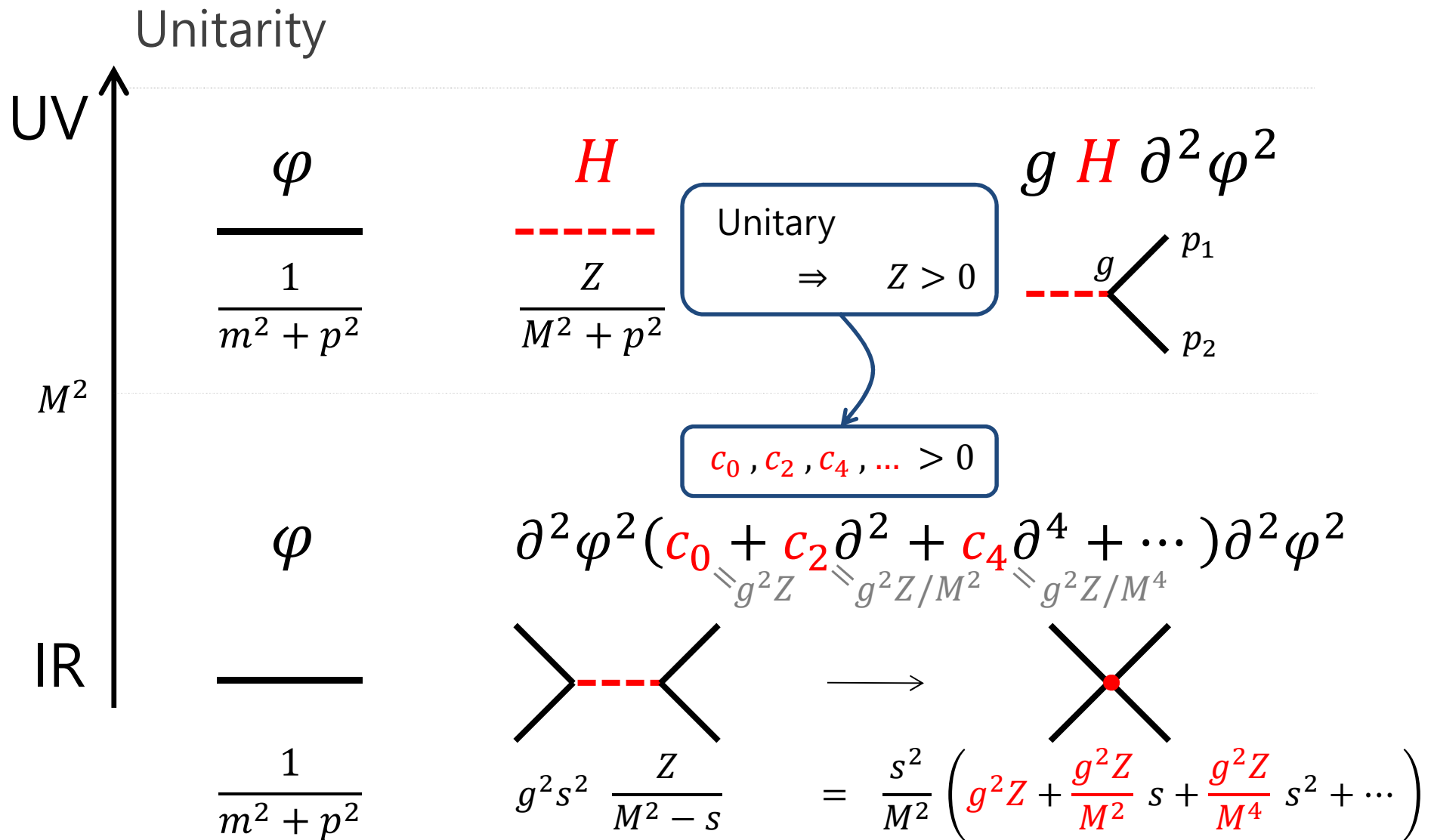






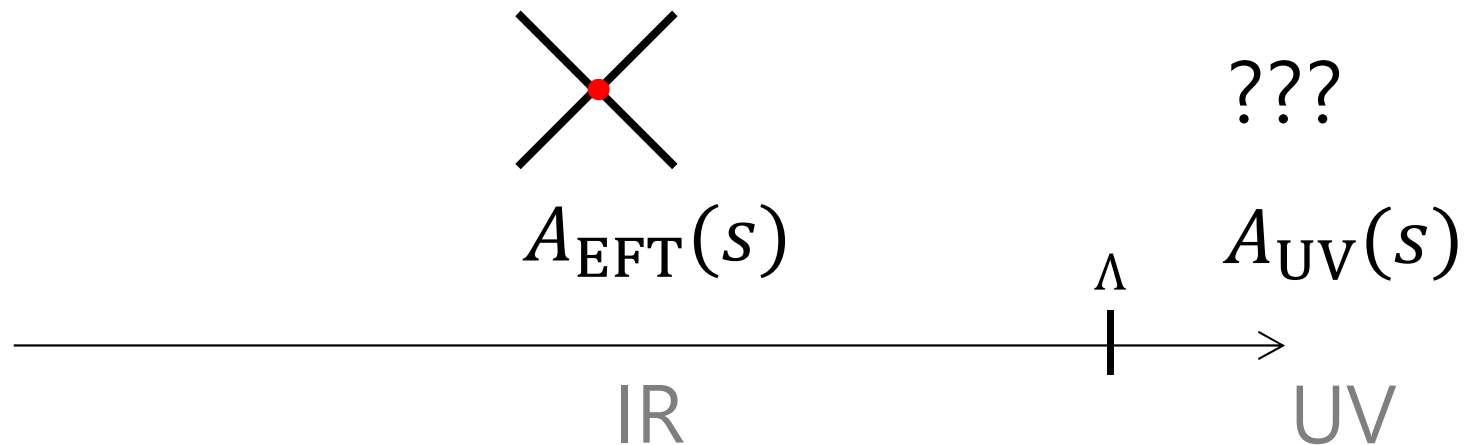


Positivity Constraints



Positivity Constraints

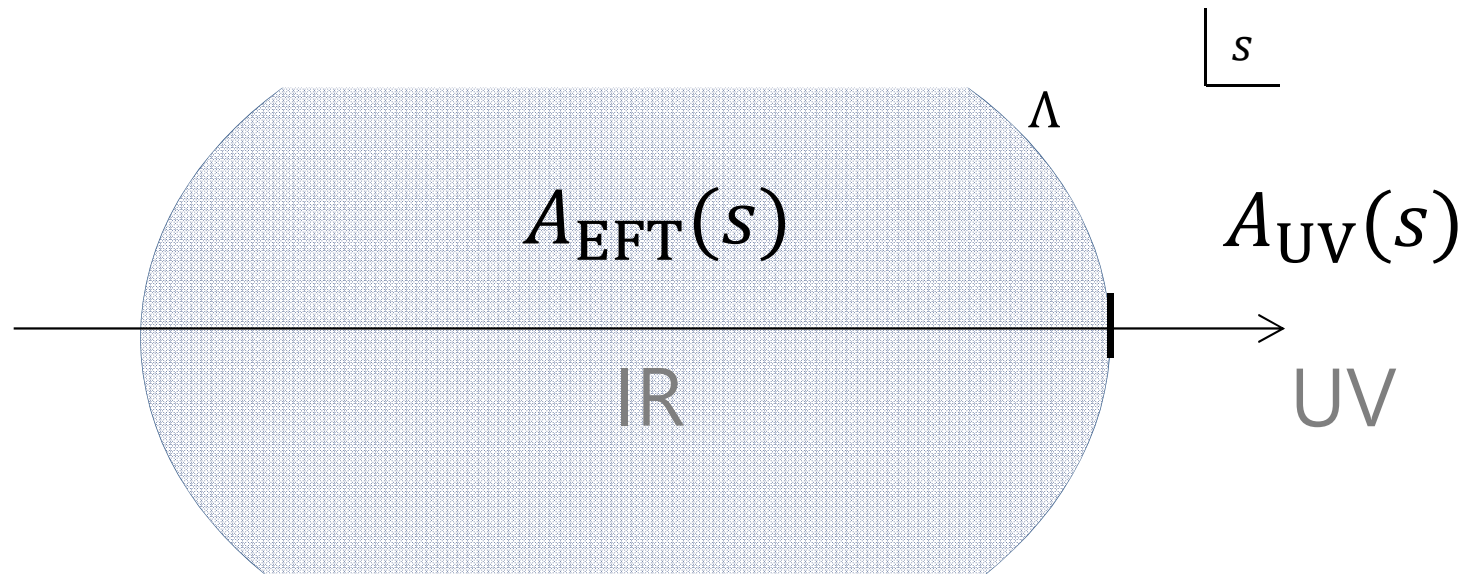
Analyticity



Causality

$\Rightarrow A(s)$ is analytic (up to known poles & branch cuts)

Analyticity

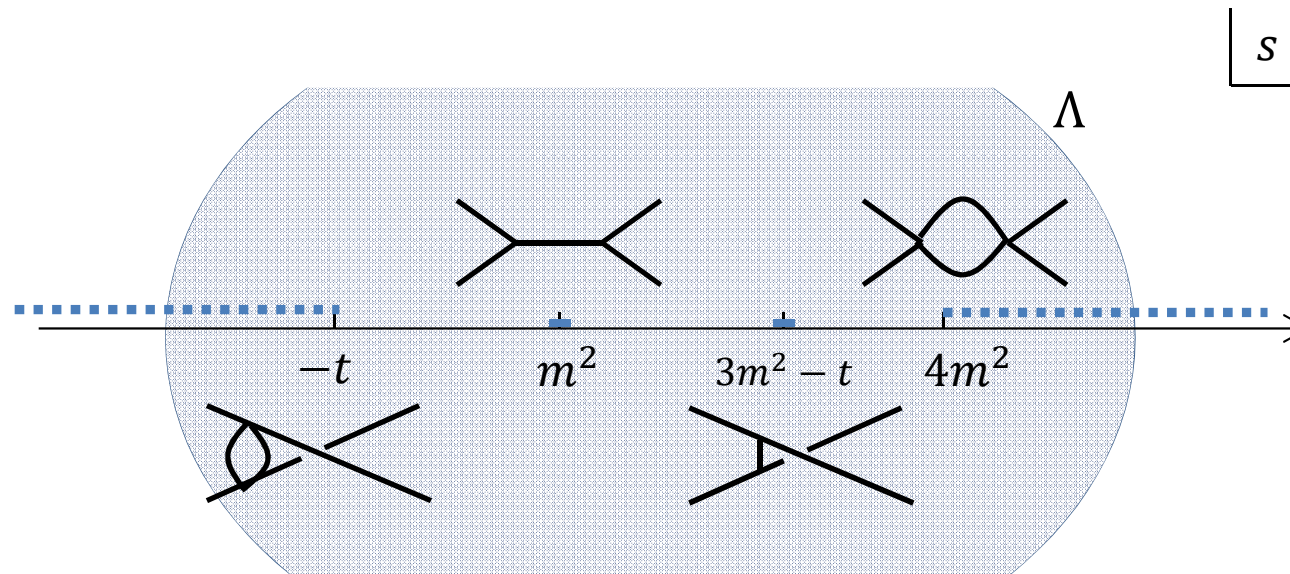


Causality

$\Rightarrow A(s)$ is analytic (up to known poles & branch cuts)

At fixed t , expect singularities from s - and u - channel exchange

Analyticity

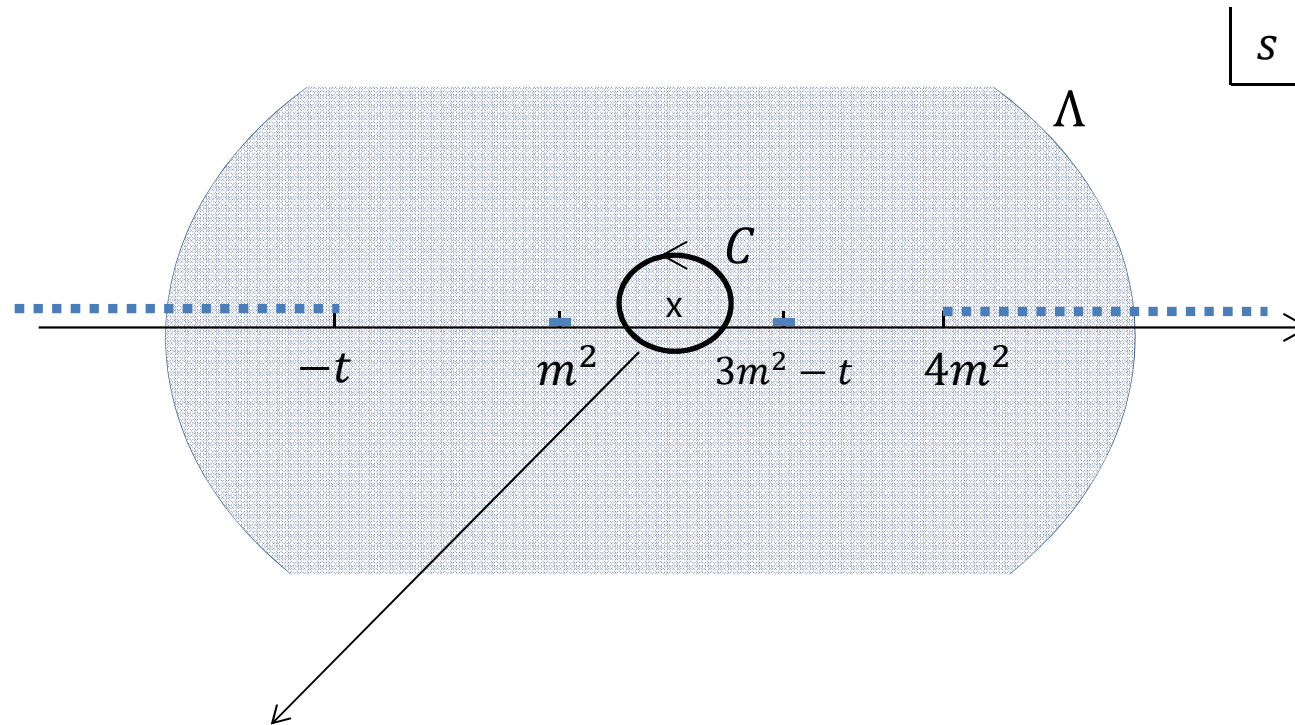


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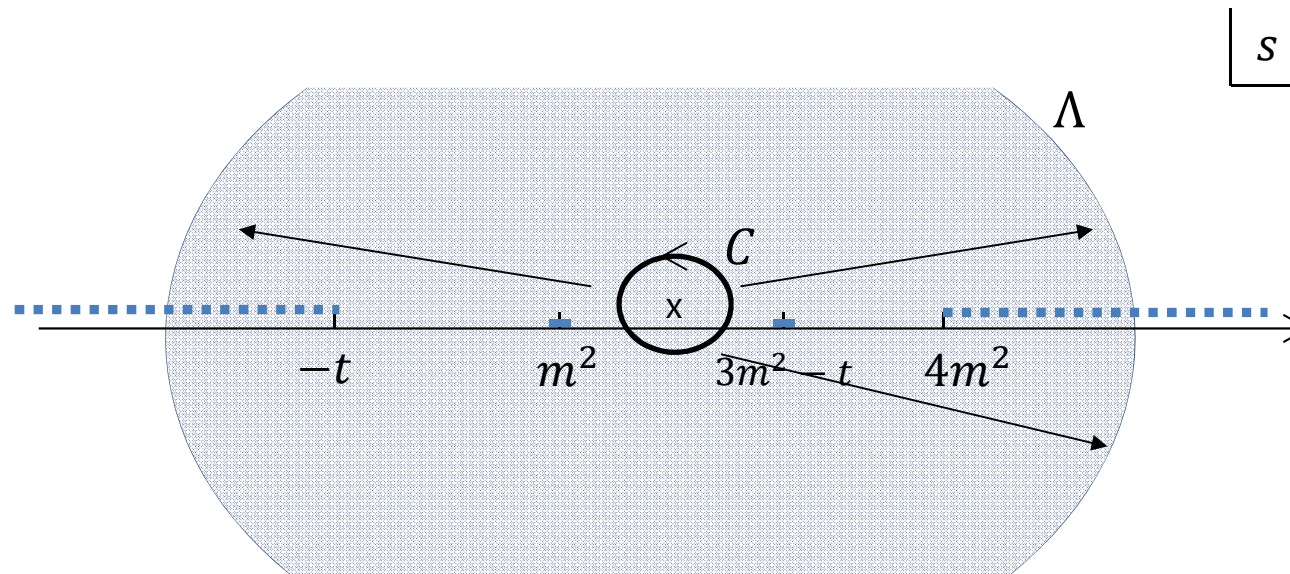
At fixed t , expect singularities from s - and u - channel exchange

Analyticity



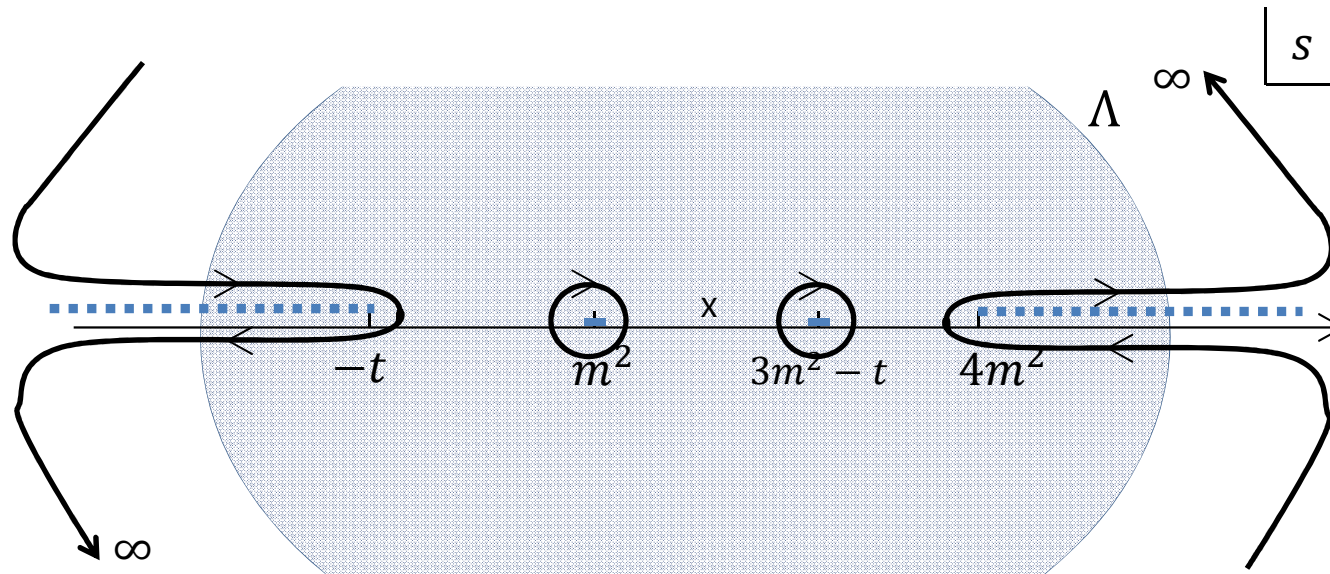
$$\partial_s^N A_{EFT}(s, t) = \oint_C \frac{d\mu}{2\pi i} \frac{A(\mu, t)}{(\mu - s)^{N+1}}$$

Analyticity



$$\partial_s^N A_{EFT}(s, t) = \oint_C \frac{d\mu}{2\pi i} \frac{A(\mu, t)}{(\mu - s)^{N+1}}$$

Analyticity



$$\partial_s^N A_{EFT}(s, t) = \text{Poles} + \left(\int_{-\infty}^{-t} + \int_{4m^2}^{\infty} \right) \frac{d\mu}{\pi} \frac{\text{Im } A(\mu, t)}{(\mu - s)^{N+1}} + \oint_{|\mu|=\infty} \frac{d\mu}{2\pi i} \frac{A(\mu, t)}{\mu^{N+1}}$$

Positivity

Analyticity
(Causality)

$$\partial_s^N \tilde{A}_{\text{EFT}}(s, t) = \left(\int_{-\infty}^{-t} + \int_{4m^2}^{\infty} \right) \frac{d\mu}{\pi} \frac{\text{Im } A(\mu, t)}{(\mu - s)^{N+1}} + \oint_{|\mu|=\infty} \frac{d\mu}{2\pi i} \frac{A(\mu, t)}{\mu^{N+1}}$$

Unitarity

$$\text{Im } A(\mu, t) \Big|_{t \rightarrow 0} = \mu \sigma(\mu) > 0$$

Positivity

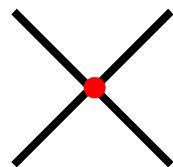
Analyticity
(Causality)

$$\partial_s^N \tilde{A}_{\text{EFT}}(s, t) = \left(\int_{-\infty}^{-t} + \int_{4m^2}^{\infty} \right) \frac{d\mu}{\pi} \frac{\text{Im } A(\mu, t)}{(\mu - s)^{N+1}} + \oint_{|\mu|=\infty} \frac{d\mu}{2\pi i} \frac{A(\mu, t)}{\mu^{N+1}}$$

Unitarity

$$\text{Im } A(\mu, t) \Big|_{t \rightarrow 0} = \mu \sigma(\mu) > 0$$

Analyticity connects EFT to the UV



$$A_{\text{EFT}}(s) = s^2 (c_0 + c_2 s + c_4 s^2 + \dots)$$

$$\text{UV Theory} \Rightarrow \underbrace{g^2 Z}_{\text{}} \quad \underbrace{g^2 Z / M^2}_{\text{}} \quad \underbrace{g^2 Z / M^4}_{\text{}}$$

$$\text{Unitarity} \Rightarrow Z > 0$$

Positivity

Analyticity
(Causality)

$$\partial_s^N \tilde{A}_{\text{EFT}}(s, t) = \left(\int_{-\infty}^{-t} + \int_{4m^2}^{\infty} \right) \frac{d\mu}{\pi} \frac{\text{Im } A(\mu, t)}{(\mu - s)^{N+1}} + \oint_{|\mu|=\infty} \frac{d\mu}{2\pi i} \frac{A(\mu, t)}{\mu^{N+1}}$$

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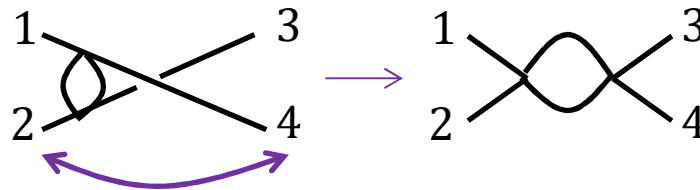
Analyticity
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Crossing



$$\int_{4m^2}^{\infty} d\mu C_N(s; \mu, t) \text{Im } A(\mu, t)$$

Positivity

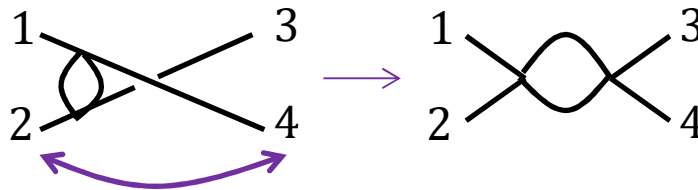
Analyticity
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$$\partial_s^N \tilde{A}_{\text{EFT}}(s, t) = \left(\int_{-\infty}^{-t} + \int_{4m^2}^{\infty} \right) \frac{d\mu}{\pi} \frac{\text{Im } A(\mu, t)}{(\mu - s)^{N+1}} + \oint_{|\mu|=\infty} \frac{d\mu}{2\pi i} \frac{A(\mu, t)}{\mu^{N+1}}$$

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Crossing



$$\int_{4m^2}^{\infty} d\mu C_N(s; \mu, t) \text{Im } A(\mu, t)$$

(Locality + Gap)
Polynomially Bounded

$$\lim_{\mu \rightarrow \infty} |A(\mu, t)| < \mu^2$$

$$= 0 \text{ if } N \geq 2$$

Positivity

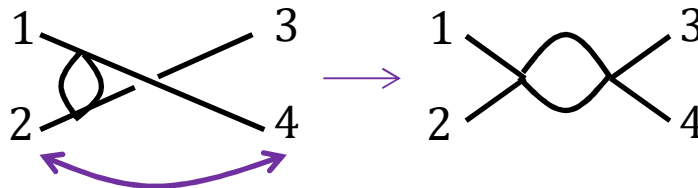
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Unitarity

$$\text{Im } A(\mu, t) \Big|_{t \rightarrow 0} = \mu \sigma(\mu) > 0$$

Crossing



$$\int_{4m^2}^{\infty} d\mu C_N(s; \mu, t) \text{Im } A(\mu, t)$$

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$$\partial_s^N \tilde{A}_{\text{EFT}}(s, 0) = \int d\mu C_N(s; \mu) \sigma(\mu) > 0$$

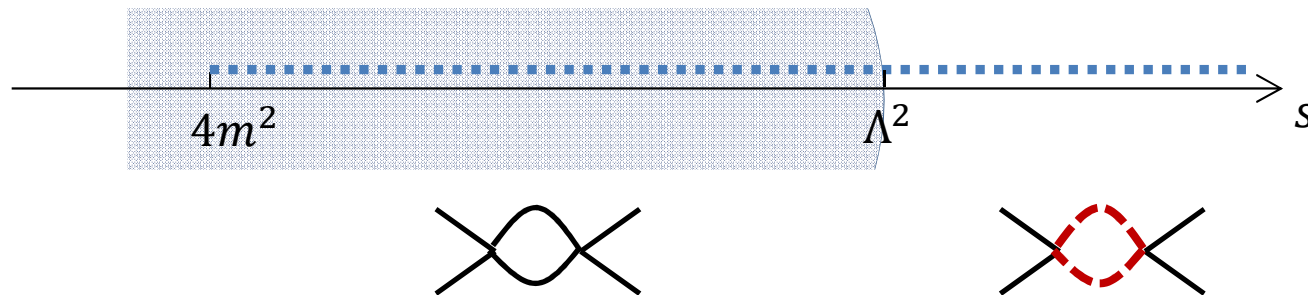
Positivity Constraints

Improving Positivity

$$\partial_s^N \tilde{A}_{\text{EFT}}(s) = \int d\mu \ C_N(s; \mu) \sigma(\mu) > 0$$

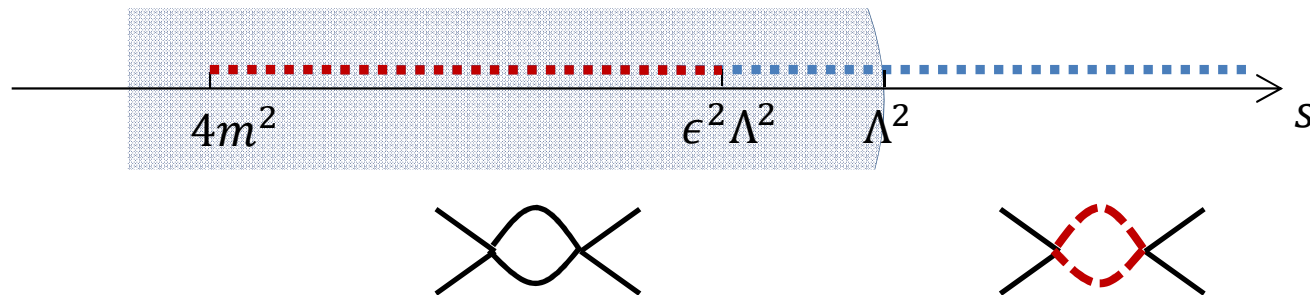
Improving Positivity

$$\partial_s^N \tilde{A}_{\text{EFT}}(s) = \int d\mu \, C_N(s; \mu) \sigma(\mu) > 0$$



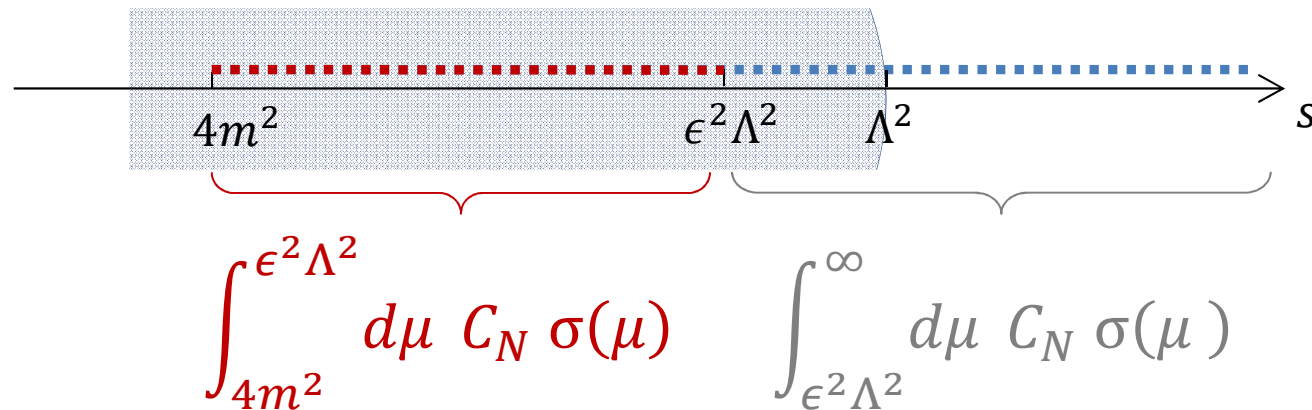
Improving Positivity

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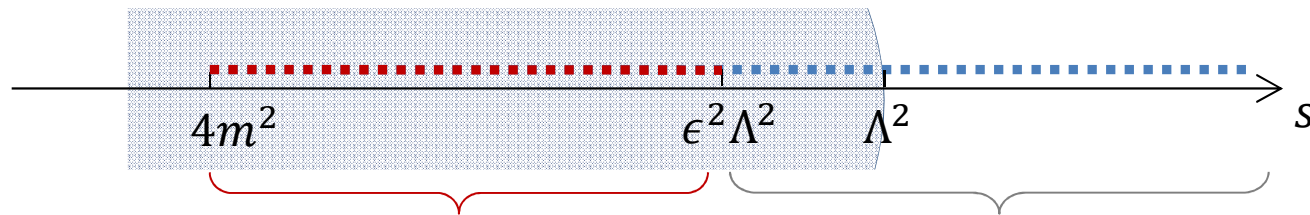
Improving Positivity

$$\partial_s^N \tilde{A}_{\text{EFT}}(s) = \int d\mu \, C_N(s; \mu) \sigma(\mu) > 0$$



Improving Positivity

$$\partial_S^N \tilde{A}_{\text{EFT}}(s) = \int d\mu C_N(s; \mu) \sigma(\mu) > 0$$



$$\partial_S^N \tilde{A}_{\text{EFT}}(s) - \int_{4m^2}^{\epsilon^2 \Lambda^2} d\mu C_N \sigma(\mu) = \int_{\epsilon^2 \Lambda^2}^{\infty} d\mu C_N \sigma(\mu) > 0$$

$$\partial_S^N \tilde{A}_{\text{EFT}}(s) > \int_{4m^2}^{\epsilon^2 \Lambda^2} d\mu C_N \sigma(\mu)$$

Improving Positivity

$$\partial_S^N \tilde{A}_{\text{EFT}}(s) > \int_{4m^2}^{\epsilon^2 \Lambda^2} d\mu C_N \sigma(\mu)$$

Improving Positivity

$$\partial_S^N \tilde{A}_{\text{EFT}}(s) > \int_{4m^2}^{\epsilon^2 \Lambda^2} d\mu C_N \sigma(\mu)$$



$$\mathcal{L}_{\text{EFT}} \left[\frac{\Phi}{M}, \frac{\partial}{M} \right]$$



$$\frac{1}{g_*^2} \mathcal{L}_{\text{EFT}} \left[\frac{g_* \Phi}{M}, \frac{\partial}{M} \right]$$

Improving Positivity

$$\partial_S^N \tilde{A}_{\text{EFT}}(s) > \int_{4m^2}^{\epsilon^2 \Lambda^2} d\mu C_N \sigma(\mu)$$

~~$$g_*^2 c_4 (\partial\Phi)^4 / M^4$$~~

$$A_{\text{EFT}} = g_*^2 \frac{c_4 s^2}{M^4} + \dots$$

For scalars, $\pi C_2(s; \mu) = \frac{\mu}{(\mu - s)^3} + \frac{\mu}{(\mu - 4m^2 + s)^3} \sim \frac{1}{\mu^2}$

$$\sigma(\mu) \sim \frac{|A(\mu)|^2}{\mu} \sim \mu^3$$

$$\frac{g_*^2 c_4}{M^4} > \frac{g_*^4 c_4^2}{16\pi^2 M^8} \int^{\epsilon^2 M^2} d\mu 2\mu \Rightarrow$$

$$1 > \frac{g_*^2 c_4}{16\pi^2} (\epsilon^4 + \dots)$$

Positivity Constraints

The Scale of New Physics

Analyticity
(Causality)

$$\partial_S^N \tilde{A}_{\text{EFT}}(s, t) = \int_{\mu_{\text{min}}}^{\infty} d\mu C_N(s; \mu, t) \text{Im}A(\mu, t)$$

Unitarity

$$\text{Im} A(\mu, t) \Big|_{t \rightarrow 0} = \mu \sigma(\mu) > 0$$

The Scale of New Physics

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$$\partial_t^P \text{Im} A(\mu, t) > 0$$

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$$\partial_t \partial_S^N \tilde{A}_{\text{EFT}}(s, t) = \int_{\mu_{\text{min}}}^{\infty} d\mu C_N(s; \mu, t) \partial_t \text{Im}A(\mu, t) + \int_{\mu_{\text{min}}}^{\infty} d\mu \underbrace{\partial_t C_N(s; \mu, t)}_{\text{Not positive, but known}} \text{Im}A(\mu, t)$$

The Scale of New Physics

Analyticity
(Causality)

$$\partial_S^N \tilde{A}_{\text{EFT}}(s, t) = \int_{\mu_{\min}}^{\infty} d\mu C_N(s; \mu, t) \text{Im}A(\mu, t)$$

Unitarity

$$\partial_t^P \text{Im} A(\mu, t) > 0$$

$$\partial_t \partial_S^N \tilde{A}_{\text{EFT}}(s, t) + \frac{2N + 1}{\mu_{\min}} \partial_S^N \tilde{A}_{\text{EFT}}(s, t) > 0$$

The Scale of New Physics

Analyticity
(Causality)

$$\partial_S^N \tilde{A}_{\text{EFT}}(s, t) = \int_{\mu_{\min}}^{\infty} d\mu C_N(s; \mu, t) \text{Im}A(\mu, t)$$

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Can constrain *every* s and t derivative of A_{EFT} !

The Scale of New Physics

Analyticity
(Causality)

$$\partial_S^N \tilde{A}_{\text{EFT}}(s, t) = \int_{\mu_{\min}}^{\infty} d\mu C_N(s; \mu, t) \text{Im}A(\mu, t)$$

Unitarity

$$\partial_t^P \text{Im} A(\mu, t) > 0$$

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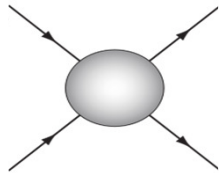
Can constrain *every* s and t derivative of A_{EFT} !

Can constrain μ_{\min} , the scale of new physics!

Positivity in Practice

$$\mathcal{L} = \mathcal{L}_{\text{renorm}} + \frac{f_6 \mathcal{O}_6}{\Lambda^2} + \frac{f_8 \mathcal{O}_8}{\Lambda^4} + \frac{f_{10} \mathcal{O}_{10}}{\Lambda^6} + \dots$$

$$A_{EFT} = c_0 + c_1 s + c_2 (f_6^2, f_8) s^2 + c_3 s^3 + \dots$$



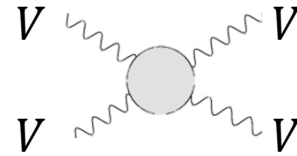
c_n Bounds
 g_* Coupling

c_n Bounds
 g_* Coupling
 M , Mass of New States

SMEFT

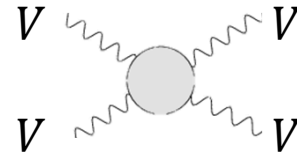


$$\mathcal{L}_{\text{EFT}} = \mathcal{L}_{\text{SM}} + \sum_i^{18} \frac{F_i}{\Lambda^4} \mathcal{O}_i$$

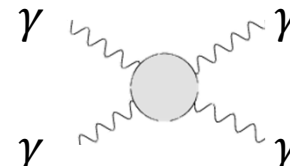
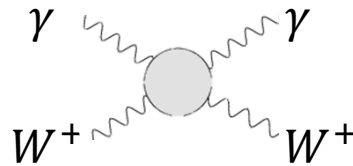
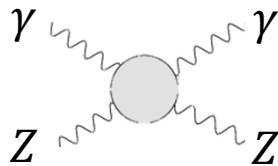
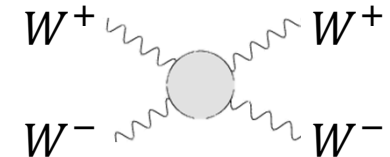
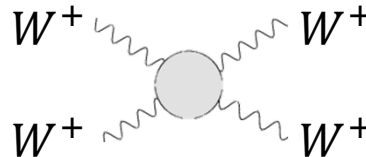
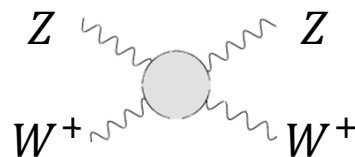
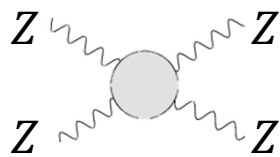


$$V = Z, W^\pm, \gamma$$

$$\mathcal{L}_{\text{EFT}} = \mathcal{L}_{\text{SM}} + \sum_i^{18} \frac{F_i}{\Lambda^4} \mathcal{O}_i$$

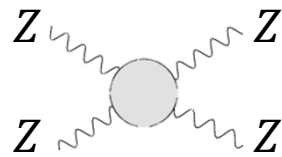


$V = Z, W^\pm, \gamma$

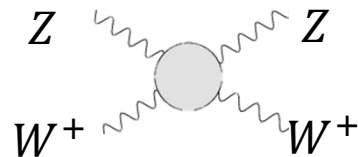


$A_{++}^{ZZ}, A_{+0}^{ZZ}, A_{+-}^{ZZ}, A_{++}^{ZW}, A_{+0}^{ZW}, A_{+-}^{ZW}, \dots \Rightarrow > 20 \text{ constraints}$
 (on 18 parameters)

$$\begin{aligned}
 F_{S,0} \text{Tr} \left[(D_\mu \Phi)^\dagger D_\nu \Phi \right]^2 & \quad \text{Tr} \left[(D_\mu \Phi)^\dagger D_\nu \Phi \right] = M_W^2 W_\mu^- W_\nu^+ \left(1 + \frac{H}{v} \right)^2 + \frac{M_Z^2}{2} Z_\mu Z_\nu \left(1 + \frac{H}{v} \right)^2 \\
 F_{S,1} \text{Tr} \left[(D_\mu \Phi)^\dagger D^\mu \Phi \right]^2 & \quad + \frac{1}{2} \partial_\mu H \partial_\nu H + \frac{iM_Z}{2} (Z_\mu \partial_\nu H - Z_\nu \partial_\mu H) \left(1 + \frac{H}{v} \right)
 \end{aligned}$$



$$F_{S,0} + F_{S,1} > 0$$



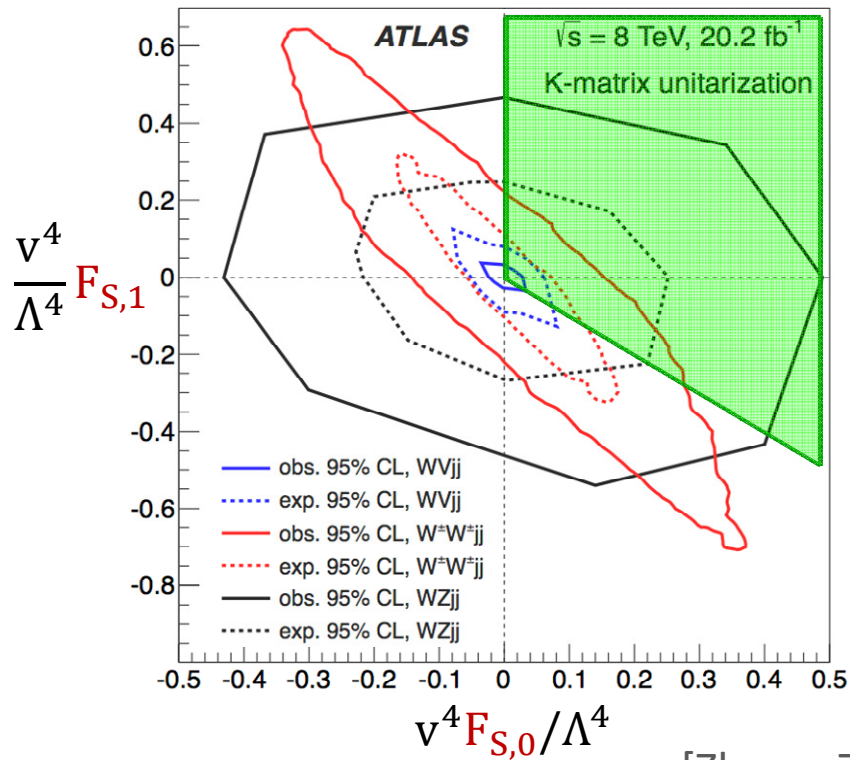
$$F_{S,0} > 0$$

$$F_{S,0} \text{Tr} \left[(D_\mu \Phi)^\dagger D_\nu \Phi \right]^2$$

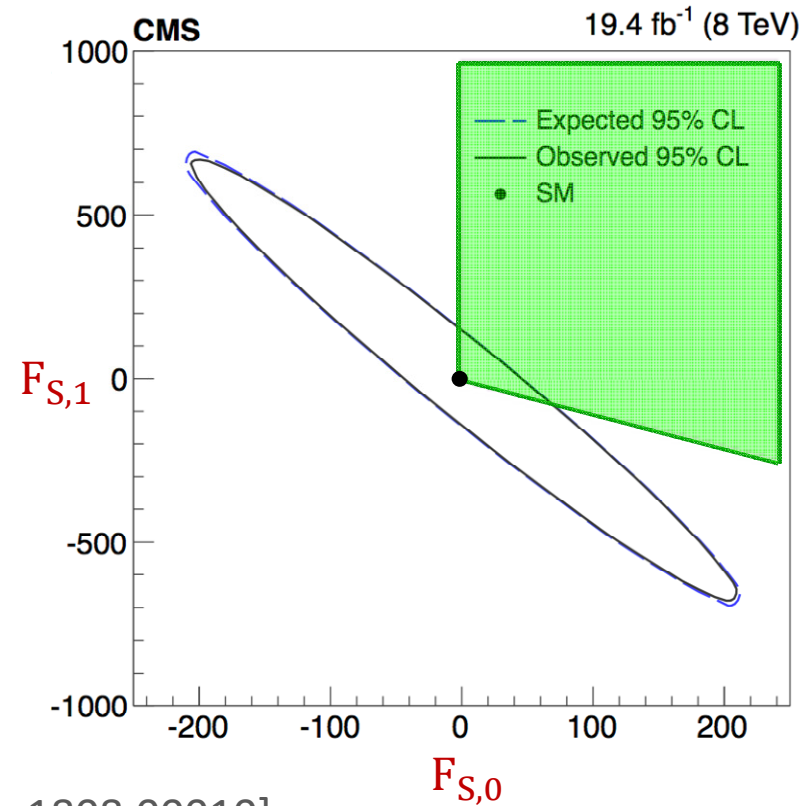
$$F_{S,1} \text{Tr} \left[(D_\mu \Phi)^\dagger D^\mu \Phi \right]^2$$

$$F_{S,0} > 0$$

$$F_{S,0} + F_{S,1} > 0$$



[Zhang, Zhou, 1808.00010]



Massive Gravity



Cosmological Constant (Problem)

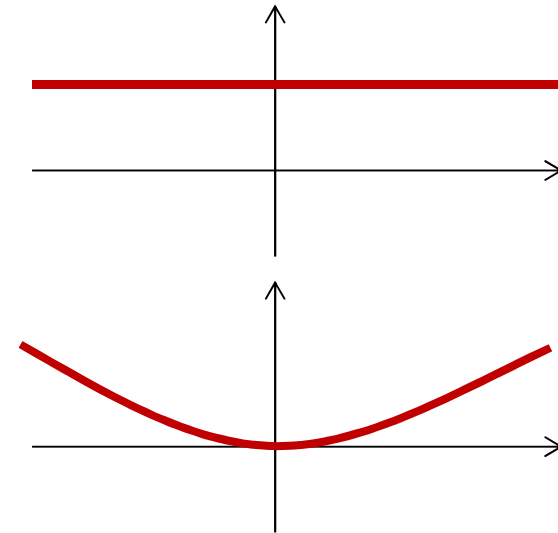
The Universe is accelerating...

$$\int d^4x \sqrt{-g} (M_P^2 R[g] + \Lambda^4)$$

or

$$\int d^4x \sqrt{-g} M_P^2 (R[g] - m^2 V(g - \bar{g}))$$

?



Cosmological Constant (Problem)

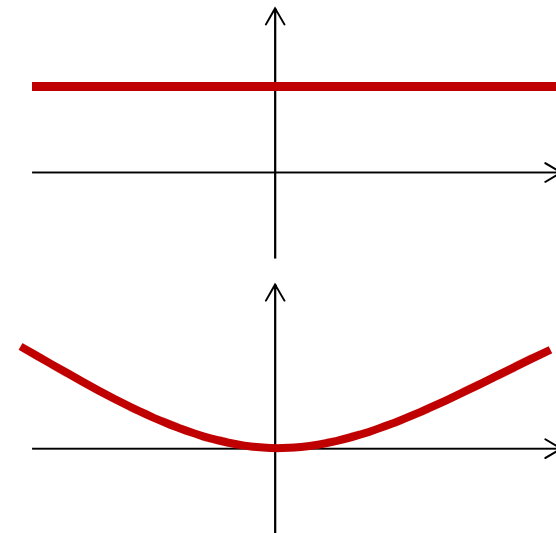
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$$\int d^4x \sqrt{-g} (M_P^2 R[g] + \Lambda^4)$$

or

$$\int d^4x \sqrt{-g} M_P^2 (R[g] - m^2 V(g - \bar{g}))$$

$$? \quad \parallel \quad h^2 + c h^3 + d h^4 + \dots$$



Pros

- Small m^2 technically natural
- Regulates IR divergences

Cons



- More free parameters
- Strongly coupled at $\Lambda_5 = (m^4 M_P)^{1/5}$

Massive Gravity

Massive Graviton Interactions

$$\int d^4x \sqrt{-g} M_P^2 (R[g] - m^2 V(g - \bar{g}))$$

$$h_{\mu\nu} \rightarrow \frac{\partial_\mu \partial_\nu \pi}{m^2 M_P} + \frac{\partial_{(\mu} A_{\nu)}}{m M_P} + \frac{h_{\mu\nu}}{M_P}$$

$$V(h) = [h^2] - [h]^2$$

$$+ c_1 [h^3] + c_2 [h^2][h]$$

$$+ d_1 [h^4] + d_2 [h^2]^2$$

$$M_P^2 m^2 V(h) = m^2 h^2 + F^2 + \partial(\dots)$$

$$+ \frac{C_5}{\Lambda_5^5} (\partial\partial\pi)^3 + \frac{C_3}{\Lambda_3^3} (\partial\partial\pi)^2 h + \dots$$

$$+ \frac{D_5 m^2}{\Lambda_5^{10}} (\partial\partial\pi)^4 + \frac{D_3}{\Lambda_3^6} (\partial\partial\pi)^3 h + \dots$$



Pros

Small m^2 technically natural

Regulates IR divergences

Cons

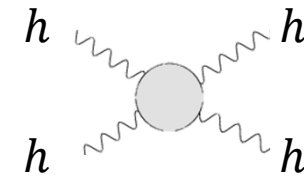
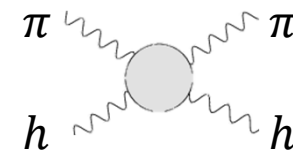
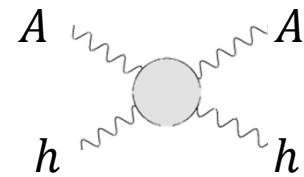
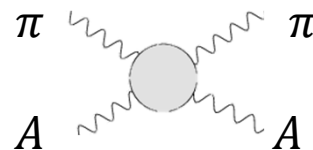
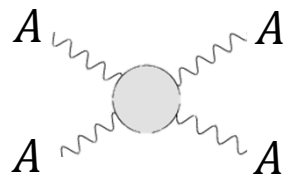
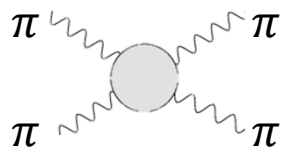


More free parameters

Strongly coupled at $\Lambda_5 = (m^4 M_P)^{1/5}$

Massive Gravity

$$\mathcal{L}_{\text{EFT}} = \mathcal{L}_{\text{GR}} + \frac{C_5}{\Lambda_5^5} \mathcal{O}_3[\pi, A, h] + \frac{D_5}{\Lambda_5^{10}} \mathcal{O}_4[\pi, A, h] \\ + \frac{C_3}{\Lambda_3^3} \mathcal{O}'_3[\pi, A, h] + \frac{D_3}{\Lambda_3^6} \mathcal{O}'_4[\pi, A, h]$$



Pros

- Small m^2 technically natural
- Regulates IR divergences

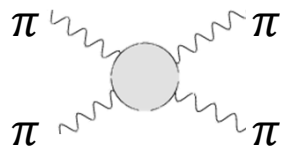


Cons

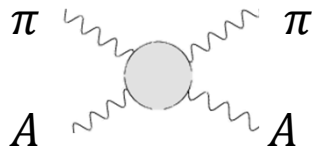
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$$D_5 = 0$$



$$C_5 = 0$$



Pros

Small m^2 technically natural
Regulates IR divergences



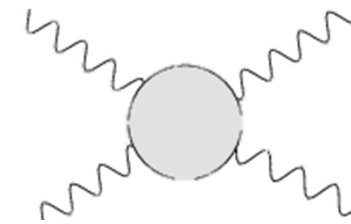
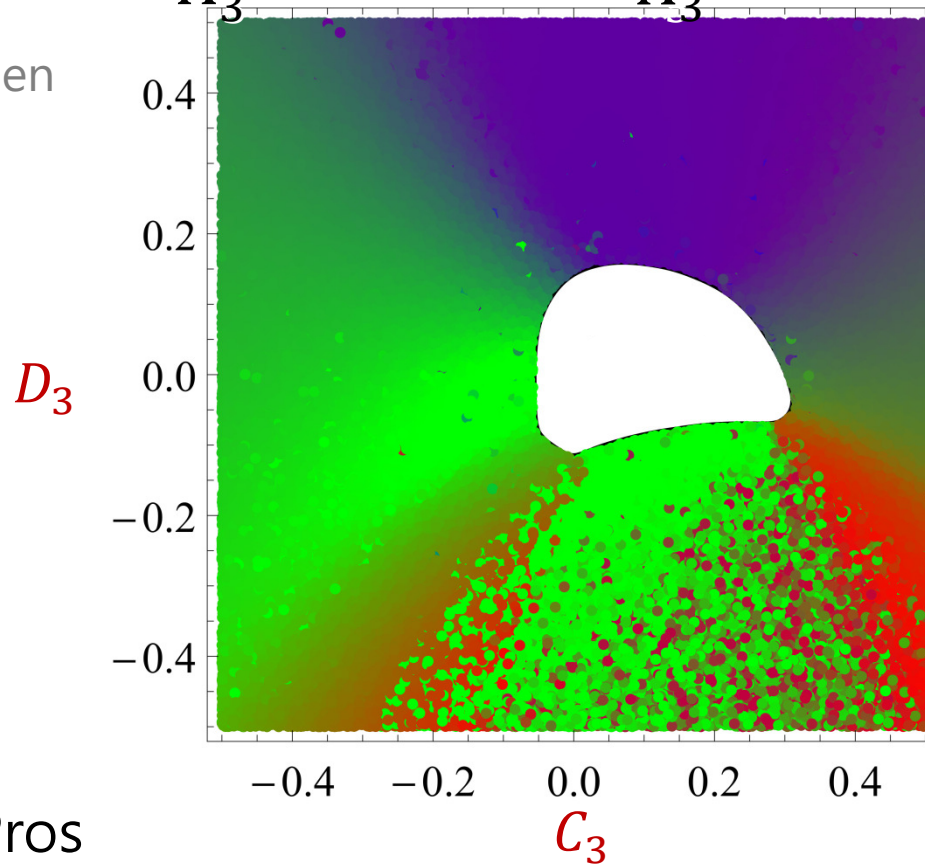
Cons

More free parameters
~~Strongly coupled at $\Lambda_5 = (m^4 M_P)^{1/5}$~~

Massive Gravity

$$\mathcal{L}_{\text{EFT}} = \mathcal{L}_{\text{GR}} + \frac{C_3}{\Lambda_3^3} \mathcal{O}'_3[\pi, A, h] + \frac{D_3}{\Lambda_3^6} \mathcal{O}'_4[\pi, A, h]$$

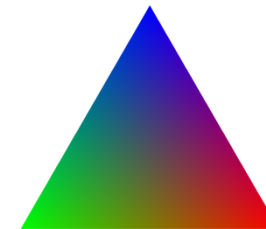
[Cheung, Remmen
1601.04068]



Tensor

Vector

Scalar



Pros

- Small m^2 technically natural
- Regulates IR divergences

Cons

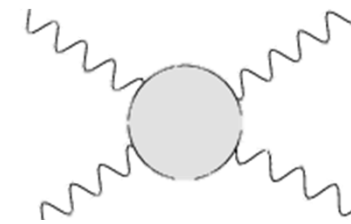
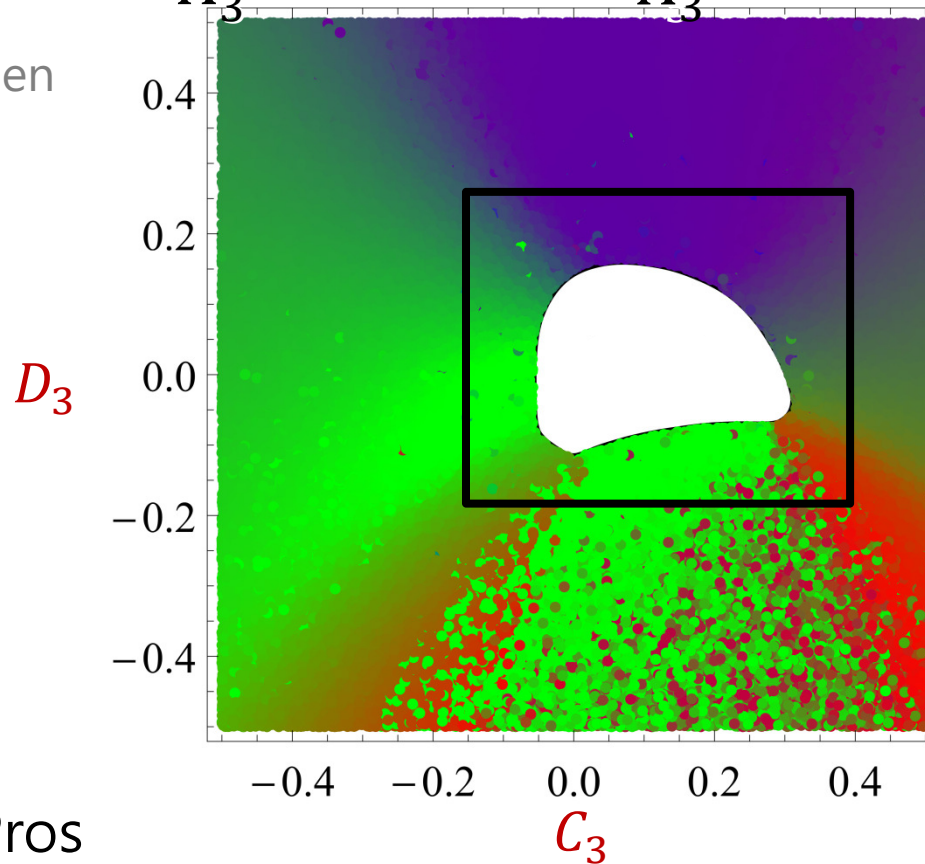


- ~~More free parameters~~
- ~~Strongly coupled at $\Lambda_5 = (m^4 M_P)^{1/5}$~~

Massive Gravity

$$\mathcal{L}_{\text{EFT}} = \mathcal{L}_{\text{GR}} + \frac{C_3}{\Lambda_3^3} \mathcal{O}'_3[\pi, A, h] + \frac{D_3}{\Lambda_3^6} \mathcal{O}'_4[\pi, A, h]$$

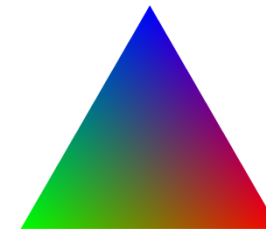
[Cheung, Remmen
1601.04068]



Tensor

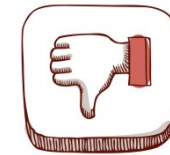
Vector

Scalar



Pros

- Small m^2 technically natural
- Regulates IR divergences

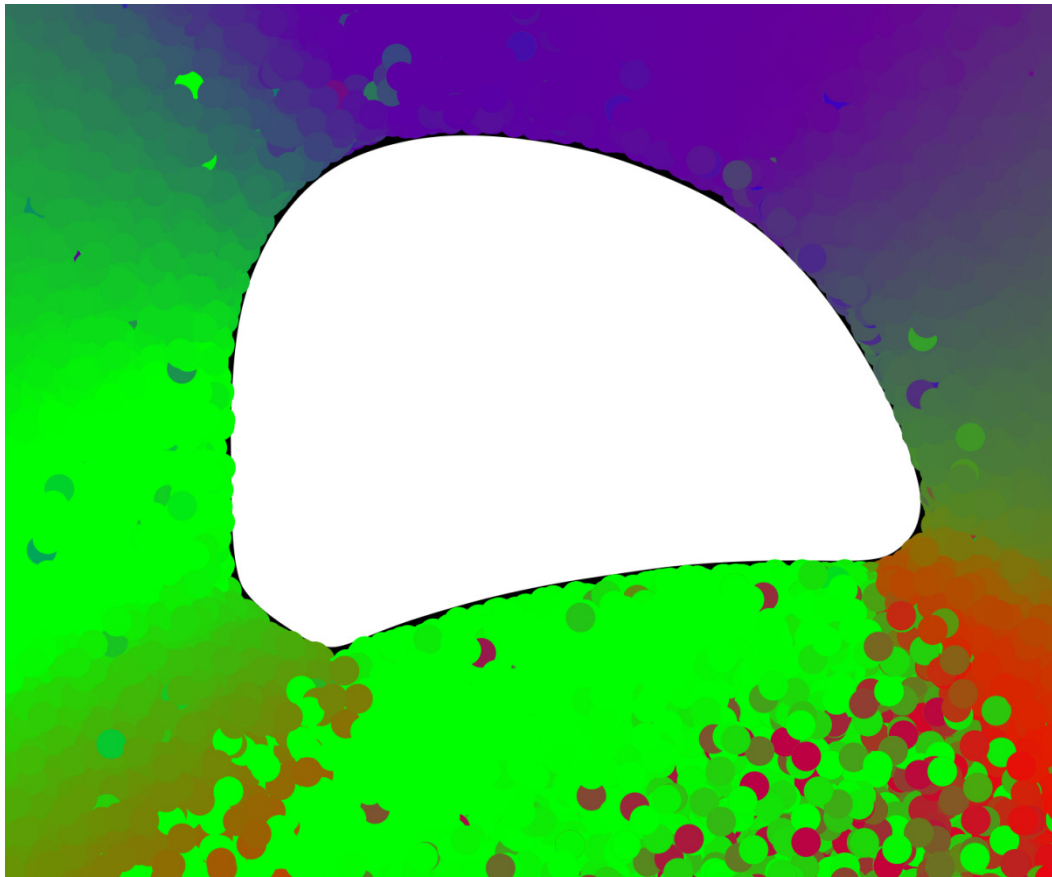


Cons

- ~~More free parameters~~
- ~~Strongly coupled at $\Lambda_5 = (m^4 M_P)^{1/5}$~~

Massive Gravity

$$\mathcal{L}_{\text{EFT}} = \mathcal{L}_{\text{GR}} + \frac{C_3}{\Lambda_3^3} \mathcal{O}'_3[\pi, A, h] + \frac{D_3}{\Lambda_3^6} \mathcal{O}'_4[\pi, A, h]$$

 D_3  C_3

Massive Gravity

$$\mathcal{L}_{\text{EFT}} = \mathcal{L}_{\text{GR}} + \frac{C_3}{\Lambda_3^3} \mathcal{O}'_3[\pi, A, h] + \frac{D_3}{\Lambda_3^6} \mathcal{O}'_4[\pi, A, h]$$

$$\rightarrow \frac{1}{g_*^2} \mathcal{L}_{\text{GR}}[g_* h] \mp \frac{g_* C_3}{\Lambda_3^3} \mathcal{O}'_3[\pi, A, h] + \frac{g_*^2 D_3}{\Lambda_3^6} \mathcal{O}'_4[\pi, A, h]$$

D_3



— $g_*^4 = 10^{-9} \frac{m}{\text{eV}}$

— $g_*^4 = 10^{-8} \frac{m}{\text{eV}}$

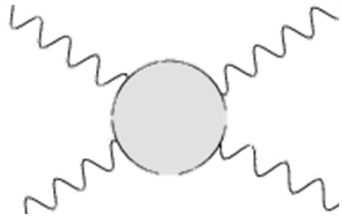
— $g_*^4 = 10^{-6} \frac{m}{\text{eV}}$

C_3

Summary

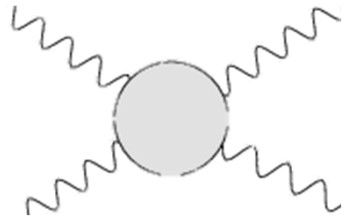


$$\frac{1}{g_*^2} \mathcal{L}_{\text{EFT}} \left[\frac{g_* \Phi}{M}, \frac{\partial}{M}; c_n \right]$$



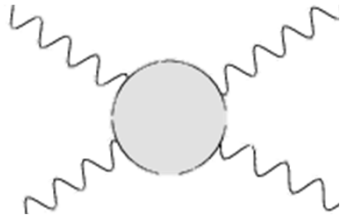
$$\frac{1}{g_*^2} \mathcal{L}_{\text{EFT}} \left[\frac{g_* \Phi}{M}, \frac{\partial}{M}; c_n \right]$$

Positivity Constraints



$$\partial_s^N A(s) > 0$$

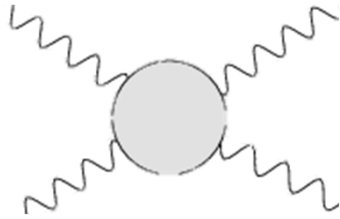
$$\frac{1}{g_*^2} \mathcal{L}_{\text{EFT}} \left[\frac{g_* \Phi}{M}, \frac{\partial}{M}; c_n \right]$$



Positivity Constraints

$$\partial_s^N A(s) > \int d\mu C_N(s, \mu) \sigma(\mu) > 0$$

$$\frac{1}{g_*^2} \mathcal{L}_{\text{EFT}} \left[\frac{g_* \Phi}{M}, \frac{\partial}{M}; c_n \right]$$

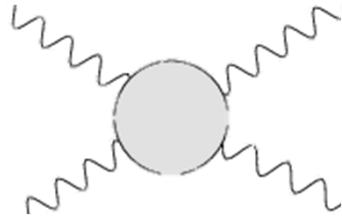


Positivity Constraints

$$\partial_s^N A(s) > \int d\mu C_N(s, \mu) \sigma(\mu) > 0$$

$$\partial_t \partial_s^N A(s, t) > \frac{2N+1}{M^2} \partial_s^N A(s, t)$$

$$\frac{1}{g_*^2} \mathcal{L}_{\text{EFT}} \left[\frac{g_* \Phi}{M}, \frac{\partial}{M}; c_n \right]$$



Positivity Constraints

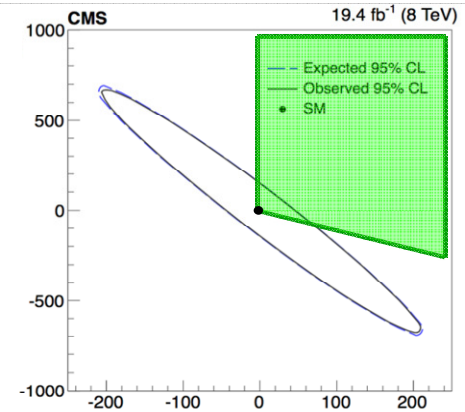
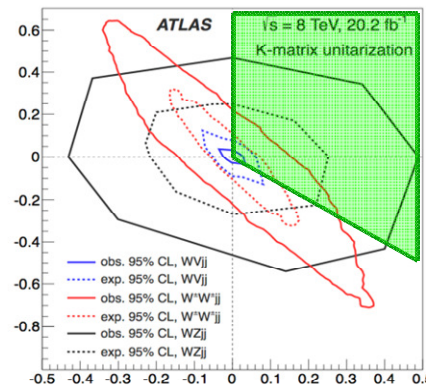
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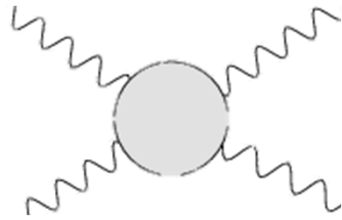
SMEFT

$$F_{S,0} \text{Tr} \left[(D_\mu \Phi)^\dagger D_\nu \Phi \right]^2$$

$$F_{S,1} \text{Tr} \left[(D_\mu \Phi)^\dagger D^\mu \Phi \right]^2$$



$$\frac{1}{g_*^2} \mathcal{L}_{\text{EFT}} \left[\frac{g_* \Phi}{M}, \frac{\partial}{M}; c_n \right]$$



Positivity Constraints

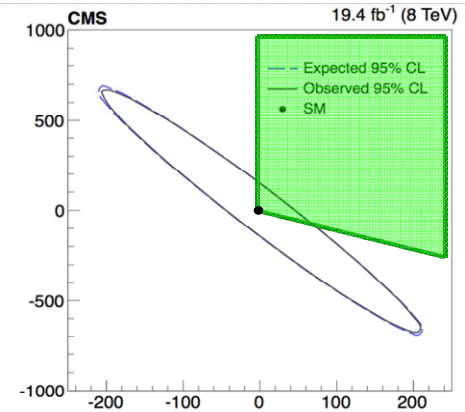
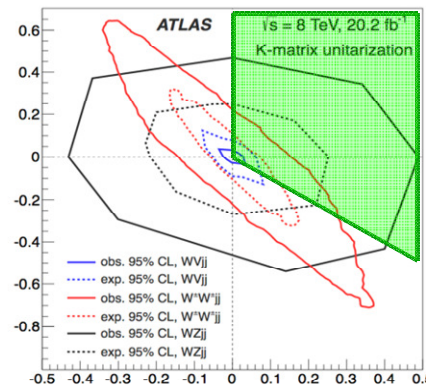
$$\partial_S^N A(s) > \int d\mu C_N(s, \mu) \sigma(\mu) > 0$$

$$\partial_t \partial_S^N A(s, t) > \frac{2N+1}{M^2} \partial_S^N A(s, t)$$

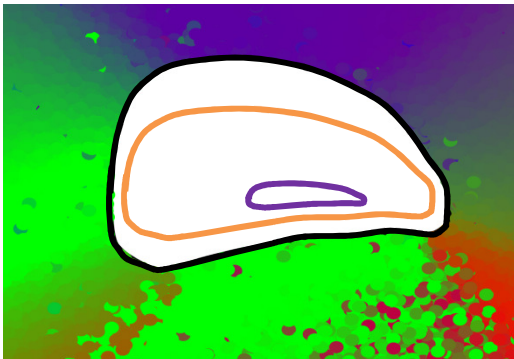
SMEFT

$$F_{S,0} \text{Tr} \left[(D_\mu \Phi)^\dagger D_\nu \Phi \right]^2$$

$$F_{S,1} \text{Tr} \left[(D_\mu \Phi)^\dagger D^\mu \Phi \right]^2$$



D_3



C_3

$$C_5 = 0$$

$$D_5 = 0$$

Massive Gravity

Summary