



# Heavy physics contributions to neutrinoless double beta decay from QCD

Nicolas Garron

University of Cambridge, HEP seminar, 19th of October, 2018

# CalLat (California Lattice)

red = postdoc and blue = grad student

- Jülich: Evan Berkowitz
- LBL/UCB: Davd Brantley, Chia Cheng (Jason) Chang, Thosrsten Kurth, Henry Monge-Camacho, André Walker-Loud
- NVIDIA: K Clark
- Liverpool: Nicolas Garron
- JLab: Balint Joó
- Rutgers: Chris Monahan
- North Carolina: Amy Nicholson
- City College of New York: Brian Tiburzi
- RIKEN/BNL: Enrico Rinaldi
- LLNL: Pavlos Vranas

• Observation of Neutrino oscillations, accumulation of evidences since the late 60's: solar  $\nu$ , atmospheric  $\nu$ ,  $\nu$  beam, ....

2015 Nobel prize in physics: Kajita and McDonald

- $\Rightarrow$  Neutrinos have non-zero mass
- $\Rightarrow$  Deviation from the Standard Model

- Observation of Neutrino oscillations, accumulation of evidences since the late 60's: solar  $\nu$ , atmospheric  $\nu$ ,  $\nu$  beam, ....
  - 2015 Nobel prize in physics: Kajita and McDonald
  - $\Rightarrow$  Neutrinos have non-zero mass
  - $\Rightarrow$  Deviation from the Standard Model
- Mass hierarchy and mixing pattern remain a puzzle

In particular, what is the nature of the neutrino mass, Dirac or Majorana ?

- Observation of Neutrino oscillations, accumulation of evidences since the late 60's: solar  $\nu$ , atmospheric  $\nu$ ,  $\nu$  beam, ....
  - 2015 Nobel prize in physics: Kajita and McDonald
  - $\Rightarrow$  Neutrinos have non-zero mass
  - $\Rightarrow$  Deviation from the Standard Model
- Mass hierarchy and mixing pattern remain a puzzle In particular, what is the nature of the neutrino mass, Dirac or Majorana ?
- Experimental searches for neutrinoless double  $\beta$  decay  $(0\nu\beta\beta)$ If measured  $\rightarrow$  Majorana particle, probe of new physics, ... Huge experimental effort

■ *β*-decay

 $n \longrightarrow p + e^- + \bar{\nu}_e$ 

• and a  $\nu_e$  can be absorbed in the process

 $\nu_e + n \longrightarrow p + e^-$ 

• so that if  $\nu_e = \bar{\nu}_e$  it is possible to have

 $n + n \longrightarrow p + p + e^{-} + e^{-}$ 

 $\Rightarrow$  Neutrinoless double beta decay

Neutrinoless double  $\beta$  decay:  $n + n \longrightarrow p + p + e^- + e^-$ 



- $0\nu\beta\beta$  violates Lepton-number conservation  $\Rightarrow$  New Physics
- Can be related to leptogensis and Matter-Antimatter asymmetry
- Can probe the absolute scale of neutrino mass (or of new physics)
- Related to dark matter ?
- Worldwide experimental effort

- Completed experiments:
  - Gotthard TPC
  - Heidelberg-Moscow, <sup>76</sup>Ge detectors (1997–2001)
  - IGEX, <sup>76</sup>Ge detectors (1999–2002)<sup>[17]</sup>
  - NEMO, various isotopes using tracking calorimeters (2003-2011)
  - Cuoricino, <sup>130</sup>Te in ultracold TeO<sub>2</sub> crystals (2003–2008)<sup>[18]</sup>
- Experiments taking data as of November 2017:
  - COBRA, <sup>116</sup>Cd in room temperature CdZnTe crystals
  - CUORE, <sup>130</sup>Te in ultracold TeO<sub>2</sub> crystals
  - EXO, a <sup>136</sup>Xe and <sup>134</sup>Xe search
  - GERDA, a <sup>76</sup>Ge detector
  - KamLAND-Zen, a <sup>136</sup>Xe search. Data collection from 2011.<sup>[18]</sup>
  - MAJORANA, using high purity <sup>76</sup>Ge p-type point-contact detectors.<sup>[19]</sup>
  - XMASS using liquid Xe
- Proposed/future experiments:
  - CANDLES, <sup>48</sup>Ca in CaF<sub>2</sub>, at Kamioka Observatory
  - MOON, developing <sup>100</sup>Mo detectors
  - AMoRE, <sup>100</sup>Mo enriched CaMoO<sub>4</sub> crystals at YangYang underground laboratory<sup>[20]</sup>
  - nEXO, using liquid <sup>136</sup>Xe in a time projection chamber <sup>[21]</sup>
  - LEGEND, Neutrinoless Double-beta Decay of <sup>76</sup>Ge.
  - LUMINEU, exploring <sup>100</sup>Mo enriched ZnMoO<sub>4</sub> crystals at LSM, France.
  - NEXT, a Xenon TPC. NEXT-DEMO ran and NEXT-100 will run in 2016.
  - SNO+, a liquid scintillator, will study <sup>130</sup>Te
  - SuperNEMO, a NEMO upgrade, will study <sup>82</sup>Se
  - TIN.TIN, a <sup>124</sup>Sn detector at INO

#### (source: Wikipedia)

- $0\nu\beta\beta$  violates Lepton-number conservation  $\Rightarrow$  New Physics
- Can be related to leptogensis and Matter-Antimatter asymmetry
- Can probe the absolute scale of neutrino mass (or of new physics)
- Related to dark matter ?
- Worldwide experimental effort
- Relating possible experimental signatures to New-Physics model requires the knowledge of QCD contributions

Computing the full process is very ambitious

- Different scales, different interactions
- Multi-particles in initial and final states
- Nucleon  $\Rightarrow$  Signal-to-noise problem

Very hard task in Lattice QCD

| gА |  |
|----|--|
|    |  |

#### The axial coupling of the nucleon





[C Chang, A Nicholson, E Rinaldi, E Berkowitz, NG, D Brantley, H Monge-Camacho, C Monahan, C Bouchard, M Clark, B Joó, T Kurth, K Orginos, P Vranas, A Walker-Loud]

Nature 558 (2018) no.7708

#### $0\nu\beta\beta$ and EFT

Process can be mediated by light or heavy particle

- E.g. light  $\nu_L$  or heavy  $\nu_R$  through seesaw mechanism
- Or heavy "New-Physics" particle

Process can be mediated by light or heavy particle

- E.g. light  $\nu_L$  or heavy  $\nu_R$  through seesaw mechanism
- Or heavy "New-Physics" particle
- Naively, one expects the long-distance contribution of a light neutrino to dominate over the short-distance contribution of a heavy particle
- But the long-range interaction requires a helicity flip and its proportional to the mass of the light neutrino
- $\Rightarrow$  Relative size of the different contributions depend on the New Physics model

Process can be mediated by light or heavy particle

- E.g. light  $\nu_L$  or heavy  $\nu_R$  through seesaw mechanism
- Or heavy "New-Physics" particle
- Naively, one expects the long-distance contribution of a light neutrino to dominate over the short-distance contribution of a heavy particle
- But the long-range interaction requires a helicity flip and its proportional to the mass of the light neutrino
- $\Rightarrow$  Relative size of the different contributions depend on the New Physics model
- Standard seesaw  $m_l \sim M_D^2/M_R \ll m_h \sim M_R$

#### $0\nu\beta\beta$ and EFT

Consider "heavy" particles contributions, integrate out heavy d.o.f.

Consider "heavy" particles contributions, integrate out heavy d.o.f.

EFT framework, see e.g. [Prézeau, Ramsey-Musolf, Vogel '03], the LO contributions are  $\pi^- \longrightarrow \pi^+ + e^- + e^-$ 

- $\blacksquare n \longrightarrow p + \pi^+ + e^- + e^-$
- $\blacksquare n + n \longrightarrow p + p + e^- + e^-$



Consider "heavy" particles contributions, integrate out heavy d.o.f.

EFT framework, see e.g. [Prézeau, Ramsey-Musolf, Vogel '03], the LO contributions are  $\pi^- \longrightarrow \pi^+ + e^- + e^-$ 

 $\blacksquare n \longrightarrow p + \pi^+ + e^- + e^-$ 

$$\blacksquare n + n \longrightarrow p + p + e^- + e^-$$



In this work we focus on the  $\pi^- \longrightarrow \pi^+$  matrix elements

- On the lattice, compute the Matrix elements of  $\pi^- \longrightarrow \pi^+$  transitions
- Extract the LEC through Chiral fits
- Use the LEC in the EFT framework to estimate a physical amplitude

#### Lattice Computation of $\pi^- \rightarrow \pi^+$ matrix elements

#### 4-quark operators

We only consider light valence quarks q = u, d

the operators of interest are

$$\mathcal{O}_{1+}^{++} = \left(\bar{q}_L \tau^+ \gamma^\mu q_L\right) \left[\bar{q}_R \tau^+ \gamma_\mu q_R\right]$$
  

$$\mathcal{O}_{2+}^{++} = \left(\bar{q}_R \tau^+ q_L\right) \left[\bar{q}_R \tau^+ q_L\right] + \left(\bar{q}_L \tau^+ q_R\right) \left[\bar{q}_L \tau^+ q_R\right]$$
  

$$\mathcal{O}_{3+}^{++} = \left(\bar{q}_L \tau^+ \gamma^\mu q_L\right) \left[\bar{q}_L \tau^+ \gamma_\mu q_L\right] + \left(\bar{q}_R \tau^+ \gamma^\mu q_R\right) \left[\bar{q}_R \tau^+ \gamma_\mu q_R\right]$$

#### 4-quark operators

We only consider light valence quarks q = u, d

the operators of interest are

$$\mathcal{O}_{1+}^{++} = \left(\bar{q}_L \tau^+ \gamma^\mu q_L\right) \left[\bar{q}_R \tau^+ \gamma_\mu q_R\right]$$
$$\mathcal{O}_{2+}^{++} = \left(\bar{q}_R \tau^+ q_L\right) \left[\bar{q}_R \tau^+ q_L\right] + \left(\bar{q}_L \tau^+ q_R\right) \left[\bar{q}_L \tau^+ q_R\right]$$
$$\mathcal{O}_{3+}^{++} = \left(\bar{q}_L \tau^+ \gamma^\mu q_L\right) \left[\bar{q}_L \tau^+ \gamma_\mu q_L\right] + \left(\bar{q}_R \tau^+ \gamma^\mu q_R\right) \left[\bar{q}_R \tau^+ \gamma_\mu q_R\right]$$

and the colour partner

$$\mathcal{O}_{1+}^{'++} = \left(\bar{q}_L \tau^+ \gamma^\mu q_L\right] \left[\bar{q}_R \tau^+ \gamma_\mu q_R\right)$$
$$\mathcal{O}_{2+}^{'++} = \left(\bar{q}_L \tau^+ q_L\right] \left[\bar{q}_L \tau^+ q_L\right) + \left(\bar{q}_R \tau^+ q_R\right] \left[\bar{q}_R \tau^+ q_R\right)$$

where () []  $\equiv$  color unmix and (] [)  $\equiv$  color unmix

# 4-quark operators (II)

In a slightly more human readable way

$$\begin{aligned} \mathcal{O}_{1+}^{++} &= \left(\bar{q}_L \tau^+ \gamma^\mu q_L\right) \left[\bar{q}_R \tau^+ \gamma_\mu q_R\right] \\ \mathcal{O}_{2+}^{++} &= \left(\bar{q}_R \tau^+ q_L\right) \left[\bar{q}_R \tau^+ q_L\right] + \left(\bar{q}_L \tau^+ q_R\right) \left[\bar{q}_L \tau^+ q_R\right] \\ \mathcal{O}_{3+}^{++} &= \left(\bar{q}_L \tau^+ \gamma^\mu q_L\right) \left[\bar{q}_L \tau^+ \gamma_\mu q_L\right] + \left(\bar{q}_R \tau^+ \gamma^\mu q_R\right) \left[\bar{q}_R \tau^+ \gamma_\mu q_R\right] \end{aligned}$$

The colour unmix are

$$\mathcal{O}_{3+}^{++} \sim \gamma_{L}^{\mu} \times \gamma_{L}^{\mu} + \gamma_{R}^{\mu} \times \gamma_{R}^{\mu} \longrightarrow VV + AA$$

$$\begin{array}{lll} \mathcal{O}_{1+}^{++} & \sim & \gamma_L^{\mu} \times \gamma_R^{\mu} \longrightarrow VV - AA \\ \mathcal{O}_{2+}^{++} & \sim & P_L \times P_L + P_R \times P_R \longrightarrow SS + PP \end{array}$$

# 4-quark operators (II)

In a slightly more human readable way

$$\begin{aligned} \mathcal{O}_{1+}^{++} &= \left(\bar{q}_L \tau^+ \gamma^\mu q_L\right) \left[\bar{q}_R \tau^+ \gamma_\mu q_R\right] \\ \mathcal{O}_{2+}^{++} &= \left(\bar{q}_R \tau^+ q_L\right) \left[\bar{q}_R \tau^+ q_L\right] + \left(\bar{q}_L \tau^+ q_R\right) \left[\bar{q}_L \tau^+ q_R\right] \\ \mathcal{O}_{3+}^{++} &= \left(\bar{q}_L \tau^+ \gamma^\mu q_L\right) \left[\bar{q}_L \tau^+ \gamma_\mu q_L\right] + \left(\bar{q}_R \tau^+ \gamma^\mu q_R\right) \left[\bar{q}_R \tau^+ \gamma_\mu q_R\right] \end{aligned}$$

The colour unmix are

$$\begin{array}{lll} \mathcal{O}_{3+}^{++} & \sim & \gamma_L^{\mu} \times \gamma_L^{\mu} + \gamma_R^{\mu} \times \gamma_R^{\mu} \longrightarrow VV + AA \\ \\ \mathcal{O}_{1+}^{++} & \sim & \gamma_L^{\mu} \times \gamma_R^{\mu} \longrightarrow VV - AA \\ \\ \mathcal{O}_{2+}^{++} & \sim & P_L \times P_L + P_R \times P_R \longrightarrow SS + PP \end{array}$$

and the colour partner

$$\begin{array}{rcl} \mathcal{O}_{1+}^{'++} & \longrightarrow & (VV - AA)_{mix} \sim (SS - PP)_{unmix} \\ \mathcal{O}_{2+}^{'++} & \longrightarrow & (SS + PP)_{mix} \sim (SS + PP)_{unmix} + c(TT)_{unmix} \end{array}$$

Nicolas Garron (University of Liverpool)

#### $\pi^- \to \pi^+$ transition

- We have to compute the matrix elements (ME) of  $\langle \pi^+ | {\cal O} | \pi^- 
  angle$
- Since QCD conserves Parity, we only consider Parity even sector

- We have to compute the matrix elements (ME) of  $\langle \pi^+ | \mathcal{O} | \pi^- \rangle$
- Since QCD conserves Parity, we only consider Parity even sector

The computation goes along the lines of  $\Delta F = 2$  ME:

- Extract the bare ME by fitting 3p and 2p functions or ratios
- Non-Perturbative Renormalisation
- Global Fit, extrapolation to physical pion mass and continuum limit

- Lattice QCD is a discretised version of Euclidean QCD
- Well-defined regularisation of the theory
- Gauge invariant (Wilson) at finite lattice spacing
- Continuum Euclidean QCD is recovered in the lmit  $a \rightarrow 0$

$$\langle O \rangle_{continuum} = \lim_{a \to 0} \lim_{V \to \infty} \langle O \rangle_{latt}$$

Allows for non-perturbative and first-principle determinations of QCD observables

Various steps of a Lattice computation (schematically)

 Generate gauge configurations (ensembles) ↔ gluons and sea quarks (or take already existing ones)

- Generate gauge configurations (ensembles) ↔ gluons and sea quarks (or take already existing ones)
- $\blacksquare$  Compute fermion propagators  $\leftrightarrow$  valence quarks

- Generate gauge configurations (ensembles) ↔ gluons and sea quarks (or take already existing ones)
- Compute fermion propagators  $\leftrightarrow$  valence quarks
- $\blacksquare$  Compute Wick contractions  $\leftrightarrow$  bare Green functions

- Generate gauge configurations (ensembles) ↔ gluons and sea quarks (or take already existing ones)
- Compute fermion propagators ↔ valence quarks
- $\blacksquare$  Compute Wick contractions  $\leftrightarrow$  bare Green functions
- Determine Z factors (if needed)  $\leftrightarrow$  renormalised Green functions

- Generate gauge configurations (ensembles) ↔ gluons and sea quarks (or take already existing ones)
- Compute fermion propagators  $\leftrightarrow$  valence quarks
- Compute Wick contractions  $\leftrightarrow$  bare Green functions
- Determine Z factors (if needed)  $\leftrightarrow$  renormalised Green functions
- $\blacksquare$  Continuum & physical pion mass extrapolations  $\leftrightarrow$  physical observables
## Remarks

Different discretizations of the Dirac operators are possible: Wilson, staggered, Twisted-mass, etc.

One difficulty is to maintain the symmetries of the continuum lagrangian at finite lattice spacing,

 $\Rightarrow$  choose the discretization adapted to the situation you want to describe

## Remarks

Different discretizations of the Dirac operators are possible: Wilson, staggered, Twisted-mass, etc.

One difficulty is to maintain the symmetries of the continuum lagrangian at finite lattice spacing,

 $\Rightarrow$  choose the discretization adapted to the situation you want to describe

In particular chiral symmetry is notoriously difficult to maintain

We consider here Domain-Wall fermions, a type of discretisation which respects chiral and flavour symmetry almost exactly.

The price to pay is a high numerical cost

#### This computation

The main features of our computation are:

- Mixed-action: Möbius Domain-Wall on  $N_f = 2 + 1 + 1$  HISQ configurations
- 3 lattice spacings, pion mass down to the physical value

The main features of our computation are:

- Mixed-action: Möbius Domain-Wall on  $N_f = 2 + 1 + 1$  HISQ configurations
- 3 lattice spacings, pion mass down to the physical value

As a consequence:

- Chiral-flavour symmetry maintained (in the valence sector)
- Lattice artefact of order  $O(a^2)$
- Good control over the chiral behaviour, continuum limit, finite volume effects
- But non-unitary setup and flavour symmetry broken in the sea

The main features of our computation are:

- Mixed-action: Möbius Domain-Wall on  $N_f = 2 + 1 + 1$  HISQ configurations
- 3 lattice spacings, pion mass down to the physical value

As a consequence:

- Chiral-flavour symmetry maintained (in the valence sector)
- Lattice artefact of order  $O(a^2)$
- Good control over the chiral behaviour, continuum limit, finite volume effects
- But non-unitary setup and flavour symmetry broken in the sea
- I am not entering the *staggered* debate
- $\blacksquare$  We take the mixed-action terms into account in the  $\chi {\rm PT}$  expressions

The main features of our computation are:

- Mixed-action: Möbius Domain-Wall on  $N_f = 2 + 1 + 1$  HISQ configurations
- 3 lattice spacings, pion mass down to the physical value

As a consequence:

- Chiral-flavour symmetry maintained (in the valence sector)
- Lattice artefact of order  $\mathcal{O}(a^2)$
- Good control over the chiral behaviour, continuum limit, finite volume effects
- But non-unitary setup and flavour symmetry broken in the sea
- I am not entering the *staggered* debate
- $\blacksquare$  We take the mixed-action terms into account in the  $\chi {\rm PT}$  expressions

In addition we perform the renormalisation non-perturbatively Only perturbative errors come from the conversion to  $\overline{\rm MS}$ 

| HISQ gauge configuration parameters |               |                  |               |           |                           |                    |  |  |  |  |
|-------------------------------------|---------------|------------------|---------------|-----------|---------------------------|--------------------|--|--|--|--|
| abbr.                               | $N_{\rm cfg}$ | volume           | $\sim a$ [fm] | $m_l/m_s$ | $\sim m_{\pi_5}$<br>[MeV] | $\sim m_{\pi_5} L$ |  |  |  |  |
| a15m400                             | 1000          | $16^3 \times 48$ | 0.15          | 0.334     | 400                       | 4.8                |  |  |  |  |
| a15m350                             | 1000          | $16^3 \times 48$ | 0.15          | 0.255     | 350                       | 4.2                |  |  |  |  |
| a15m310                             | 1960          | $16^3 \times 48$ | 0.15          | 0.2       | 310                       | 3.8                |  |  |  |  |
| a15m220                             | 1000          | $24^3 \times 48$ | 0.15          | 0.1       | 220                       | 4.0                |  |  |  |  |
| a15m130                             | 1000          | $32^3 \times 48$ | 0.15          | 0.036     | 130                       | 3.2                |  |  |  |  |
| a12m400                             | 1000          | $24^3 \times 64$ | 0.12          | 0.334     | 400                       | 5.8                |  |  |  |  |
| a12m350                             | 1000          | $24^3 \times 64$ | 0.12          | 0.255     | 350                       | 5.1                |  |  |  |  |
| a12m310                             | 1053          | $24^3 \times 64$ | 0.12          | 0.2       | 310                       | 4.5                |  |  |  |  |
| a12m220S                            | 1000          | $24^3 \times 64$ | 0.12          | 0.1       | 220                       | 3.2                |  |  |  |  |
| a12m220                             | 1000          | $32^3 \times 64$ | 0.12          | 0.1       | 220                       | 4.3                |  |  |  |  |
| a12m220L                            | 1000          | $40^3 \times 64$ | 0.12          | 0.1       | 220                       | 5.4                |  |  |  |  |
| a12m130                             | 1000          | $48^3 \times 64$ | 0.12          | 0.036     | 130                       | 3.9                |  |  |  |  |
| a09m400                             | 1201          | $32^3 \times 64$ | 0.09          | 0.335     | 400                       | 5.8                |  |  |  |  |
| a09m350                             | 1201          | $32^3 \times 64$ | 0.09          | 0.255     | 350                       | 5.1                |  |  |  |  |
| a09m310                             | 784           | $32^3 \times 96$ | 0.09          | 0.2       | 310                       | 4.5                |  |  |  |  |
| a09m220                             | 1001          | $48^3\times96$   | 0.09          | 0.1       | 220                       | 4.7                |  |  |  |  |

# The setup (II)

#### For this analysis we only consider

|               | $m_\pi \sim 310~{ m MeV}$ |            | $m_\pi \sim 220~{ m MeV}$ |            | $m_\pi \sim 130~{ m MeV}$ |            |
|---------------|---------------------------|------------|---------------------------|------------|---------------------------|------------|
| <i>a</i> (fm) | V                         | $m_{\pi}L$ | V                         | $m_{\pi}L$ | V                         | $m_{\pi}L$ |
| 0.15          | $16^{3} \times 48$        | 3.78       | $24^{3} \times 48$        | 3.99       |                           |            |
| 0.12          |                           |            | $24^{3} \times 64$        | 3.22       |                           |            |
| 0.12          | $24^{3} \times 64$        | 4.54       | $32^{3} \times 64$        | 4.29       | $48^3 	imes 64$           | 3.91       |
| 0.12          |                           |            | $40^3 	imes 64$           | 5.36       |                           |            |
| 0.09          | $32^{3} \times 96$        | 4.50       | $48^3 	imes 96$           | 4.73       |                           |            |

#### Bare results

Define usual 2p and 3p functions

$$C_{\pi}(t) = \sum_{\mathbf{x}} \sum_{\alpha} \langle \alpha | \Pi^{+}(t, \mathbf{x}) \Pi^{-}(0, \mathbf{0}) | \alpha \rangle$$
$$= \sum_{n} \frac{|Z_{n}^{\pi}|^{2}}{2E_{n}^{\pi}} \left( e^{-E_{n}^{\pi}t} + e^{-E_{n}^{\pi}(T-t)} \right) + \cdots$$

where  $z_n^{\pi} = \langle \Omega | \Pi^+ | n \rangle$ ,  $\Omega = vaccum$  and

$$C_i^{\mathrm{3pt}}(t_f,t_i) = \sum_{\mathbf{x},\mathbf{y},lpha} \langle lpha | \Pi^+(t_f,\mathbf{x}) \mathcal{O}_i(0,\mathbf{0}) \Pi^+(t_i,\mathbf{y}) | lpha 
angle$$

#### Bare results

Define usual 2p and 3p functions

$$C_{\pi}(t) = \sum_{\mathbf{x}} \sum_{\alpha} \langle \alpha | \Pi^{+}(t, \mathbf{x}) \Pi^{-}(0, \mathbf{0}) | \alpha \rangle$$
$$= \sum_{n} \frac{|z_{n}^{\pi}|^{2}}{2E_{n}^{\pi}} \left( e^{-E_{n}^{\pi}t} + e^{-E_{n}^{\pi}(T-t)} \right) + \cdots$$

where  $z_n^{\pi} = \langle \Omega | \Pi^+ | n \rangle$ ,  $\Omega = vaccum$  and

$$C_i^{3 ext{pt}}(t_f, t_i) = \sum_{\mathbf{x}, \mathbf{y}, lpha} \langle lpha | \Pi^+(t_f, \mathbf{x}) \mathcal{O}_i(0, \mathbf{0}) \Pi^+(t_i, \mathbf{y}) | lpha 
angle$$

for example fit ratio such as

$$egin{array}{rcl} \mathcal{R}_i(t) &\equiv & C_i^{
m 3pt}(t,T-t)/\left(C_\pi(t)C_\pi(T-t)
ight) \ &\longrightarrow & rac{a^4\langle\pi|\mathcal{O}_{i+}^{++}|\pi
angle}{(a^2z_0^\pi)^2}+\dots \end{array}$$

#### Bare results



Example of results for  $a \simeq 0.12$  fm , near physical pion mass ensemble

#### Non Perturbative Renormalisation (NPR)

#### A few words on the renormalisation

First step: remove the divergences

Non-perturbative Renormalisation à la Rome-Southampton [Martinelli et al '95]

$$Q^{\textit{lat}}_i(a) 
ightarrow Q^{MOM}_i(\mu,a) = Z(\mu,a)_{ij} Q^{\textit{lat}}_j(a)$$

and take the continuum limit

$$Q_i^{MOM}(\mu,0) = \lim_{a^2 \to 0} Q_i^{MOM}(\mu,a)$$

Second step: Matching to  $\overline{\rm MS}$ , done in perturbation theory [Sturm et al., Lehner and Sturm, Gorbahn and Jäger, Gracey, ...]

$$Q_i^{MOM}(\mu,0) 
ightarrow Q_i^{\overline{ ext{MS}}}(\mu) = (1+r_1lpha_{\mathcal{S}}(\mu)+r_2lpha_{\mathcal{S}}(\mu)^2+\ldots)_{ij}Q_j^{MOM}(\mu,0)$$

Consider a quark bilinear  $O_{\Gamma} = \bar{\psi}_2 \Gamma \psi_1$ 

Define

 $\mathsf{\Pi}(x_2,x_1) = \langle \psi_2(x_2) \mathcal{O}_{\mathsf{\Gamma}}(0) \overline{\psi}_1(x_1) \rangle = \langle S_2(x_2,0) \mathsf{\Gamma} S_1(0,x_1) \rangle$ 

In Fourier space  $S(p) = \sum_{x} S(x, 0)e^{ip.x}$ 

 $\Pi(p_2,p_1) = \langle S_2(p_2) \Gamma S_1(p_1)^{\dagger}) \rangle$ 

Consider a quark bilinear  $O_{\Gamma} = \bar{\psi}_2 \Gamma \psi_1$ 

Define

 $\Pi(x_2, x_1) = \langle \psi_2(x_2) O_{\Gamma}(0) \overline{\psi}_1(x_1) \rangle = \langle S_2(x_2, 0) \Gamma S_1(0, x_1) \rangle$ 

In Fourier space  $S(p) = \sum_{x} S(x, 0)e^{ip.x}$ 

 $\Pi(p_2,p_1) = \langle S_2(p_2) \Gamma S_1(p_1)^{\dagger}) \rangle$ 

Amputated Green function

 $\Lambda(p_2, p_1) = \langle S_2(p_2)^{-1} \rangle \langle S_2(p_2) \Gamma S_1(p_1)^{\dagger} \rangle \rangle \langle (S_2(p_1)^{\dagger^{-1}}) \rangle$ 

Consider a quark bilinear  $O_{\Gamma} = \bar{\psi}_2 \Gamma \psi_1$ 

Define

 $\Pi(x_2,x_1) = \langle \psi_2(x_2) O_{\Gamma}(0) \overline{\psi}_1(x_1) \rangle = \langle S_2(x_2,0) \Gamma S_1(0,x_1) \rangle$ 

In Fourier space  $S(p) = \sum_{x} S(x, 0)e^{ip.x}$ 

 $\Pi(p_2,p_1) = \langle S_2(p_2) \Gamma S_1(p_1)^{\dagger}) \rangle$ 

Amputated Green function

 $\Lambda(p_2, p_1) = \langle S_2(p_2)^{-1} \rangle \langle S_2(p_2) \Gamma S_1(p_1)^{\dagger} \rangle \rangle \langle (S_2(p_1)^{\dagger^{-1}}) \rangle$ 

Rome Southampton original scheme (RI-MOM),  $p_1 = p_2 = p$  and  $\mu = \sqrt{p^2}$ 

$$Z(\mu, a) imes \lim_{m o 0} \operatorname{Tr}(\Gamma \Lambda(p, p))_{\mu^2 = p^2} = \operatorname{Tree}$$

Remarks

• Can be generalised to the four-quark operator mixing case

Remarks

- Can be generalised to the four-quark operator mixing case
- Non-perturbative off-shell and massless scheme(s)
- Requires gauge fixing (unlike Schrödinger Functional)

Remarks

- Can be generalised to the four-quark operator mixing case
- Non-perturbative off-shell and massless scheme(s)
- Requires gauge fixing (unlike Schrödinger Functional)

Note that the choice of projector and kinematics is not unique In particular, SMOM scheme

$$p_1 \neq p_2$$
 and  $p_1^2 = p_2^2 = (p_1 - p_2)^2$ 

Can use *q* as projector

Remarks

- Can be generalised to the four-quark operator mixing case
- Non-perturbative off-shell and massless scheme(s)
- Requires gauge fixing (unlike Schrödinger Functional)

Note that the choice of projector and kinematics is not unique In particular, SMOM scheme

$$p_1 \neq p_2$$
 and  $p_1^2 = p_2^2 = (p_1 - p_2)^2$ 

Can use *q* as projector

In principle the results should agree after conversion to  $\overline{\rm MS},$  and extrapolation to the continuum limit

#### Renormalisation basis of the $\Delta F = 2$ operators

As for BSM neutral meson mixing one needs to renormalise 5 operators ,

(27,1) 
$$O_1^{\Delta S=2} = \gamma_\mu \times \gamma_\mu + \gamma_\mu \gamma_5 \times \gamma_\mu \gamma_5$$

$$\begin{array}{lll} (8,8) & \left\{ \begin{array}{l} O_2^{\Delta s=2} &=& \gamma_\mu \times \gamma_\mu - \gamma_\mu \gamma_5 \times \gamma_\mu \gamma_5 \\ O_3^{\Delta s=2} &=& 1 \times 1 - \gamma_5 \times \gamma_5 \end{array} \right. \\ (6,\overline{6}) & \left\{ \begin{array}{l} O_4^{\Delta s=2} &=& 1 \times 1 + \gamma_5 \times \gamma_5 \\ O_5^{\Delta s=2} &=& \sigma_{\mu\nu} \times \sigma_{\mu\nu} \end{array} \right. \end{array}$$

So the renormalisation matrix has the form

$$\mathcal{Z}^{\Delta S=2}=\left(egin{array}{cccc} \mathcal{Z}_{11} & & & \ & \mathcal{Z}_{22} & \mathcal{Z}_{23} & & \ & \mathcal{Z}_{32} & \mathcal{Z}_{33} & & \ & & & \mathcal{Z}_{44} & \mathcal{Z}_{45} & \ & & & & \mathcal{Z}_{54} & \mathcal{Z}_{55} \end{array}
ight)$$

## More details on NPR

- Setup is the similar to RBC-UKQCD
   In particular we follow [Arthur & Boyle '10]
- We implement momentum sources [Gockeler et al '98] to achieve high stat. accuracy
- Non exceptional kinematic with symmetric point  $p_1^2 = p_2^2 = (p_2 p_1)^2$



to suppress IR contaminations [Sturm et al', RBC-UKQCD '09 '10]

## Choice of SMOM scheme

Orientation of the momenta kept fixed

$$p_1 = \frac{2\pi}{L}[n, 0, n, 0]$$
  $p_2 = \frac{2\pi}{L}[0, n, n, 0]$ 

 $\Rightarrow$  Well defined continuum limit

## Choice of SMOM scheme

Orientation of the momenta kept fixed

$$p_1 = \frac{2\pi}{L}[n, 0, n, 0]$$
  $p_2 = \frac{2\pi}{L}[0, n, n, 0]$ 

 $\Rightarrow$  Well defined continuum limit

• We chose  $\gamma_{\mu}$  projectors, for example

$$P^{(\gamma_{\mu})} \quad \leftrightarrow \quad \gamma_{\mu} imes \gamma_{\mu} + \gamma_{\mu} \gamma_{5} imes \gamma_{\mu} \gamma_{5}$$

 $\Rightarrow$  Z factor of a four quark operator O in the scheme  $(\gamma_{\mu}, \gamma_{\mu})$  defined by

$$\lim_{m \to 0} \left. \frac{Z_O^{(\gamma_\mu, \gamma_\mu)}}{Z_V^2} \frac{P^{(\gamma_\mu)} \left\{ \Lambda_O \right\}}{\left( P^{(\gamma_\mu)} \left\{ \Lambda_V \right\} \right)^2} \right|_{\mu^2 = p^2} = Tree$$

Note that this defines an off-shell massless scheme

## Step-scaling

Rome-Southampton method requires a *windows* 

 $\Lambda^2_{QCD} \ll \mu^2 \ll (\pi/a)^2$ 

• And our lattice spacings are  $a^{-1} \sim 2.2, 1.7, 1.3 GeV$ 

## Step-scaling

Rome-Southampton method requires a windows

 $\Lambda^2_{QCD} \ll \mu^2 \ll (\pi/a)^2$ 

And our lattice spacings are  $a^{-1} \sim 2.2, 1.7, 1.3 GeV$ 

we follow [Arthur & Boyle '10] and [Arthur, Boyle, NG, Kelly, Lytle '11] and define

$$\sigma(\mu_2,\mu_1) = \lim_{a^2 \to 0} \lim_{m \to 0} \left[ (P\Lambda(\mu_2,a))^{-1} P\Lambda(\mu_1,a) \right] = \lim_{a^2 \to 0} Z(\mu_2,a) Z(\mu_1,a)^{-1}$$

## Step-scaling

Rome-Southampton method requires a windows

 $\Lambda^2_{QCD} \ll \mu^2 \ll (\pi/a)^2$ 

And our lattice spacings are  $a^{-1} \sim 2.2, 1.7, 1.3 GeV$ 

we follow [Arthur & Boyle '10] and [Arthur, Boyle, NG, Kelly, Lytle '11] and define

$$\sigma(\mu_2,\mu_1) = \lim_{a^2 \to 0} \lim_{m \to 0} \left[ (P\Lambda(\mu_2,a))^{-1} P\Lambda(\mu_1,a) \right] = \lim_{a^2 \to 0} Z(\mu_2,a) Z(\mu_1,a)^{-1}$$

• We use 3 lattice spacings to compute  $\sigma(2 \text{ GeV}, 1.5 \text{ GeV})$  but only the two finest to compute  $\sigma(3 \text{ GeV}, 2 \text{ GeV})$  and get

 $Z(3 \text{ GeV}, a) = \sigma(3 \text{ GeV}, 2 \text{ GeV}) \sigma(2 \text{ GeV}, 1.5 \text{ GeV}) Z(1.5 \text{ GeV}, a)$ 

#### Intermezzo: the importance of SMOM schemes

based on RBC-UKQCD 2010-now

... [NG Hudspith Lytle'16], [Boyle NG Hudspith Lehner Lytle '17] [...Kettle, Khamseh, Tsang 17-18]

#### BSM kaon mixing - Results



- $\blacksquare$  The Green functions might suffer from IR poles,  $\sim 1/p^2,$  or  $\sim 1/m_\pi^2$  which can pollute the signal
- In principle these poles are suppressed at high  $\mu$  but they appear to be quite important at  $\mu\sim$  3 GeV for some quantities which allow for pion exchanges
- The traditional way is to "subtract " these contamination by hand

- $\blacksquare$  The Green functions might suffer from IR poles,  $\sim 1/p^2,$  or  $\sim 1/m_\pi^2$  which can pollute the signal
- In principle these poles are suppressed at high  $\mu$  but they appear to be quite important at  $\mu\sim$  3 GeV for some quantities which allow for pion exchanges
- The traditional way is to "subtract " these contamination by hand
- However these contaminations are highly suppressed in a SMOM scheme, with non-exceptional kinematics
- We argue that this pion pole subtractions is not-well under control and that schemes with exceptional kinematics should be discarded





#### BSM kaon mixing - Results



#### BSM kaon mixing - Results


#### Better MOM schemes ?

# More MOM schemes

Renormalisation scale is  $\mu$ , given by the choice of kinematics

Original RI-MOM scheme

$$p_1=p_2$$
 and  $\mu^2\equiv p_1^2=p_2^2$ 

But this lead to "exceptional kinematics' and bad IR poles

# More MOM schemes

Renormalisation scale is  $\mu$ , given by the choice of kinematics

Original RI-MOM scheme

$$p_1 = p_2$$
 and  $\mu^2 \equiv p_1^2 = p_2^2$ 

But this lead to "exceptional kinematics' and bad IR poles

then RI-SMOM scheme

$$p_1 \neq p_2$$
 and  $\mu^2 \equiv p_1^2 = p_2^2 = (p_1 - p_2)^2$ 

Much better IR behaviour [Sturm et al., Lehner and Sturm, Gorbahn and Jäger, Gracey, ...]

## More MOM schemes

Renormalisation scale is  $\mu$ , given by the choice of kinematics

Original RI-MOM scheme

 $\textit{p}_1=\textit{p}_2$  and  $\mu^2\equiv\textit{p}_1^2=\textit{p}_2^2$ 

But this lead to "exceptional kinematics' and bad  $\operatorname{IR}$  poles

then RI-SMOM scheme

$$p_1 \neq p_2$$
 and  $\mu^2 \equiv p_1^2 = p_2^2 = (p_1 - p_2)^2$ 

Much better IR behaviour [Sturm et al., Lehner and Sturm, Gorbahn and Jäger, Gracey, ...] We are now studying a generalisation (see also [Bell and Gracey])

$$p_1
eq p_2$$
 and  $\mu^2\equiv p_1^2=p_2^2,~~(p_1-p_2)^2=\omega\mu^2$  where  $\omega\in[0,4]$ 

Note that  $\omega = 0 \leftrightarrow RI - MOM$  and  $\omega = 1 \leftrightarrow RI - SMOM$ 

In collaboration with [...,Cahill, Gorbahn, Gracey, Perlt , Rakow, ... ]

Nicolas Garron (University of Liverpool)

#### Back to $0\nu\beta\beta$ : Physical results

#### Chiral extrapolations

With

$$\Lambda_{\chi} = 4\pi F_{\pi} , \qquad \qquad \epsilon_{\pi} = \frac{m_{\pi}}{\Lambda_{\chi}} ,$$

we find in the continuum at NLO ( $\beta_i$  and  $c_i$  are free parameters)

$$\begin{split} O_1 &= \frac{\beta_1 \Lambda_{\chi}^4}{(4\pi)^2} \bigg[ 1 + \frac{7}{3} \epsilon_{\pi}^2 \ln(\epsilon_{\pi}^2) + c_1 \epsilon_{\pi}^2 \bigg] \\ O_2 &= \frac{\beta_2 \Lambda_{\chi}^4}{(4\pi)^2} \bigg[ 1 + \frac{7}{3} \epsilon_{\pi}^2 \ln(\epsilon_{\pi}^2) + c_2 \epsilon_{\pi}^2 \bigg] \\ \frac{O_3}{\epsilon_{\pi}^2} &= \frac{\beta_3 \Lambda_{\chi}^4}{(4\pi)^2} \bigg[ 1 + \frac{4}{3} \epsilon_{\pi}^2 \ln(\epsilon_{\pi}^2) + c_3 \epsilon_{\pi}^2 \bigg] \end{split}$$

### Chiral extrapolations

With

$$\Lambda_{\chi} = 4\pi F_{\pi} , \qquad \qquad \epsilon_{\pi} = \frac{m_{\pi}}{\Lambda_{\chi}} ,$$

we find in the continuum at NLO ( $\beta_i$  and  $c_i$  are free parameters)

$$\begin{split} O_1 &= \frac{\beta_1 \Lambda_{\chi}^4}{(4\pi)^2} \bigg[ 1 + \frac{7}{3} \epsilon_{\pi}^2 \ln(\epsilon_{\pi}^2) + c_1 \epsilon_{\pi}^2 \bigg] \\ O_2 &= \frac{\beta_2 \Lambda_{\chi}^4}{(4\pi)^2} \bigg[ 1 + \frac{7}{3} \epsilon_{\pi}^2 \ln(\epsilon_{\pi}^2) + c_2 \epsilon_{\pi}^2 \bigg] \\ \frac{O_3}{\epsilon_{\pi}^2} &= \frac{\beta_3 \Lambda_{\chi}^4}{(4\pi)^2} \bigg[ 1 + \frac{4}{3} \epsilon_{\pi}^2 \ln(\epsilon_{\pi}^2) + c_3 \epsilon_{\pi}^2 \bigg] \end{split}$$

In practice, these expressions are modified to incorporate  $a^2$ , mixed-action effects and finite volume effects

#### Extrapolations to the physical point



## Extrapolations to the physical point



#### Extrapolations to the physical point



### "Pion bag parameter"

We define  $B_{\pi} = O_3 / (\frac{8}{3} m_{\pi}^2 F_{\pi}^2)$ 



## Physical results



 $1-2\sigma$  agreement with [V. Cirigliano, W. Dekens, M. Graesser, E. Mereghetti 1701.01443] where they use an estimate from SU(3)  $\chi$ PT

but the uncertainty decreases from 20-40% to 5-8%

## Physical results



 $1-2\sigma$  agreement with [V. Cirigliano, W. Dekens, M. Graesser, E. Mereghetti 1701.01443] where they use an estimate from  $SU(3)~\chi{\rm PT}$ 

but the uncertainty decreases from 20-40% to 5-8%

```
and B_{\pi} = 0.430(16)[0.432(16)]
```

low value, far from 1 as anticipated for example by [Pich & de Rafael]

#### Conclusions

## Conclusions and outlook

- First computation of heavy contributions to  $0\nu\beta\beta$ , the  $\langle \pi^+|O|\pi^-\rangle$  MEs
- Accepted by PRL [A. Nicholson, E. Berkowitz, H. Monge-Camacho, D. Brantley, N.G., C.C. Chang, E. Rinaldi, M.A. Clark, B. Joo, T. Kurth, B. Tiburzi, P. Vranas, A. Walker-Loud] arXiv:1805.02634

# Conclusions and outlook

- First computation of heavy contributions to  $0\nu\beta\beta$ , the  $\langle \pi^+|O|\pi^-\rangle$  MEs
- Accepted by PRL [A. Nicholson, E. Berkowitz, H. Monge-Camacho, D. Brantley, N.G., C.C. Chang, E. Rinaldi, M.A. Clark, B. Joo, T. Kurth, B. Tiburzi, P. Vranas, A. Walker-Loud] arXiv:1805.02634

Our computation features

- Good Chiral symmetry
- Non-perturbative renormalisation
- Physical pion masses, three lattice spacings

As for BSM neutral meson meson mixing, chiral symmetry and SMOM schemes are crucial !

## Conclusions and outlook

There is still some work to do:

- Compute contributions within nuclei  $\langle N | O | N \rangle$
- Other unknown short-distance contributions
- Long-distance contributions ?

### Backup



ØА

Insertion of the axial current between two nucleon state,

$$\langle N(p')|\overline{\psi}\gamma_{\mu}\gamma_{5}\psi|N(p)
angle = \overline{u}(p')\left[\gamma_{\mu}\gamma_{5}G_{A}(q^{2}) + \gamma_{5}rac{q_{\mu}}{2m_{N}}G_{P}(q^{2})
ight]\overline{u}(p)$$

where q is the momentum transfer q = p' - p

The nucleon axial coupling is then

$$g_A = G_A(0)$$

 $g_A$  is the strength at which the nucleon couples to the axial current

#### $g_A$ the Nucleon axial coupling

Nuclear  $\beta$  decay:  $n \longrightarrow p + e^- + \bar{\nu}_e$ 



 $\rightarrow$  Well-measured experimentally  $g_A = 1.2723(23)$  error < 0.2%

# A problem on the lattice

- It should be a relatively "simple" quantity
- But turned out to be a long standing puzzle
- Can we believe in lattice results for nucleons ?
- Or is there a problem with QCD ?

#### A problem on the lattice



Summary plot from [Martha Constantinou @ Lat2014]

# Our computation

With CalLat (California Lattice) Collaboration

- Möbius fermions on  $N_f = 2 + 1 + 1$  HISQ ensembles  $\Rightarrow$  Chiral symmetry
- 3 lattice spacings  $a \sim 0.15, 0.012, 0.09$  fm, several volumes
- Multiple pion mass and physical pion mass on a ~ 0.15, 0.012 ensembles

 $\Rightarrow$  Good control over Chiral/cont./ infinite Vol. extrapolations

# Our computation

Main improvements (compared to recent computations)

- New method to extract the signal "kills" the noise problem
- Chiral fermions, so dominant Lattice artefacts are a<sup>2</sup> and a<sup>4</sup>
- Non-perturbative renormalisation  $Z_A/Z_V = 1$

$$g_A = rac{Z_A}{Z_V} \left(rac{g_A}{g_V}
ight)^{\mathrm{bare}}$$









 $g_A^{QCD} = 1.271(13)$   $g_A^{PDG} = 1.2723(23)$ [Chang, Nicholson, Rinaldi, Berkowitz, N.G., Brantley, Monge-Camacho, Monahan, Bouchard, Clark, Joó, Kurth, Orginos, Vranas, Walker-Loud]

Published in Nature 558 (2018) no.7708

Nicolas Garron (University of Liverpool)

## Error budget

#### $g_A = 1.2711(103)^{\mathrm{s}}(39)^{\chi}(15)^{\mathrm{a}}(19)^{\mathrm{v}}(04)^{\mathrm{I}}(55)^{\mathrm{M}}$

where the errors are statistical (s), chiral ( $\chi$ ), continuum (a), infinite volume (v), isospin breaking (I) and model-selection (M)

To be compare to the experimental value  $g_A = 1.2723(23)$